

The ionospheric quiet-time variation as a challenge for model validation tests: preliminary considerations

I. Tsagouri¹, J. S. Shim², M. M. Kuznetsova³

¹National Observatory of Athens, Greece

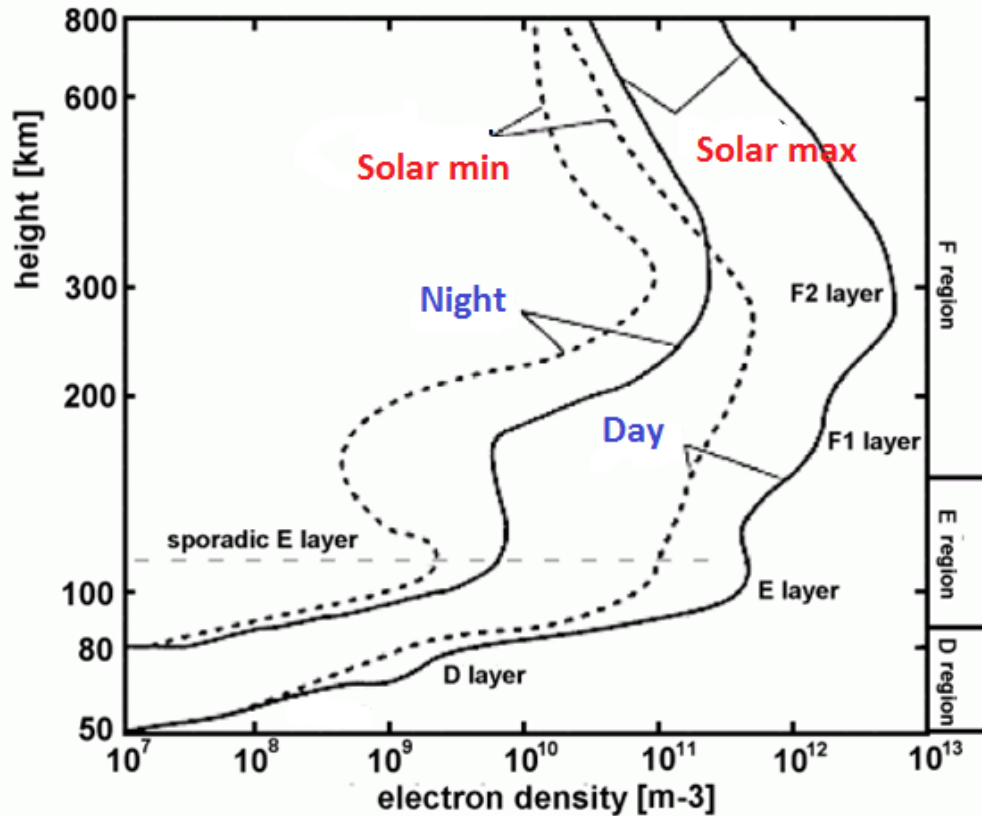
²CUA/ NASA GSFC, Greenbelt, MD, USA

³ NASA GSFC, Greenbelt, MD, USA

International CCMC-LWS Working Meeting:

April 3-7, 2017 Cape Canaveral (FL, USA)

Ionospheric variability



The ionosphere is not the same every day since it is a highly coupled system: ionization production, loss and transport

