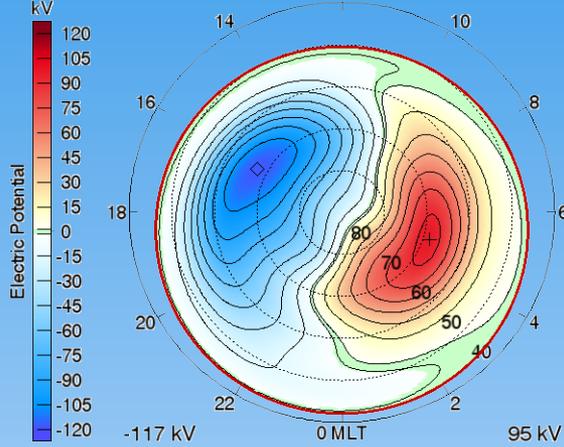


Process 1: Quantifying the storm energy input

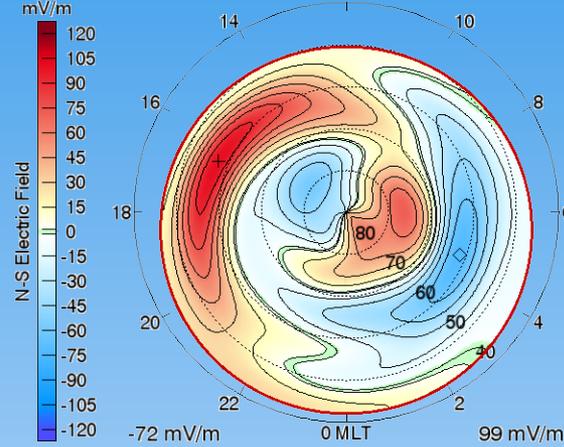
- Increase in magnetospheric/ionospheric high latitude convection and auroral precipitation
- Enhances conductivity at high latitudes and NO production
- Joule heating increase
- NO cooling IR radiation measured by SABER (\propto NO and T)
- Rate of temperature/density response and recovery

Daniel R. Weimer

8 Nov. 2004
06:00 UT

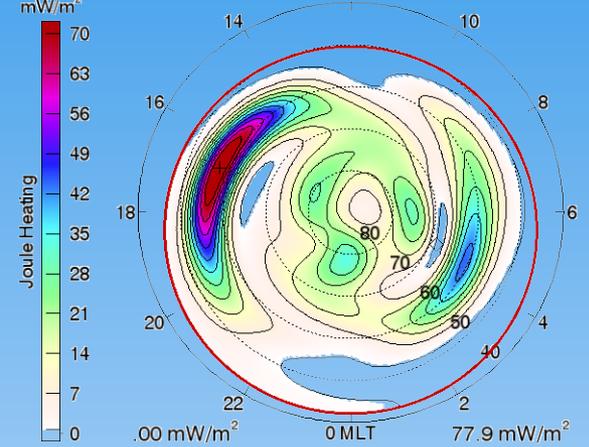


8 Nov. 2004
06:00 UT



$$\mathbf{E} = -\nabla_S \Phi$$

8 Nov. 2004
06:00 UT



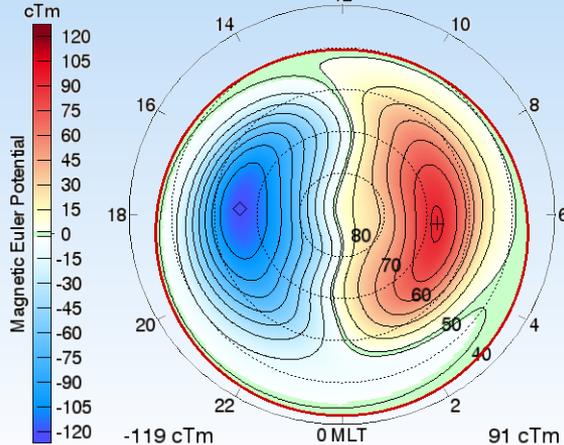
$$\mathbf{S} = \mathbf{E} \times \Delta \mathbf{B} / \mu_o$$

Originally developed for a FAC model, the magnetic potentials are even more useful in combination with the electric potentials to obtain the Poynting flux

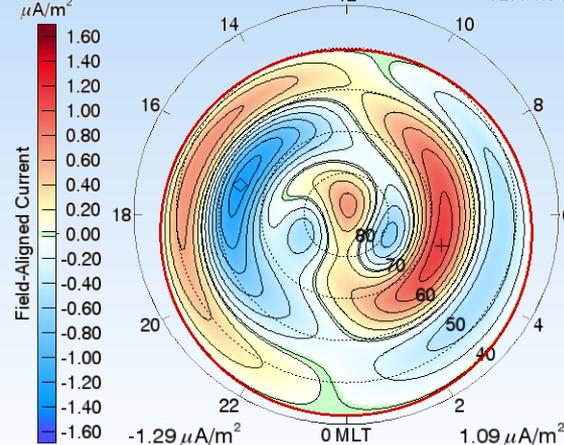
$$J_{\parallel} = \nabla_S^2 \psi / \mu_o$$

