



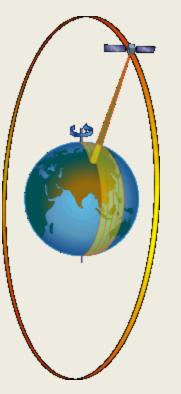
# DMSP Measurements of Poynting Flux into the Ionosphere-Thermosphere System: Examples and Statistics

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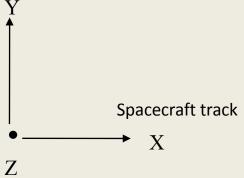
with contributions from:

L. Kilcommons, W. Li, M. Hairston, G. Wilson, J. Lei, K. Drake, C. Lin, A. Richmond



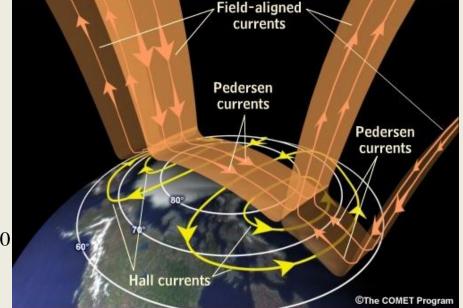
## Poynting Vector from Defense Meteorological Satellite Program S/C

DMSP instruments sense Electric and Magnetic Fields



$$\mathbf{S} = \mathbf{E} \times \delta \mathbf{B}_{DMSPHorizontal} / \mu_0$$

$$S_{||} = (E_x \delta B_y - E_y \delta B_x) / \mu_0$$



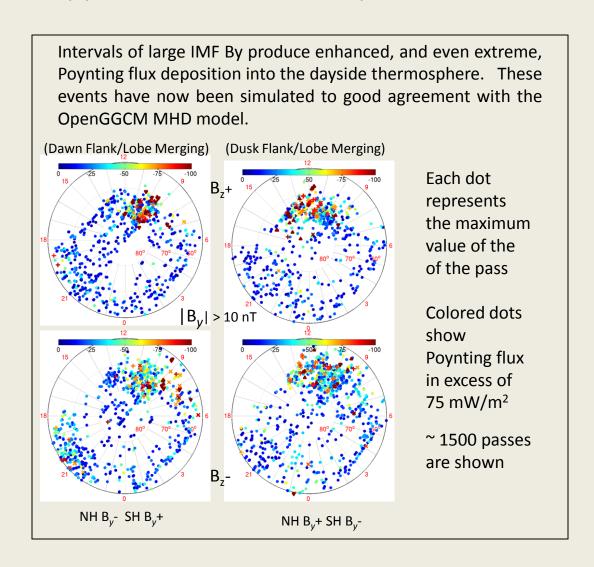
where

$$\mathbf{E} = -\mathbf{V} \times \mathbf{B}_{IGRF}$$
 and  $\delta \mathbf{B}_{DMSP\,Horizontal} = \mathbf{B}_{DMSP} - \mathbf{B}_{Main}$ 

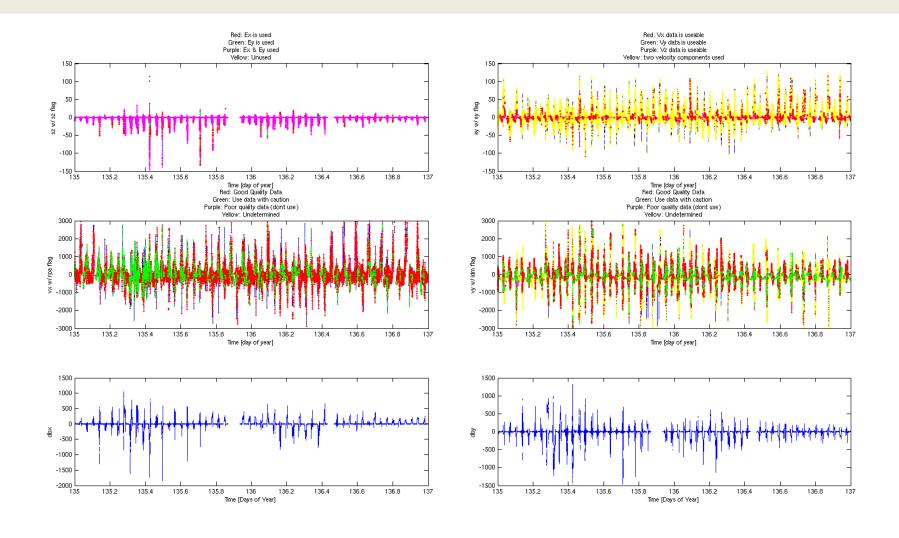
#### Poynting Flux Comparison 2000-2005 IMF By Influence Extreme Poynting Flux in the Dayside Thermosphere: Examples and Statistics [Knipp et al., 2011] GRL in press

- -We combine DMSP F-15 Poynting flux data from 2000-2005, in the northern and southern hemispheres according to the location of the preferred magnetopause merging region.
- -When the IMF By component is large, significant Poynting flux is deposited in the dayside.