$$\frac{\partial N_e}{\partial t} = q - I(N_e) - \text{div}(N_e V)$$

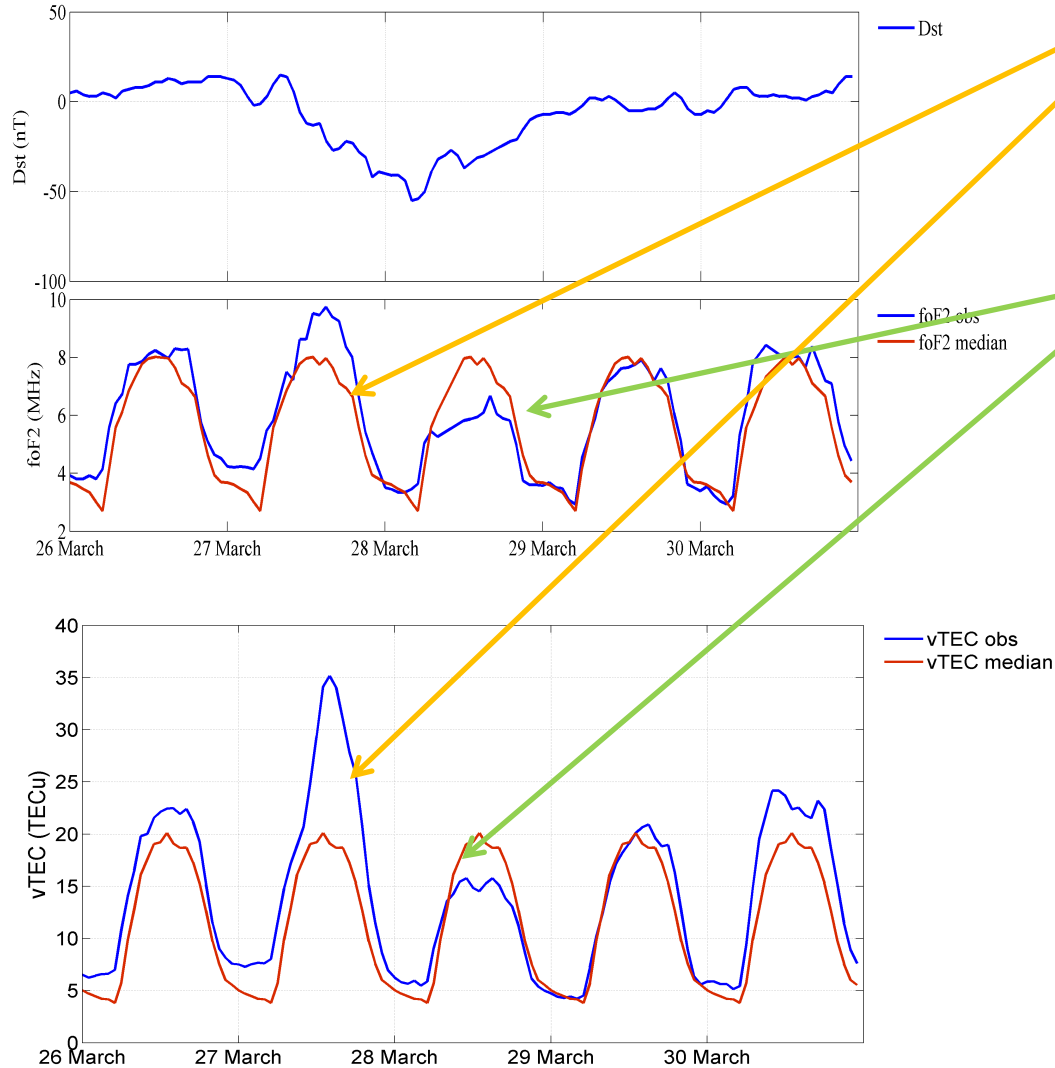
Electron density: changes over multiple timescales ranging from approximately minutes (e.g., solar flare effects) to solar cycle durations (~11 years).

Normal (diurnal, monthly, seasonal, solar cycle, latitudinal, longitudinal dependence)
Transient (e.g. space weather effects) defined with respect to normal changes



Large scale ionospheric storm-time disturbances in plasma density

Dourbes (26 - 30 March 2012)



Positive storm effects:

increase in ionospheric ionization wrt background conditions

Negative storm effects:

decrease in the ionospheric ionization below background conditions

Motivation

Quantification of the quiet time ionospheric variation

- Assessment of climatological models (long-term predictions)
- Quantification of the storm impact – assessment of modeling capabilities for ionospheric short-term forecasting applications

Physical quantities:

NmF2/foF2; hmF2; vTEC

Options

- ❑ Average over 5-quietest days within a month
- ❑ Average over 5-quietest days within 30-days prior to an event

Standard “prediction” approaches

- ❑ Monthly medians
- ❑ Running medians (30-days prior to an event) – Suitable for “real time” applications

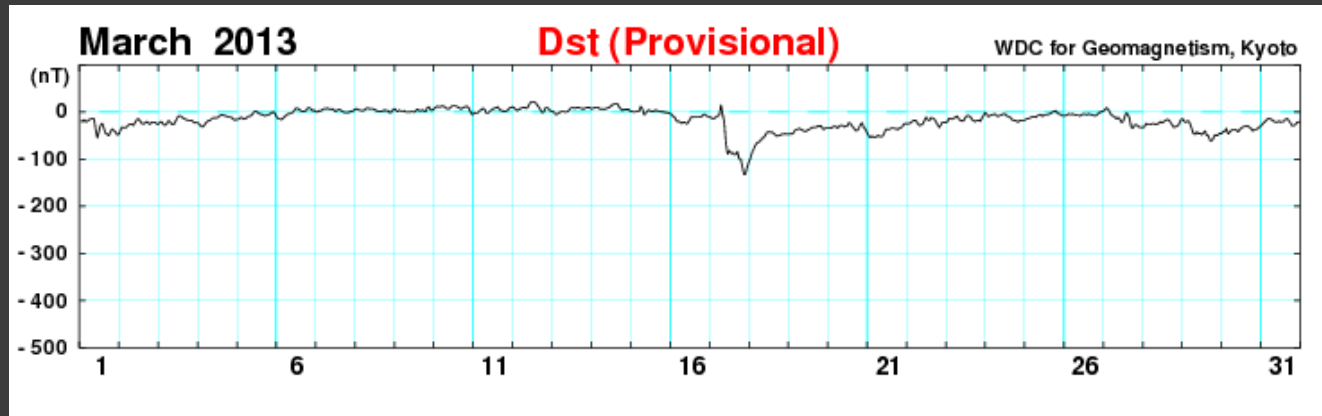


Selection of the quiet days

For the selection of the five days we use the following criteria:

- ▶ Min Dst index ≥ -30 nT for the day and the previous one
- ▶ Max AE index ≤ 250 nT for the day and the previous one