$$\Delta \mathbf{B} = \hat{\mathbf{r}} \times \nabla_S \psi$$

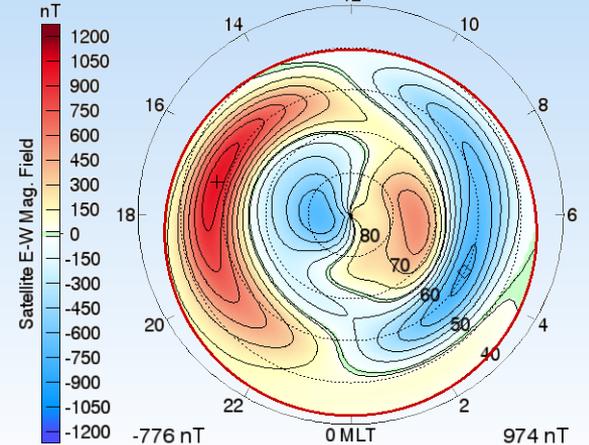
8 Nov. 2004
06:00 UT



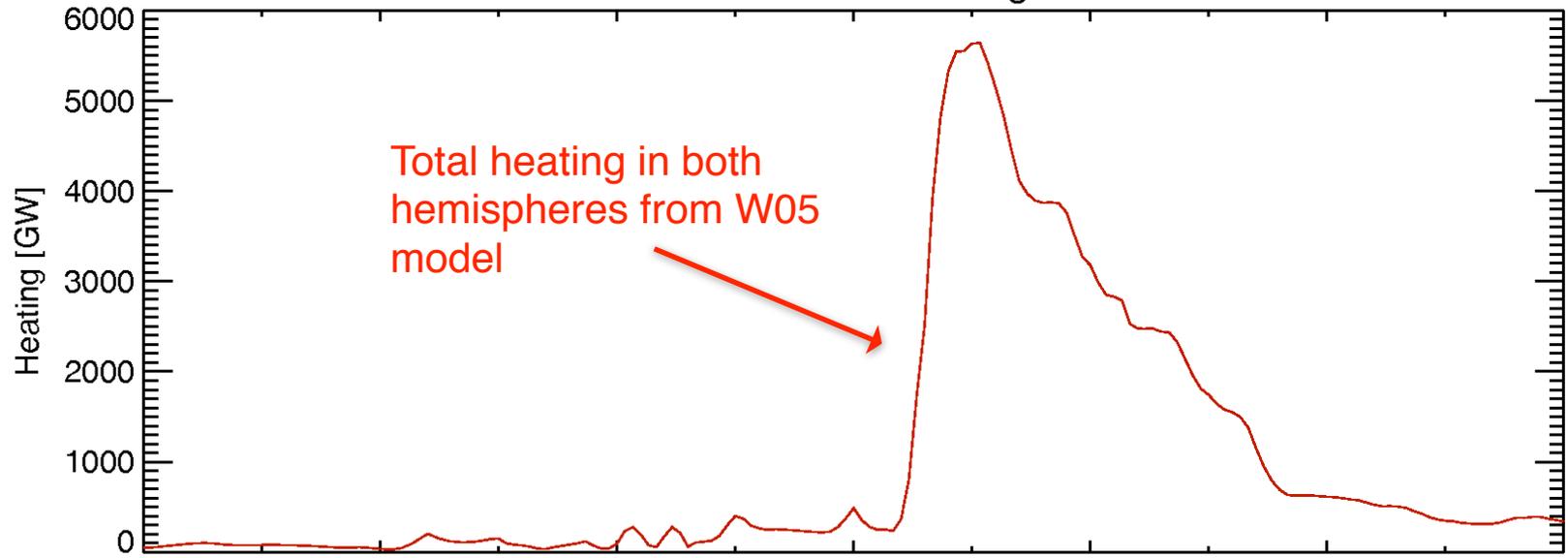
8 Nov. 2004
06:00 UT



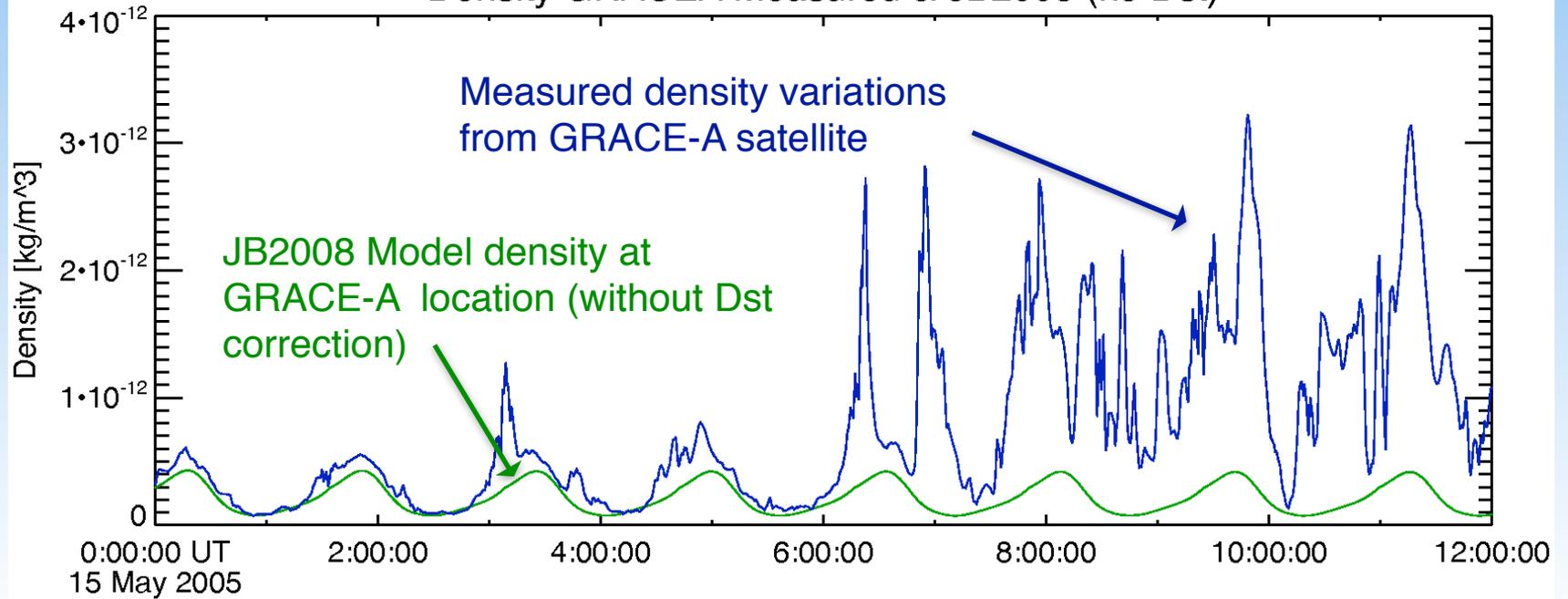
8 Nov. 2004
06:00 UT



W05 Model Heating



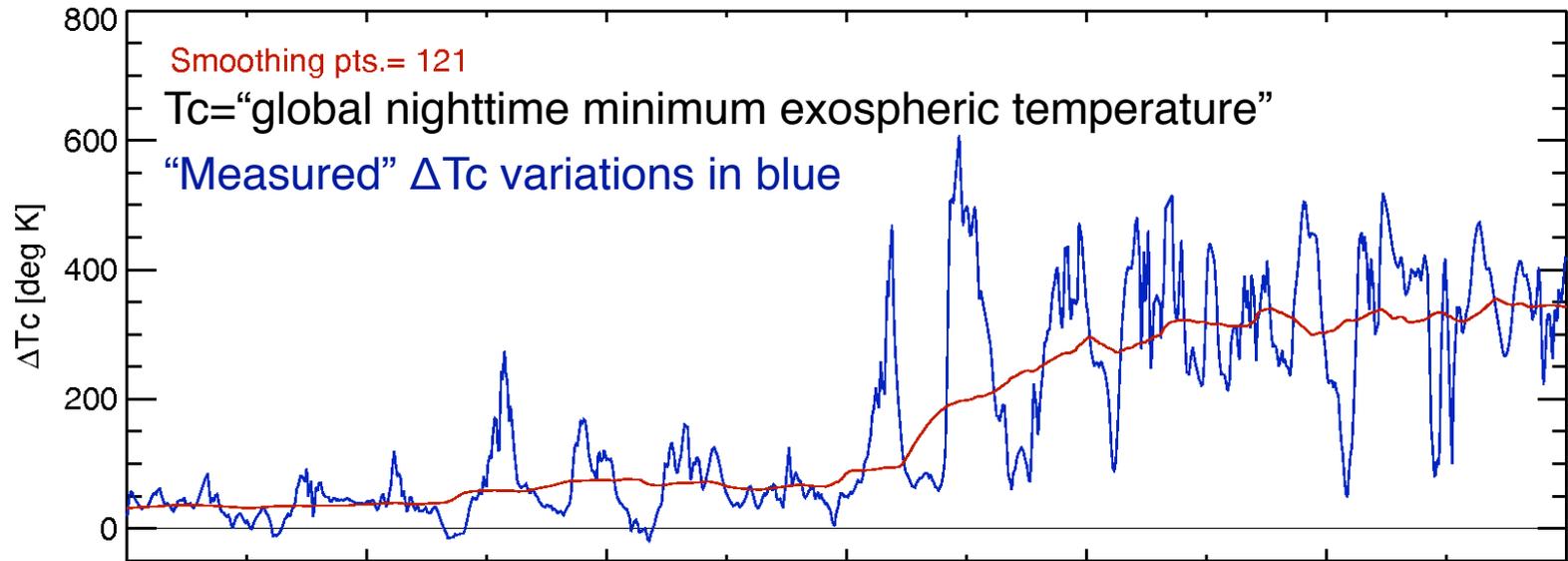
Density GRACEA Measured & JB2008 (no Dst)



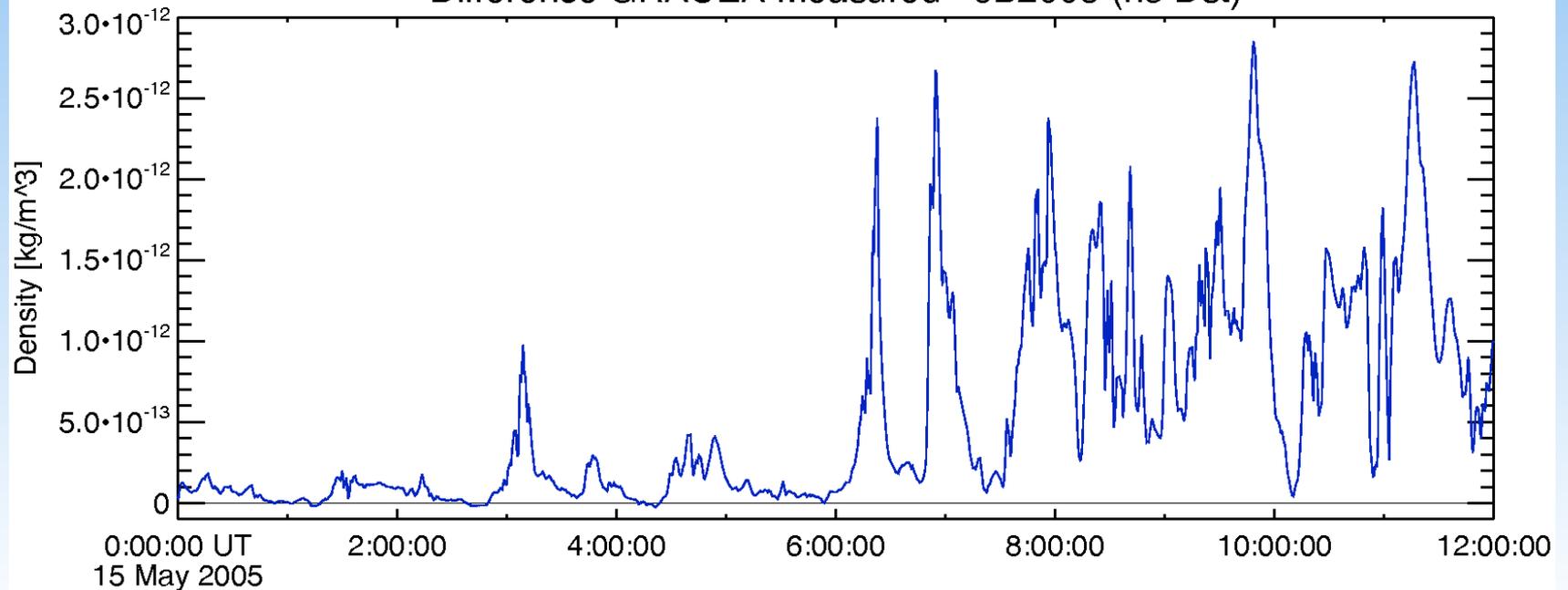
- ❖ The Jacchia-Bowman 2008 (JB2008)* code uses a parameter named T_c , representing the “global nighttime minimum exospheric temperature.” Global values of exospheric temperatures and densities are derived from T_c .
- ❖ The slowly varying, “background” level of T_c is calculated from several indices of solar EUV radiation and X-ray flux.
- ❖ Faster variations, named ΔT_c , are due to the auroral heating in the ionosphere. In the JB2008 model, ΔT_c is derived from the Dst geomagnetic index (but Dst is not real-time data).
- ❖ ΔT_c can be derived by matching densities from JB2008 code with densities measured by CHAMP or GRACE.
- ❖ The total energy content of the thermosphere is proportional to ΔT_c . *Burke [2008]* found that a 1°K increase in ΔT_c raises the total energy in the thermosphere above 100 km altitude by $1.01 \cdot 10^{14}$ Joules.

* Bowman, B. R., W. K. Tobiska, F. A. Marcos, C. Y. Huang, C. S. Lin, and W. J. Burke, A new empirical thermospheric density model JB2008 using new solar and geomagnetic indices, in *AIAA 2008-6438*, AIAA Astrodynamics Conference, Honolulu, HI, 2008.

