#### Transitioning AMPS to the CCMC

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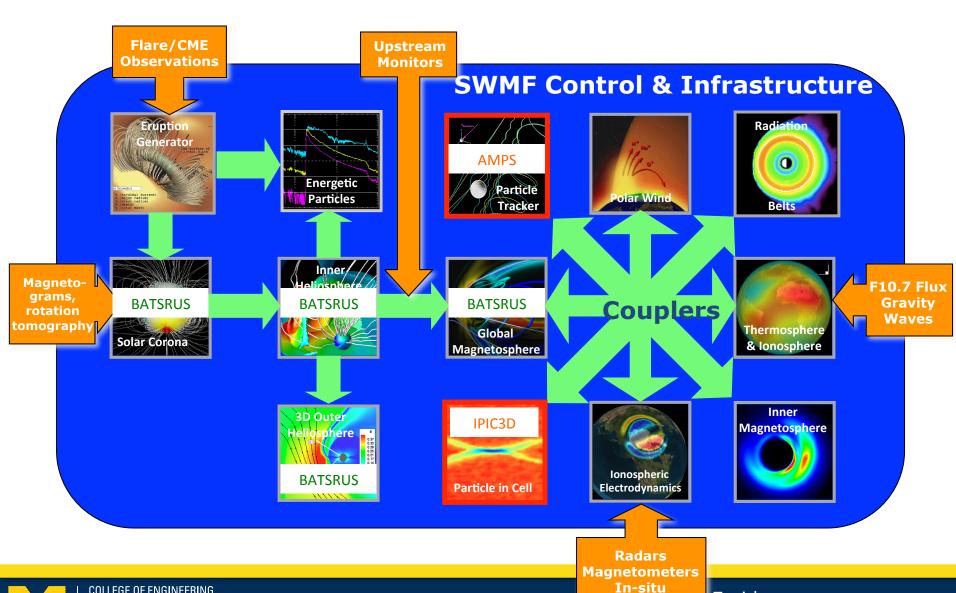
University of Michigan

### Particle Based Modeling

- Non thermal component of the simulated environment
  - Energetic particle radiation environment in Earth's magnetosphere
  - Charged dust, etc...

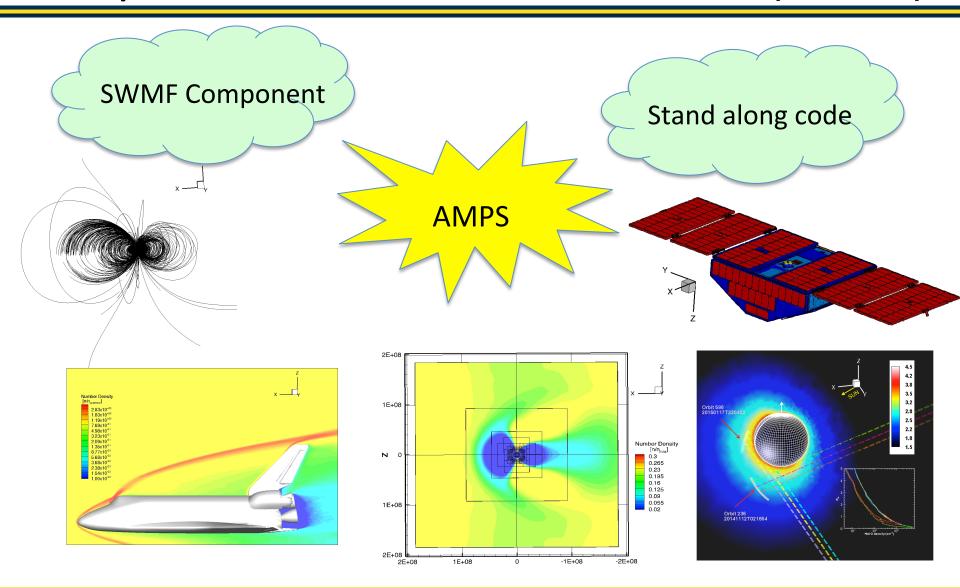
2. Illustration of the particle dynamics on a kinetic level

#### AMPS in the SWMF



9<sup>th</sup> CCMC Community Workshop, April 2018

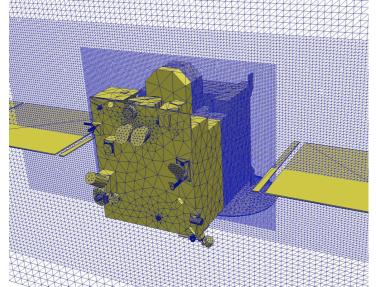
# Adaptive Mesh Particle Simulator (AMPS)



#### Adaptive Mesh Particle Simulator (AMPS)

#### Applications

- Comets, planetary satellites and exospheres
- Tracking of the ions and charged dust
- Integrated with SPICE and SWMF
- Optimizations
  - Local time step and particle weight
  - Individual particle weight correction



- Adaptive Mesh Refinement (AMR) with cuτ-ceiis
- Parallel implementation
  - Domain decomposition
  - Static (volume, cell number ) and dynamic load balancing (particle number, execution time)
  - Hybrid implementation: MPI (between nodes) + OpenMP (within each

#### Recent Development

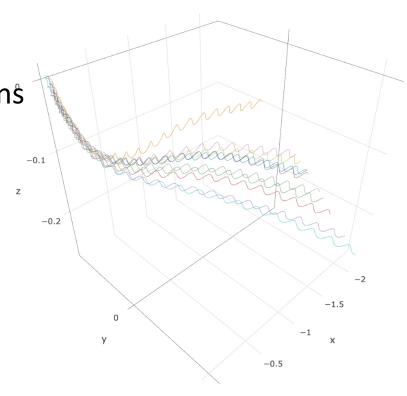
- 1. Backward particle tracking
- Calculation of the cutoff rigidity, energy spectra, and flux of the energetic particles