Storm time interval	5 Quietest days within the month	5 Quietest days prior to the event
16-20 March 2013	6/3, 7/3, 8/3, 25/3, 26/3	25/2, 27/2, 6/3, 7/3, 8/3



Data presentation

Autoscaled values of foF2 and hmF2 from Chilton ionosonde: Autoscaling error less than 0.7 MHz (Bamford et al., Radio Science, 2008)

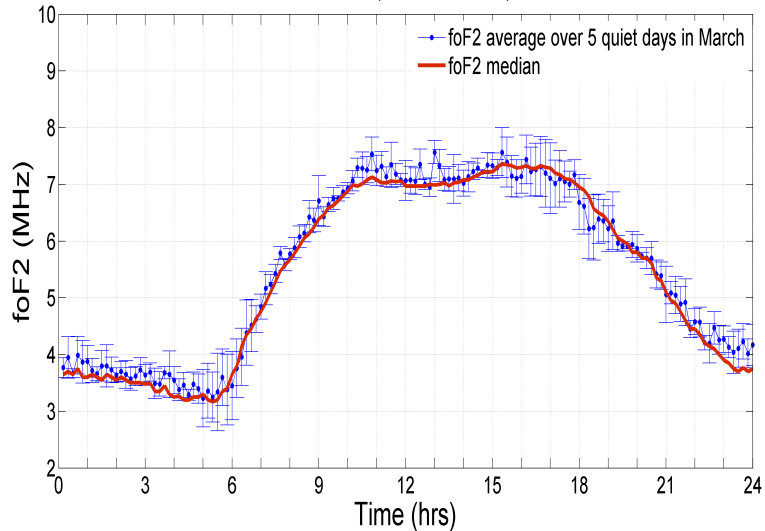
Ionospheric Station	Geographic longitude (°E)	Geographic latitude (°N)	GPS Station	Geographic longitude (°E)	Geographic latitude (°N)
Chilton	359.4	51.5	HERT	0.334	50.867

vTEC estimates used here are based on data from HERT GPS receiver. They are calculated from Receiver Independent Exchange Format (RINEX) files with 30 s sampling, using the single station solution proposed by Ciralo (2005) and Ciralo et al. (2007) that assumes that the ionosphere is a thin layer at 300 km altitude.

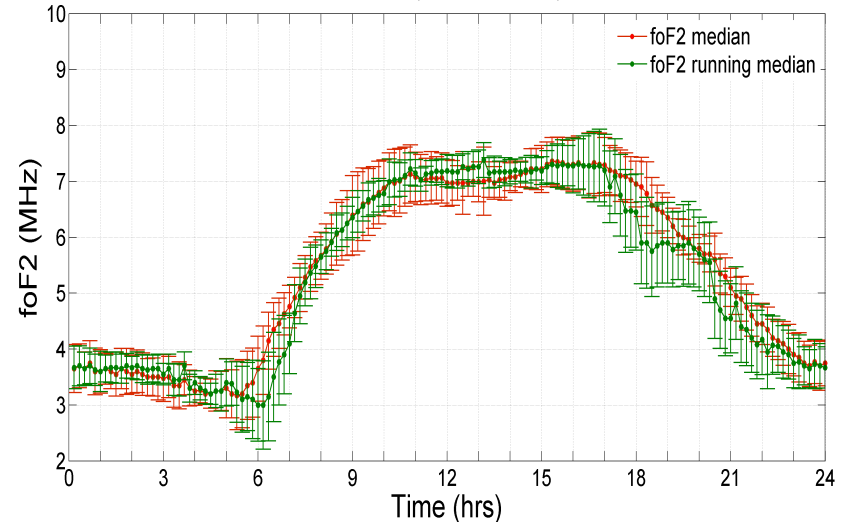


Chilton (foF2) – March 2013

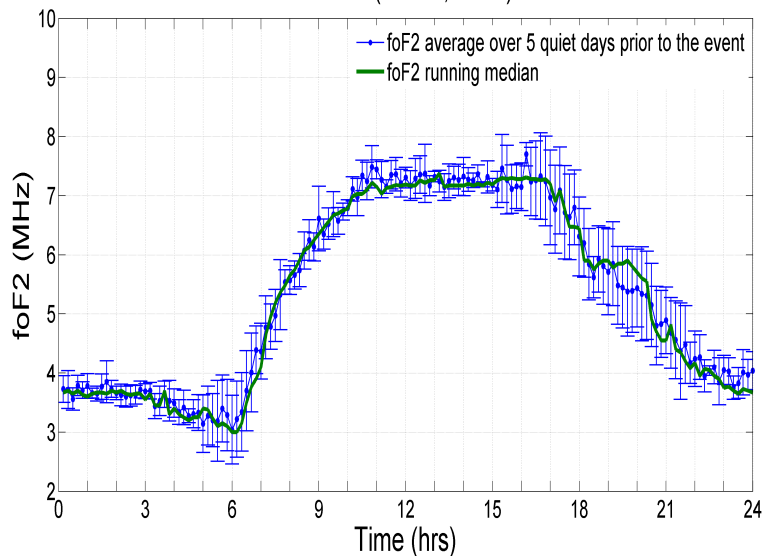
Chilton (March, 2013)