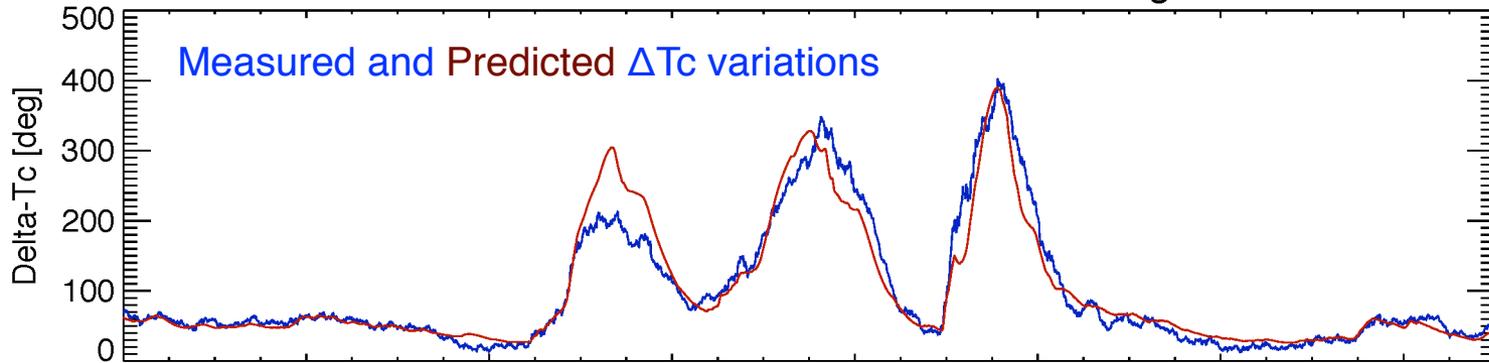
ΔT_c



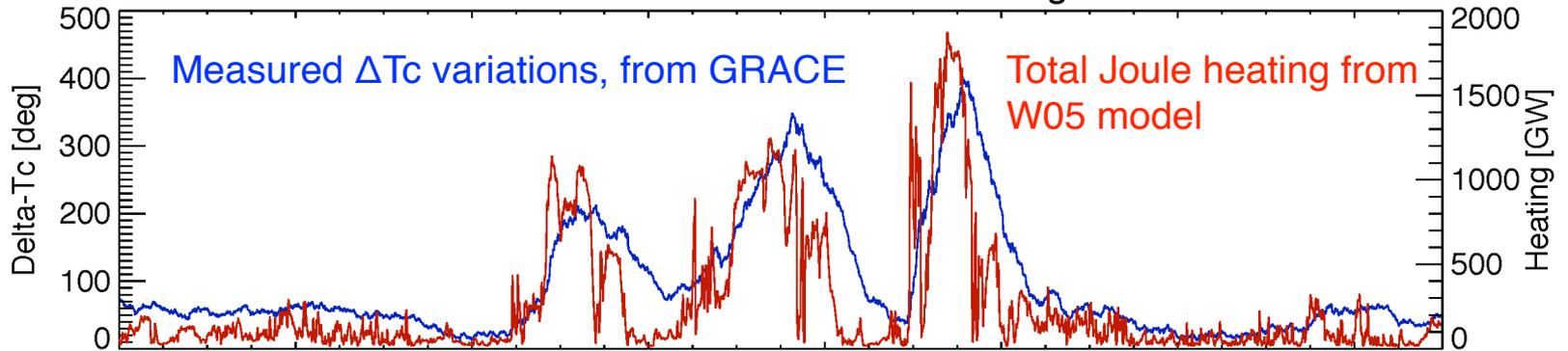
Difference GRACEA Measured - JB2008 (no Dst)



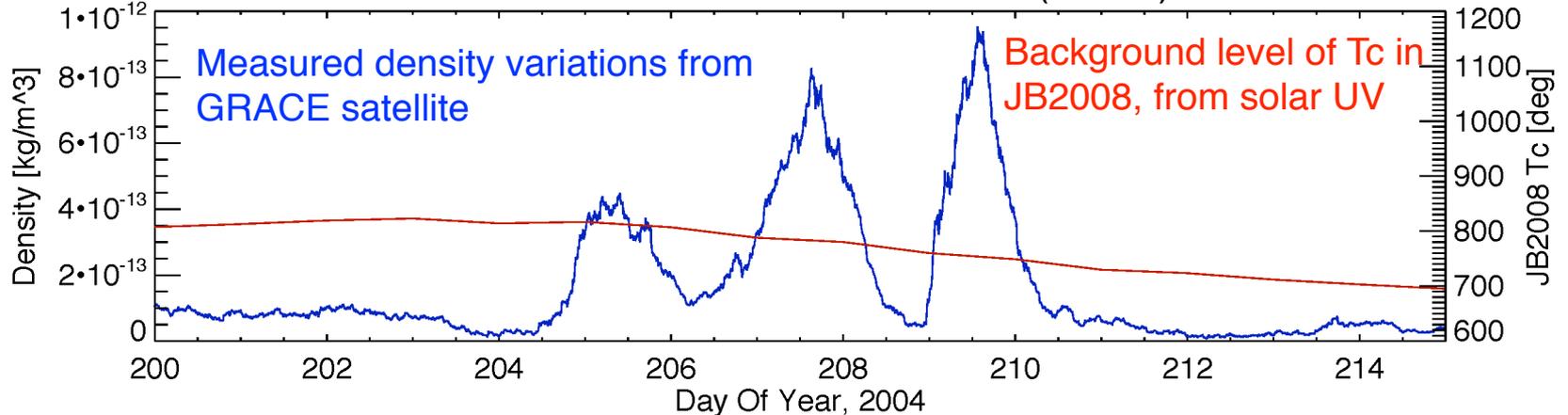
Prediction of Delta-Tc from W05 Heating



Delta-Tc and W05 Model Heating



Difference GRACEA Data - JB2008 (no Dst)



The ΔT_c prediction technique:

$$\Delta T_c(t_{n+1}) = \Delta T_c(t_n) \left(1 - \frac{\Delta t}{\tau_c}\right) + \beta H_J \Delta t$$

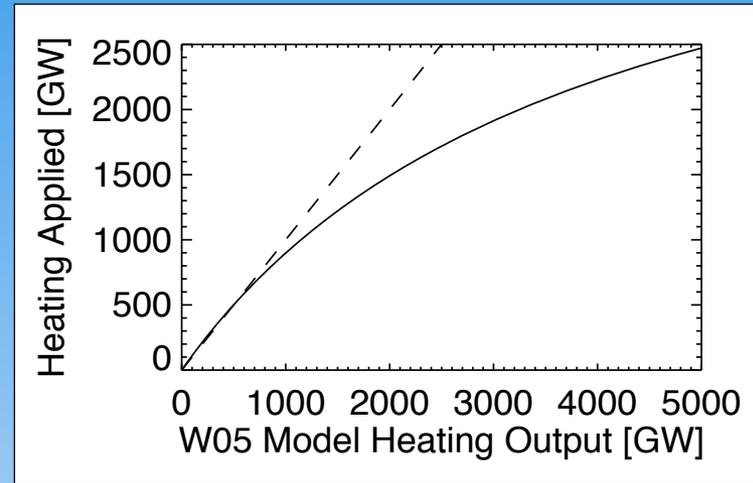
$$\beta = 6.9 \cdot 10^{-4} \text{ (}^\circ\text{K/GW-min)}$$