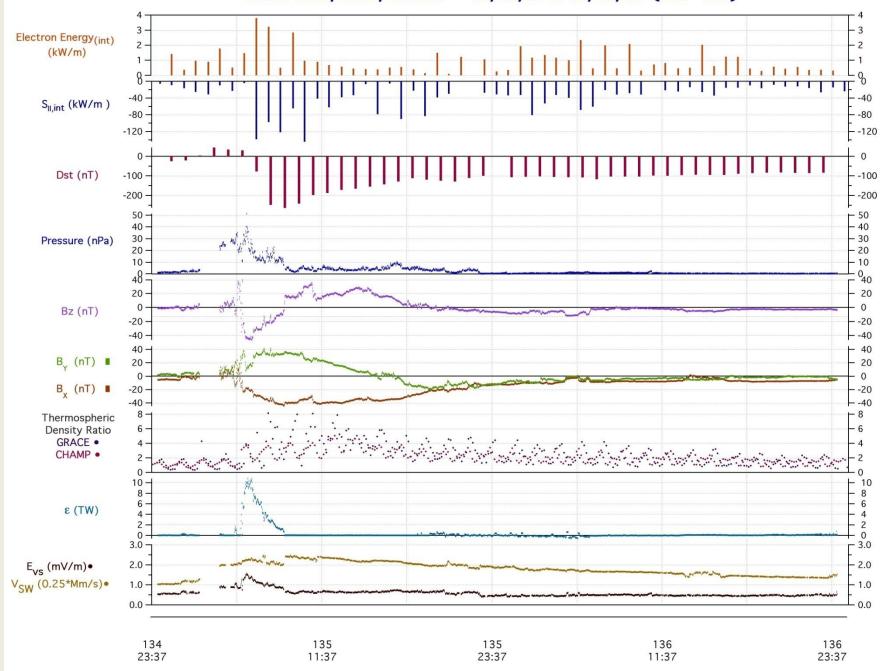
  Deposition may exceed 170 mW/m<sup>2</sup>—an order of magnitude above typical auroral values.
- -Empirical Joule heat models do not capture this result.

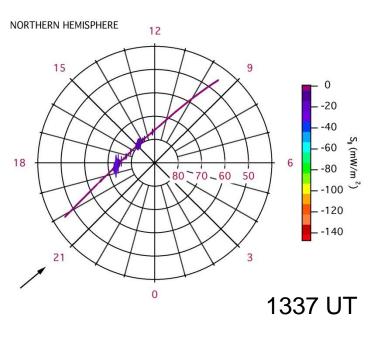


### May 15 and 16 2005



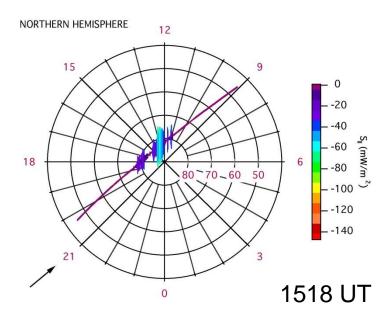
#### DMSP F15 / ACE / GRACE 05/15/05 to 05/16/05 (135 - 136)

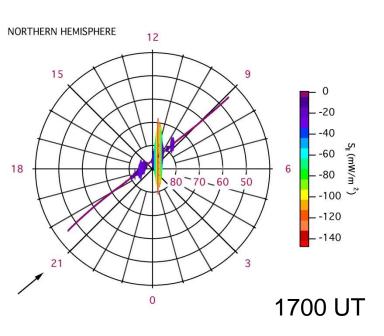


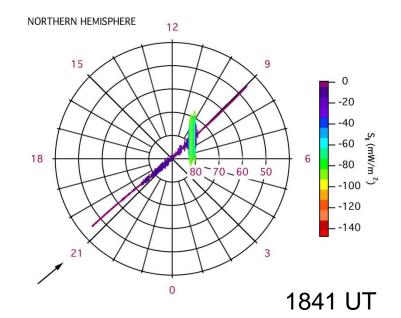


May 15 2005

DMSP F-15 Poynting Flux







# Oct 29 2003 DMSP- SWMF-Weimer Comparison

Compared along DMSP F-15 satellite track