Chilton (March, 2013)



Chilton (March, 2013)



Error bars: STDs (Uncertainties mainly due to ionogram autoscaling errors and quiet time variability)

Mean STD monthly medians: **0.4 MHz**

Mean STD 5 quiet days in the month: **0.3 MHz**

Mean STD running medians: **0.4 MHz**

Mean STD 5 quiet days before the event: **0.4 MHz**

Chilton (foF2) – March 2013

$$\text{STD (\%)} = (\text{STD_foF2x} / \text{foF2x}) * 100$$

x: median, running median, average over 5 quiet days

STD (%) is estimated over each time of the day

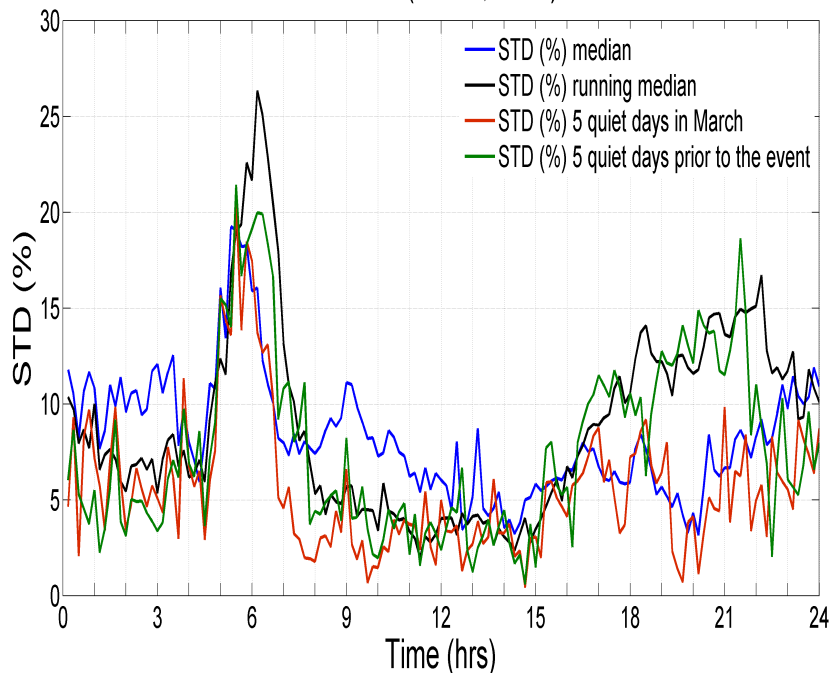
Mean STD (%) monthly medians: **8 %**

Mean STD (%) 5 quiet days in the month: **6 %**

Mean STD (%) running medians: **9 %**

Mean STD (%) 5 quiet days before the event: **8 %**

Chilton (March, 2013)

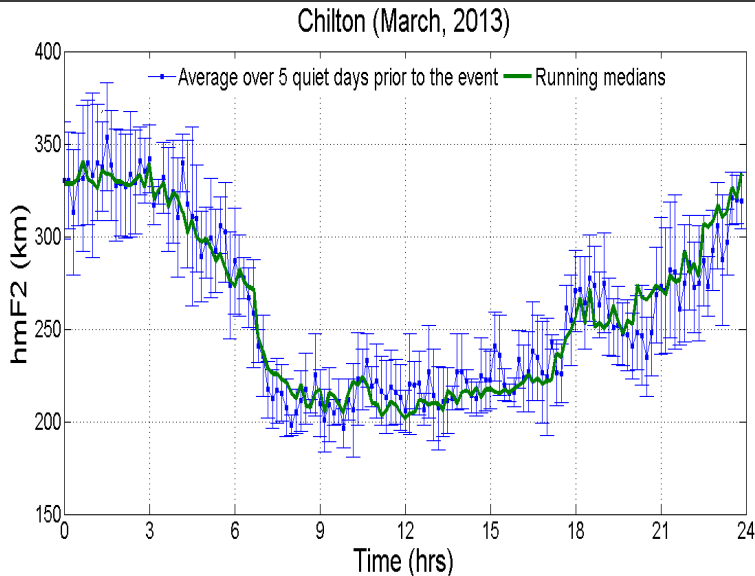
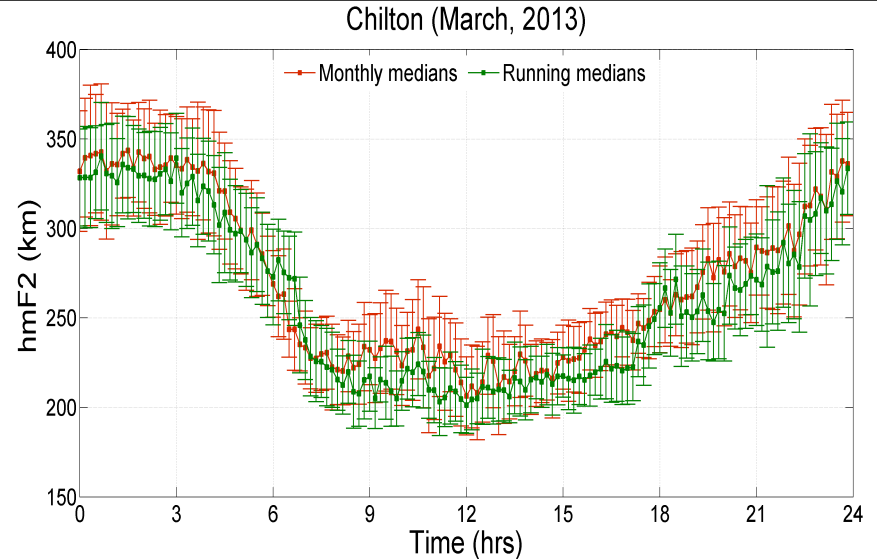
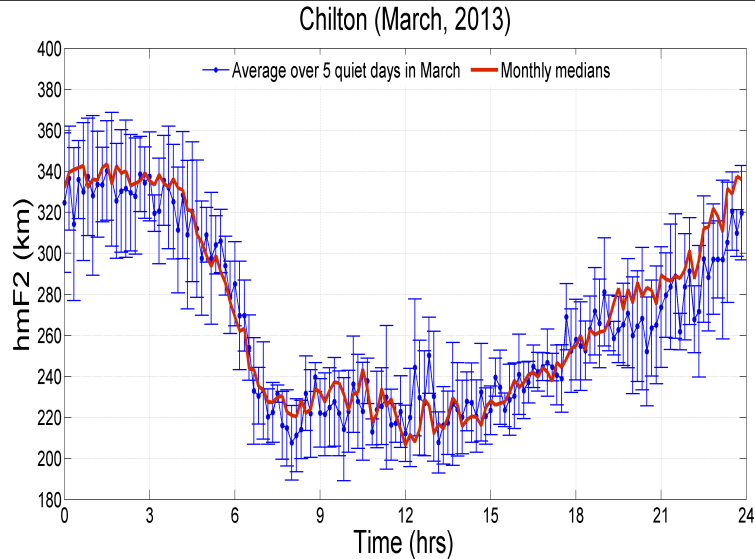