$$\tau_c = 14.6 \text{ (hours)} - 0.281 \text{ NO}$$

$$\text{NO}(t_{n+1}) = \text{NO}(t_n) \left(1 - \frac{\Delta t}{\tau_{\text{NO}}}\right) + \gamma H_J \Delta t$$

$$\gamma = 2.5 \cdot 10^{-5} \text{ (units/GW-min)}$$

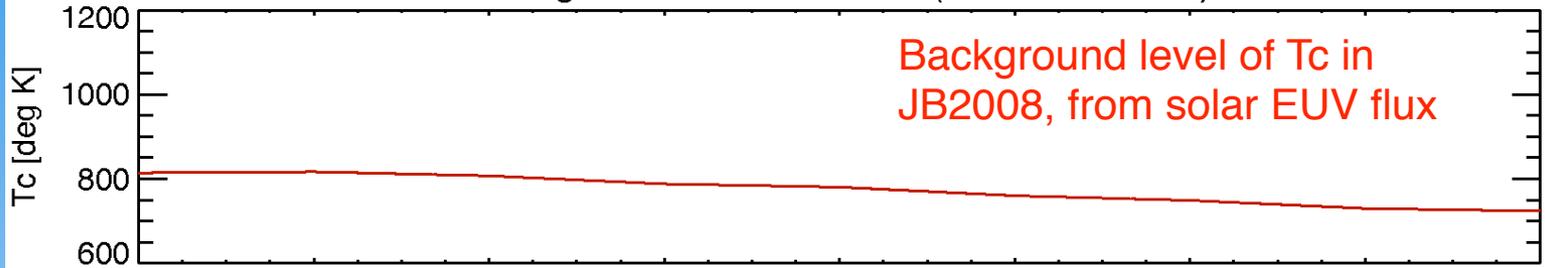
$$\tau_{\text{NO}} = 28.0 \text{ (hours)}$$



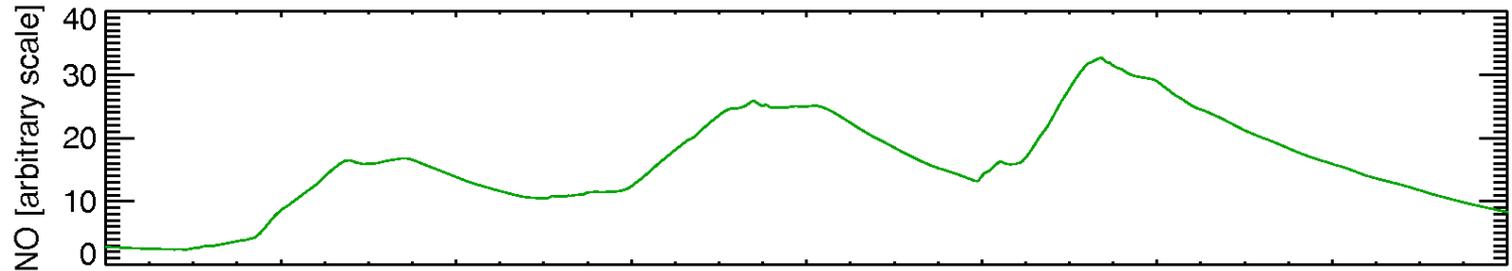
Saturation curve applied to heating

H_J is total Joule heating/Poynting flux from W05 model, with additional saturation applied. All constants (except fixed γ) obtained by fitting five years of H_J with CHAMP and GRACE measurements of ΔT_c . Details are provided in: Weimer, D. R., B. R. Bowman, E. K. Sutton, and W. K. Tobiska (2011), Predicting global average thermospheric temperature changes resulting from auroral heating, *J. Geophys. Res.*, 116, A01312, doi:10.1029/2010JA015685.

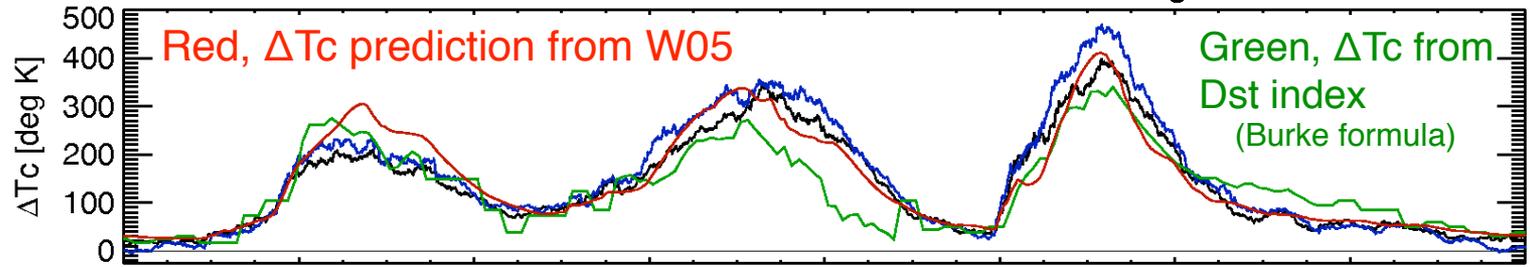
Background Tc in JB2008 (From EUV flux)



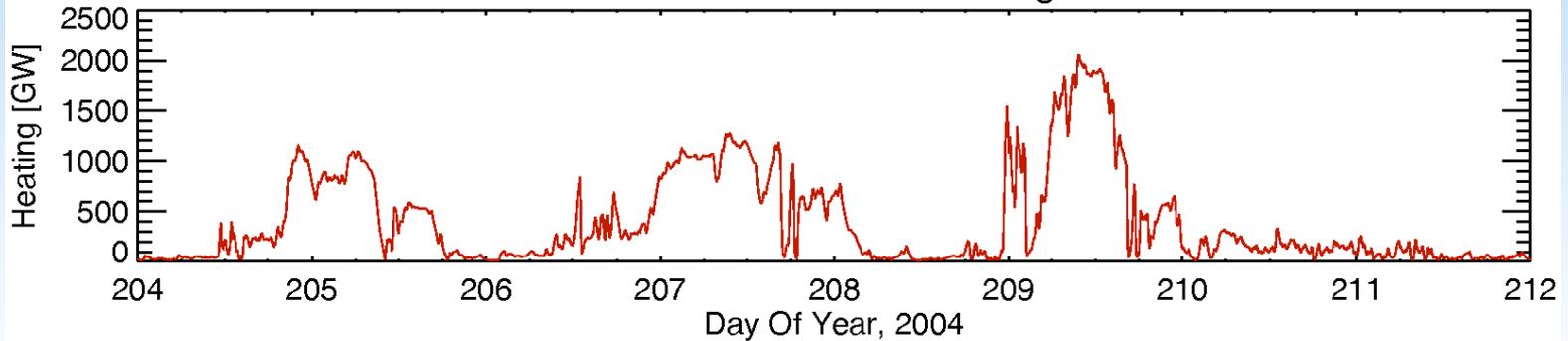
Modeled Nitric Oxide



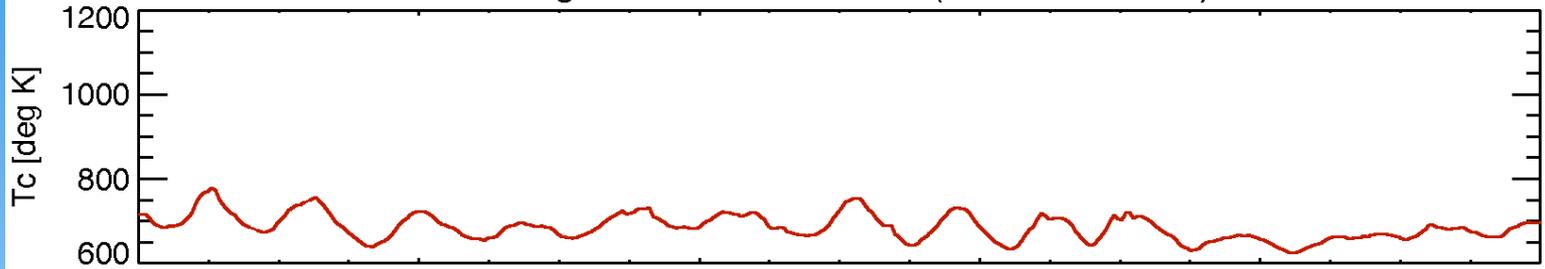
GRACE & CHAMP Measured ΔTc and W05 Heating Prediction



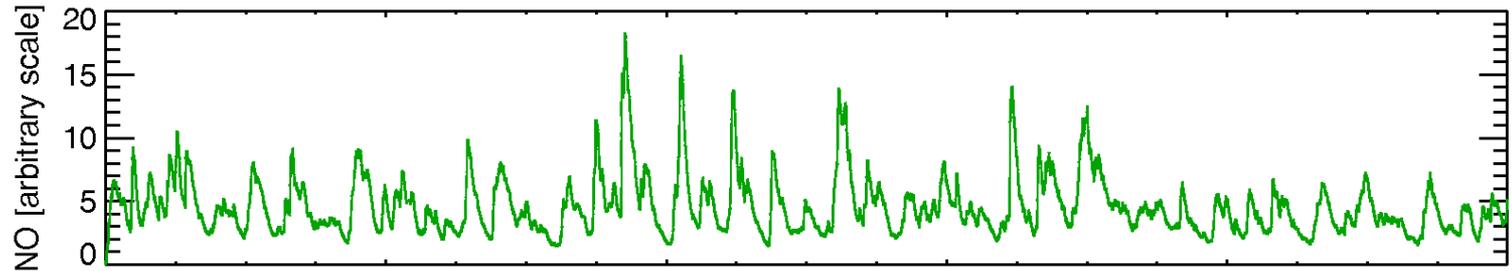
Total W05 Model Heating



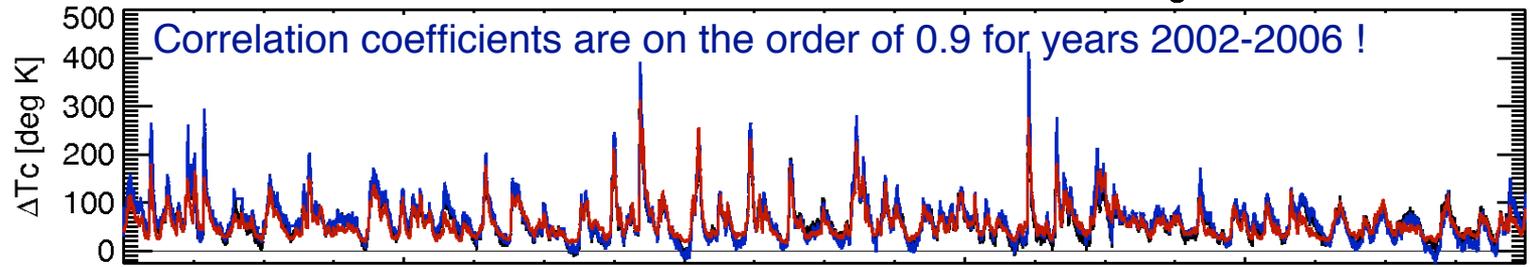
Background Tc in JB2008 (From EUV flux)