Information that may be extracted

- Local time dependence of the uncertainties: e.g., for the case under study here the uncertainties are significantly larger in dawn sector in all terms (for Chilton UT=LT)
- Monthly medians are comparable to the average of 5 quiet days within the month, while running medians are comparable to the average of 5 quiet days prior to the storm event. On average, all approaches may be considered comparable
- On average, ionospheric variations of about 10% wrt quiet conditions may be ignored.



Chilton (hmF2) – March 2013



Error bars: STDs (Uncertainties mainly due to ionogram autoscaling errors and quiet time variability)

Mean STD monthly medians: **26 km**

Mean STD 5 quiet days in the month: **21 km**

Mean STD running medians: **23 km**

Mean STD 5 quiet days before the event: **23 km**



Chilton (hmF2) – March 2013

$$\text{STD (\%)} = (\text{STD_hmF2x} / \text{hmF2x}) * 100$$

x: median, running median, average
over 5 quiet days

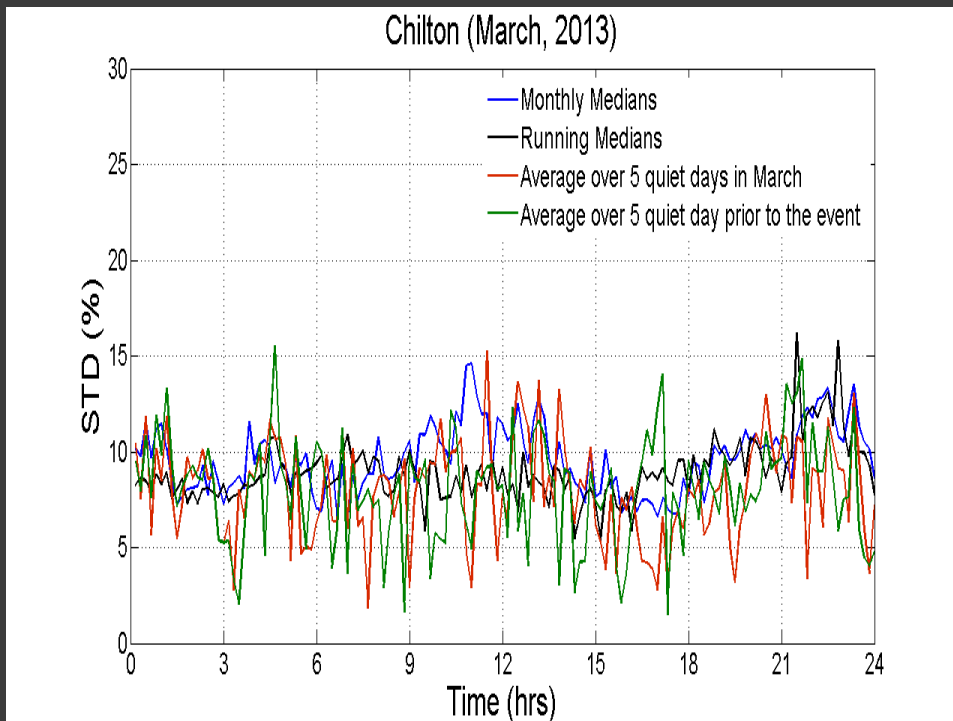
Mean STD (%) monthly medians: **10 %**

Mean STD (%) 5 quiet days in the month: **8 %**

Mean STD (%) running medians: **9%**

Mean STD (%) 5 quiet days before the event: **8%**

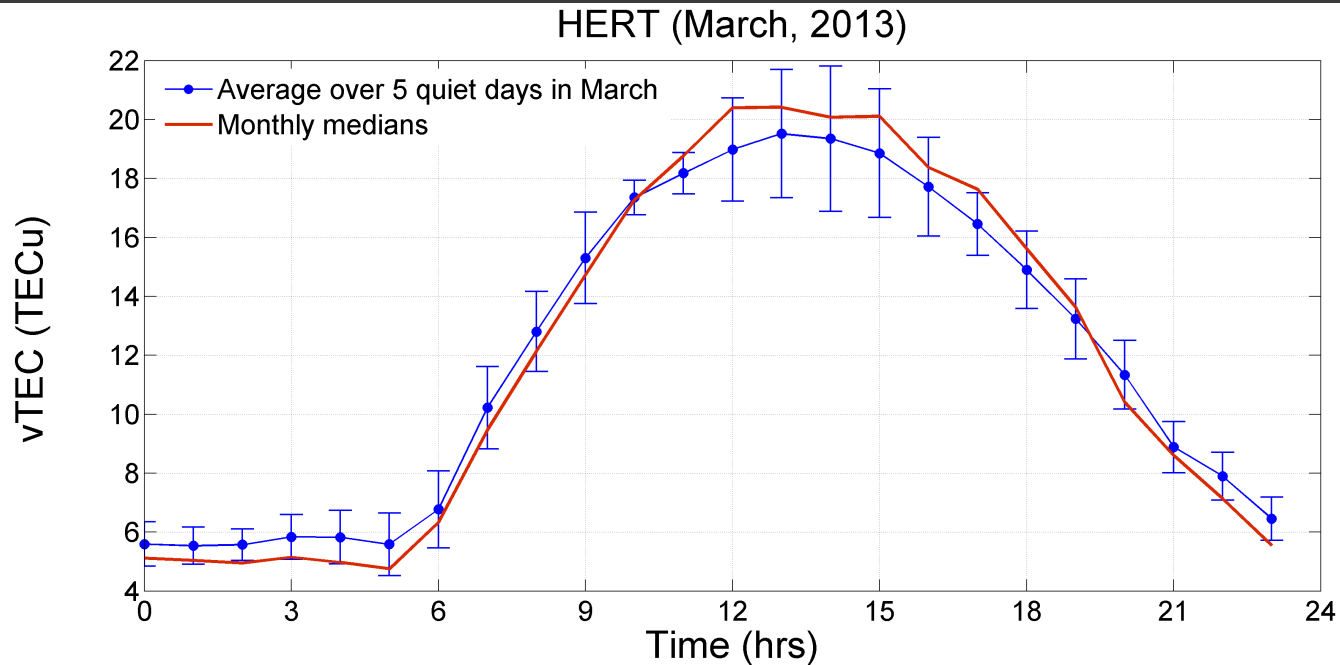
STD (%) is estimated over each time of
the day



Extracted information

- Some local time dependence of the uncertainties in monthly medians in the prenoon sector (for Chilton UT=LT)
- On average, all approaches may be considered comparable
- On average, ionospheric variations of about 10% wrt quiet conditions may be neglected.

HERT (vTEC) – March 2013



Mean STD monthly medians:

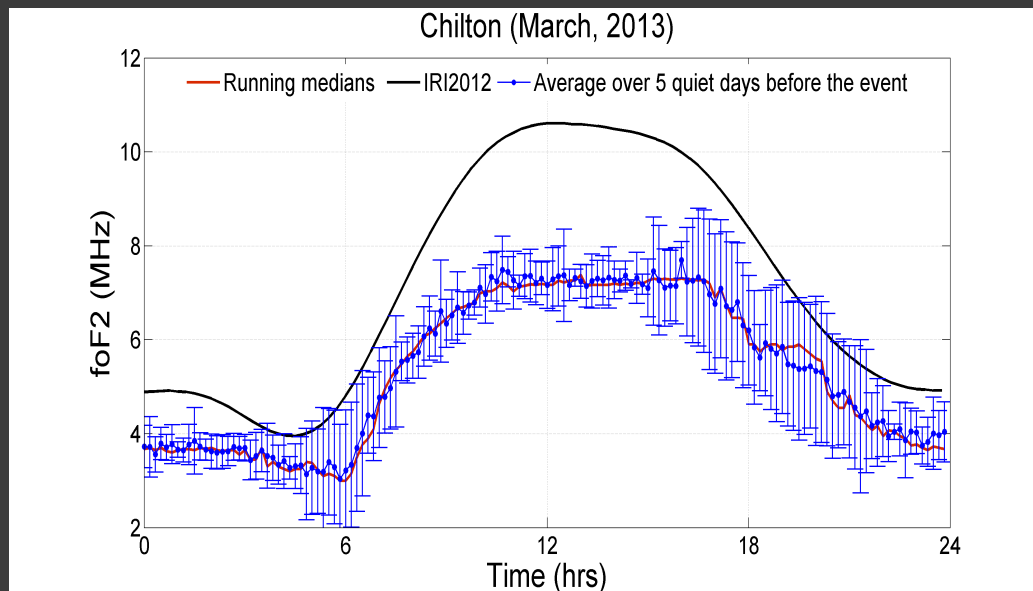
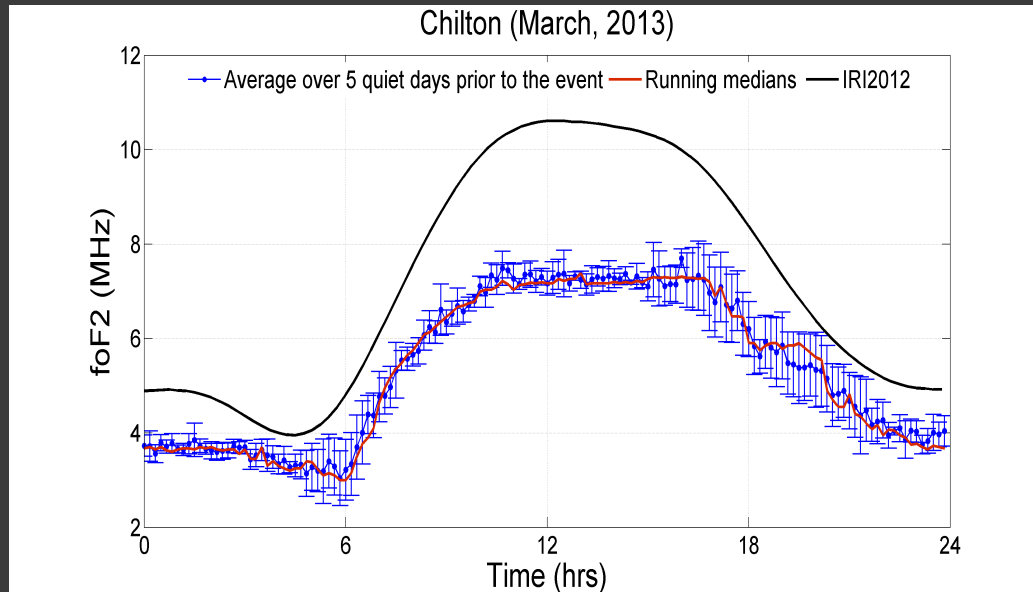
Mean STD 5 quiet days in the month: **1.2 TECu**

Mean STD (%) monthly medians:

Mean STD (%) 5 quiet days in the month: **11 %**

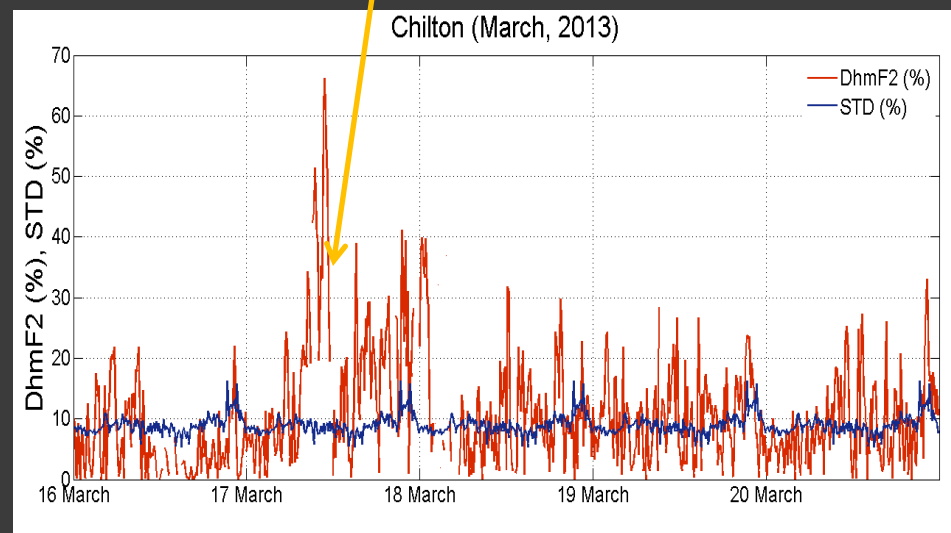
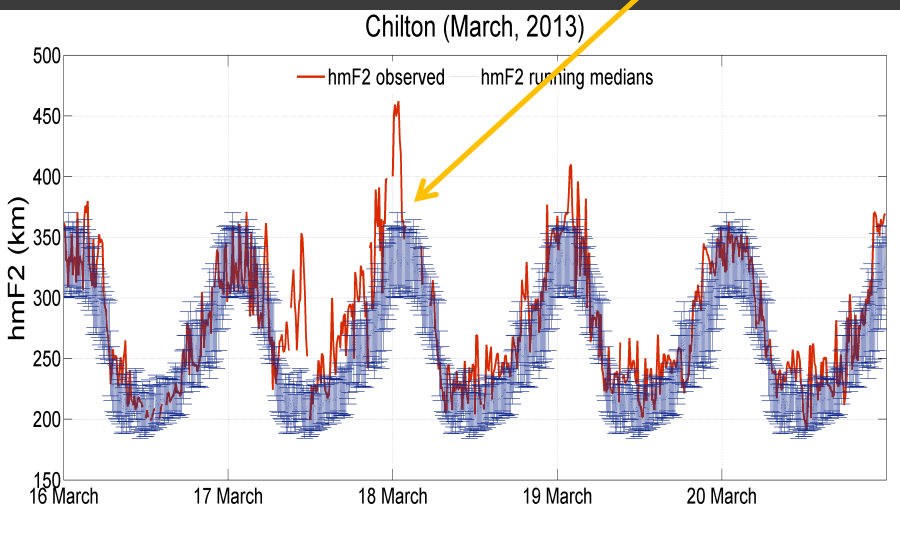
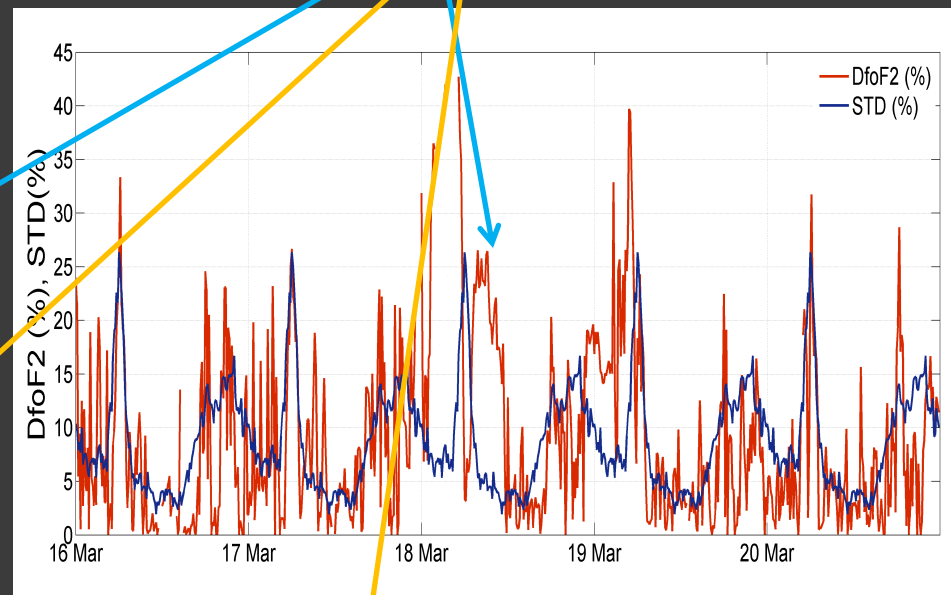
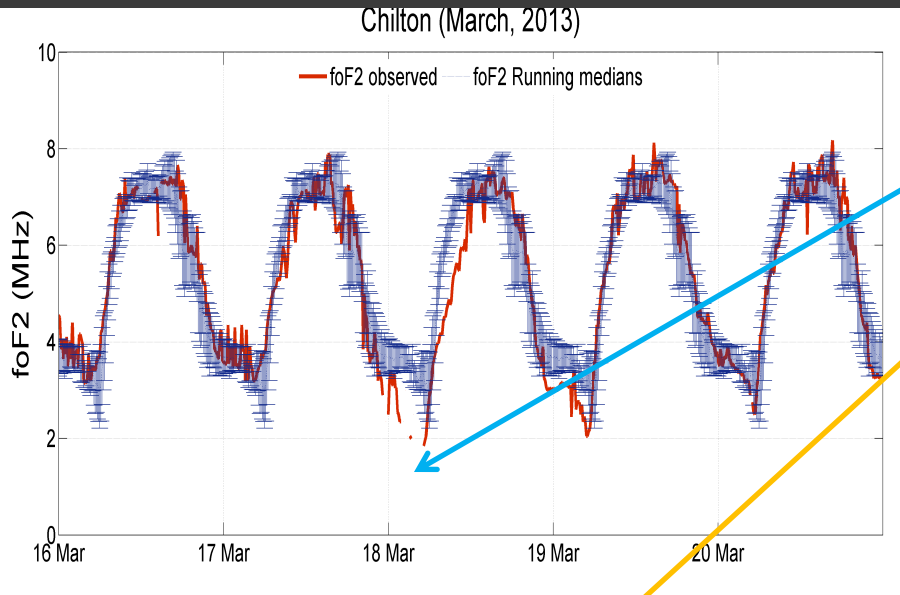


Chilton (foF2) – March 2013



Chilton (foF2, hmF2) – March 2013

Storm Impact



▶
$$DfoF2(\%) = \left[\frac{foF2_{obs} - foF2_{runningmedian}}{foF2_{runningmedian}} \right] * 100$$