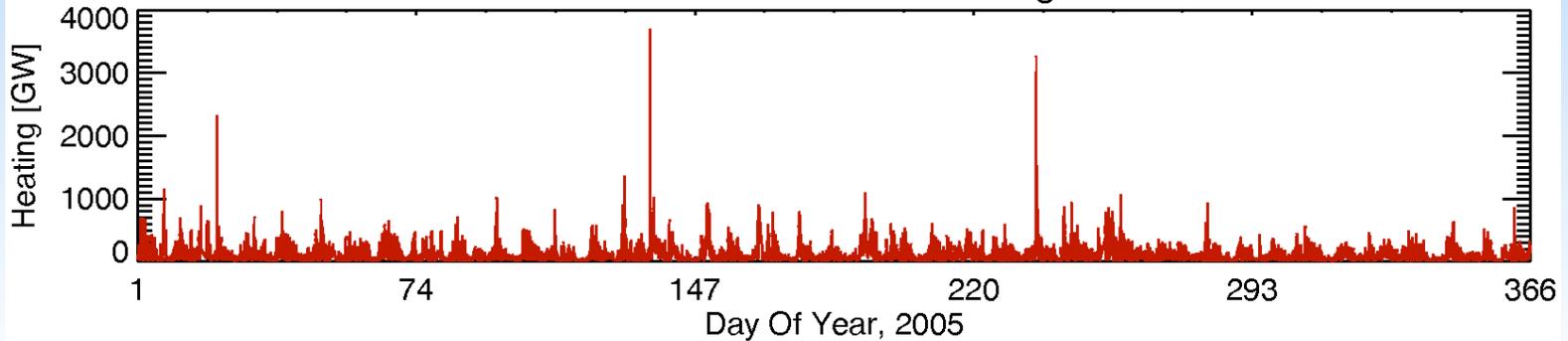
Modeled Nitric Oxide

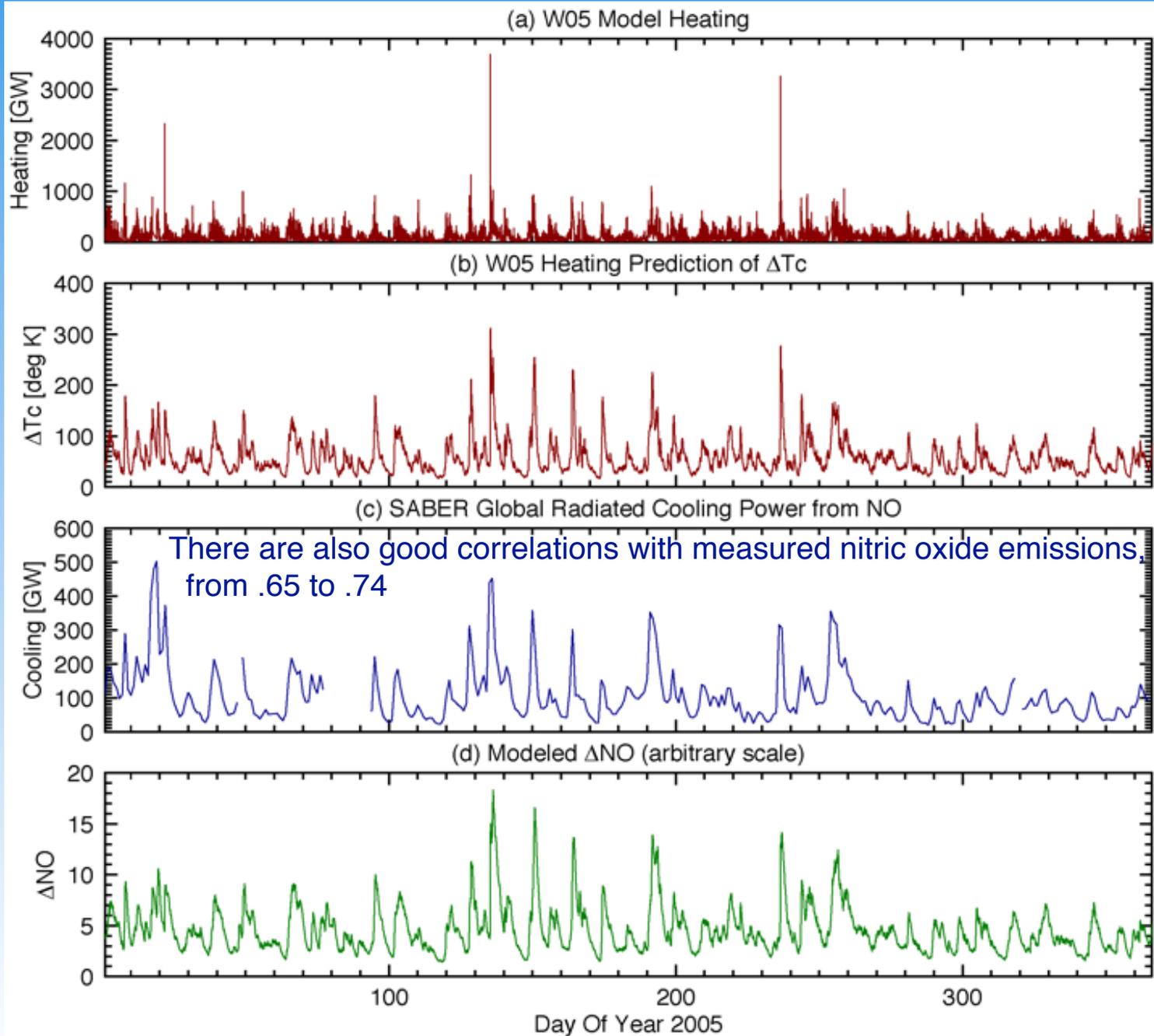


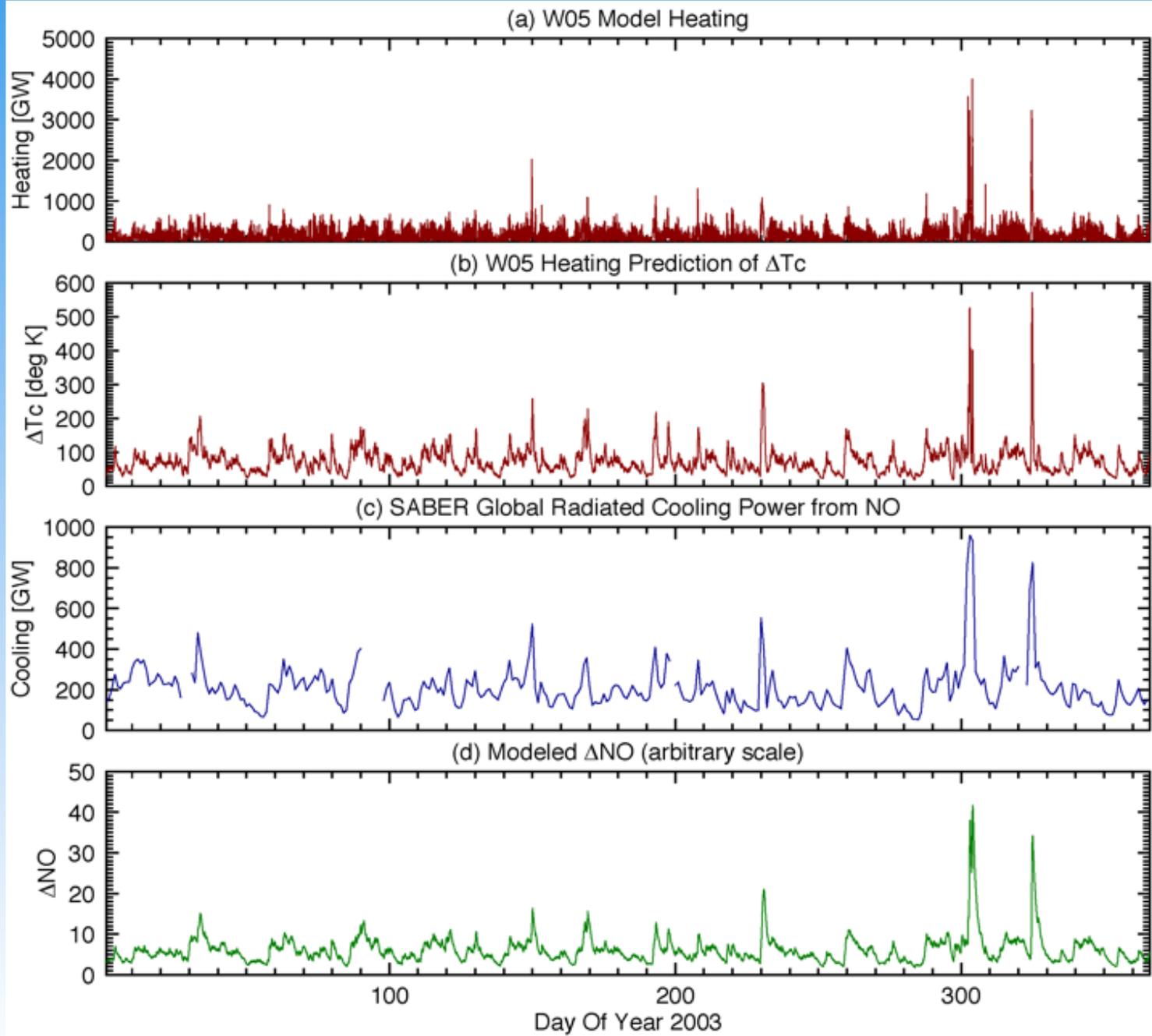
GRACE & CHAMP Measured ΔT_c and W05 Heating Prediction



Total W05 Model Heating

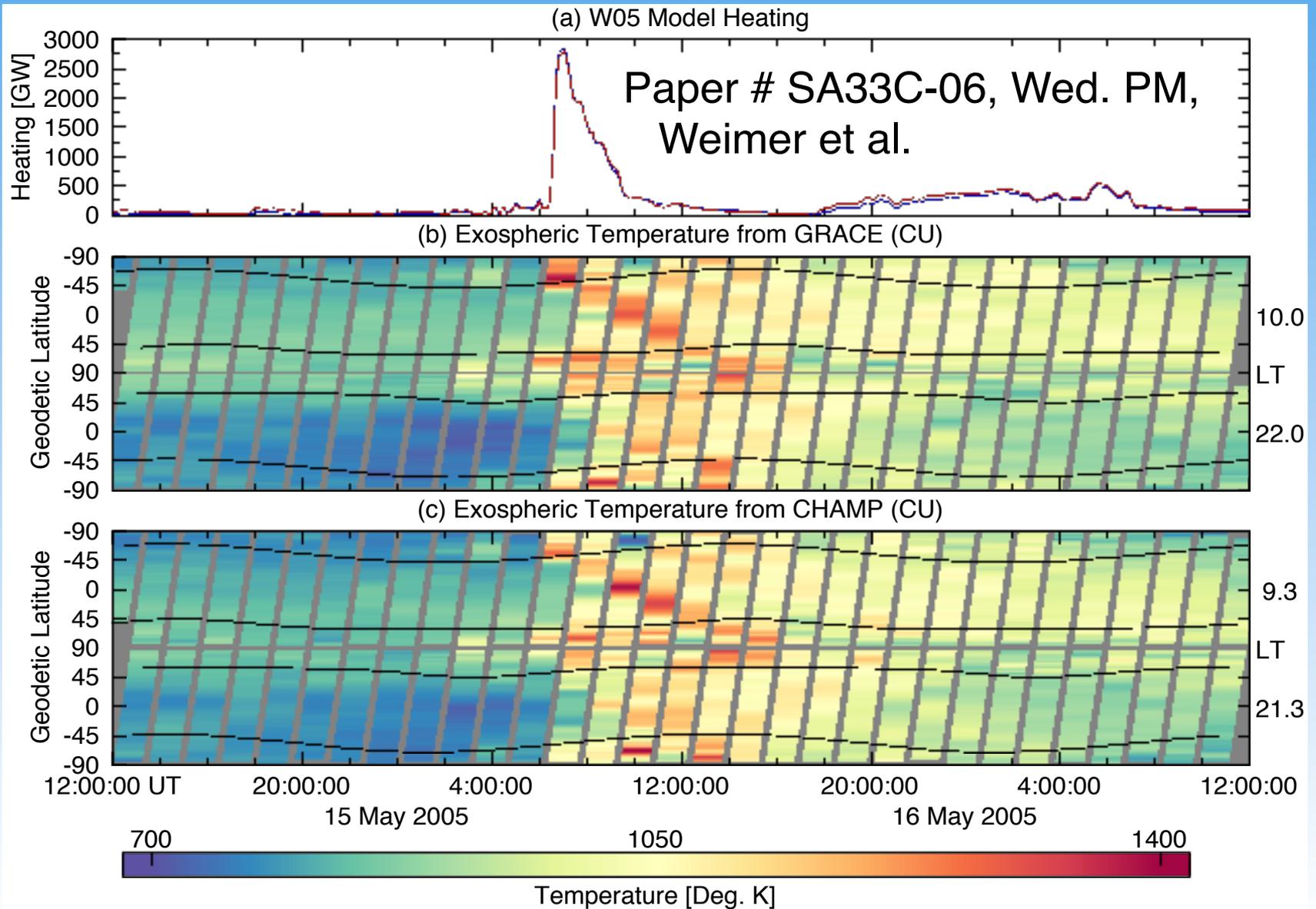






Extreme Heating Event on 15 May 2005

Peak heating at 7 UT, and density spike seen at equator by 9 to 9:45 UT !



Acknowledgments

The author thanks Eric Sutton and Eelco Doornbos for providing calibrated CHAMP and GRACE measurements of neutral density.

Bruce Bowman and Kent Tobiska provided the code for the JB2008 model and solar indices.

TIMED SABER measurements of nitric oxide emissions were provided by Martin Mlynczak, NASA LaRC.

Appendix A: Energy budget of the thermosphere

The total energy content of the thermosphere is proportional to ΔT_c . *Burke* [2008] found that a 1°K increase in ΔT_c raises the total energy in the thermosphere above 100 km altitude by $1.01 \cdot 10^{14}$ Joules.

$$\Delta T_c(t_{n+1}) = \Delta T_c(t_n) \left(1 - \frac{\Delta t}{\tau_c}\right) + \beta H_J \Delta t$$

$$\beta = 6.9 \cdot 10^{-4} \text{ (°K/GW-min)}$$

$$\tau_c = 14.6 \text{ (hours)} - 0.281 NO$$

$$NO(t_{n+1}) = NO(t_n) \left(1 - \frac{\Delta t}{\tau_{NO}}\right) + \gamma H_J \Delta t$$

$$\gamma = 2.5 \cdot 10^{-5} \text{ (units/GW-min)}$$

$$\tau_{NO} = 28.0 \text{ (hours)}$$

With β equal to $6.9 \cdot 10^{-4}$ °K/GW/min, an output of 362 GW from the W05 model over a period of 4 min is needed to raise the T_c temperature by 1°K. A heat input of 362 GW during a 4 min interval amounts to $0.869 \cdot 10^{14}$ J, which is just slightly under the $1.01 \cdot 10^{14}$ Joule figure obtained by *Burke* [2008] for the change in energy. The particle precipitation can account for the other 14%.