#### **Under construction:**

1. Field solver for adding implicit PIC capabilities

#### Transitioning to the CCMC

- Analyze charged particle trajectories moving in plasma
  - Post processing tool: use pre-calculated plasma and magnetic field
- Import of the plasma flow simulations into AMPS
  - Time-dependent results of the plasma modeling is saved in files
- Output
  - Individual particle trajectories
  - Samples of the distribution functions
  - Energy spectra
  - Flux



#### Transitioning to the CCMC

Web based user interface is under testing

	← → C ☐ ccmc.gsfc.nasa.gov/requests/PP/user_registration.php	Q 🖒 H 🚺 \land 🗏
	COMMUNITY COORDINATED MODELING CENTER Related Links   1	Frequently Asked Questions   Community Feedback   Downloads   Sitemap
	About   Models at CCMC   Request A Run   View Results   Instant Run   Metrics and Validation   Education   RT Simulations	
	Step 1: Generate a Registration Number for each Requested Run and select the Model.  Registration Number  The Registration Number is composed of your first name (GivenName), your last name (FamilyName), date (yyyymmdd),model type (GM - Globa Solar/Heliosphere, PP - Postprocessing), and run identification number (RunNumber):	l Magnetosphere, IT - Ionosphere/Thermosphere, SH -
	GivenName_FamilyName_mmddyy_RunNumber, e.g., Lutz_Rastaetter_20160101_PP_1.	
	At the present time you are allowed to make up to 4 different submissions on the the same date (mmddyy). For each new submission made on the "4", or "5"). Multiple submissions made on the same date with the same Run Number will overwrite the previous submission. You can use this feat cancel or modify your submission at later date, please contact the CCMC staff:	same date you need to choose a new <b>Run Number</b> ("1", "2", "3", ure to resubmit the request on the same date. If you decide to
	e-mail: ror@ccmc.gsfc.nasa.gov tel.: Masha Kuznetsova (+1) 301-286-9571, Lutz Rastaetter (+1) 301-286-1085.	
	Please have registration numbers when making inquiries about your requests. You will need your registration number to view the results when the	simulations have finished.
	First (Given) Name: (required)	
	Last (Family) Name: (required)	
	E-mail: (required)	
	Run Number: 1 : Select a different Run Number if you have already requested any Magnetosphere Postprocessing run today (unless you want to overwrite the requested any Magnetosphere Postprocessing run today).	est)
	Title/Introduction Key words	
<b>4                                    </b>	Global Magnetosphere Post-Processing (PP) model:	
AMPS	<ul><li>MPS (version as of 2016/02/02)</li></ul>	
	⊚ Separatrix and Null Point Finder	
	⊕ Glocer (version as of 2016/02/02)	
	Submit Reset	
	Next page: Existing Run To Work on and Options	

#### Transitioning to the CCMC

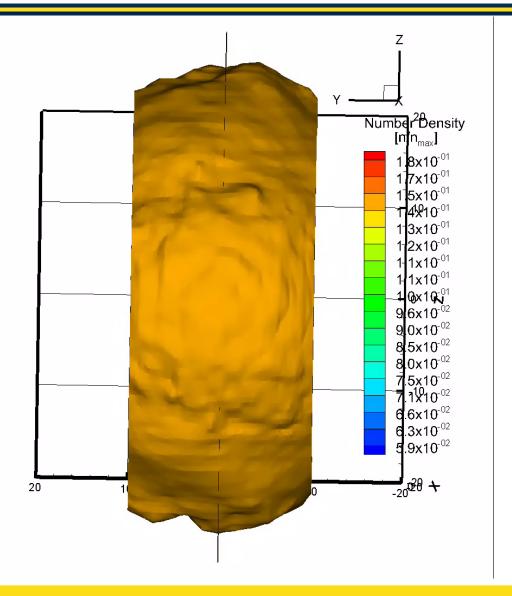
Initial parameters of the simulated trajectories

Models at CCMC   Request A Run   View Results   Instant Run   Metrics and	Validation   Education   RT Simulations
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# GCRs in Earth's Magnetosphere

#### Energy dependence:

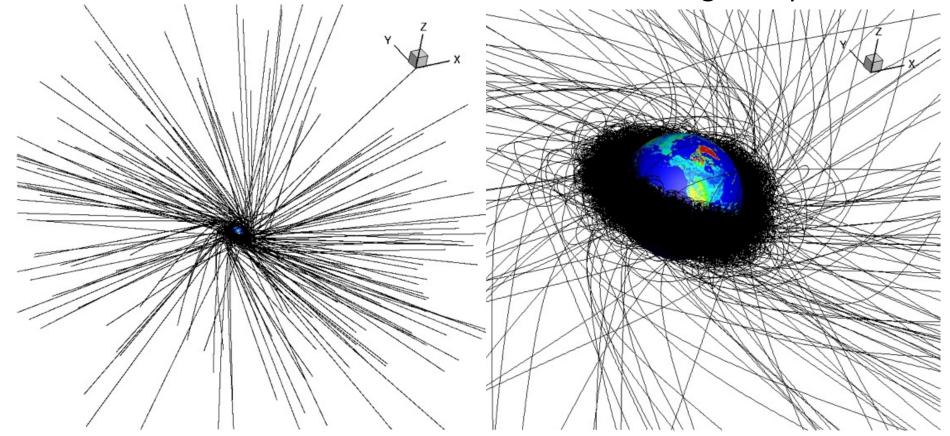
- Cutoff at low latitudes
- Shielding by the magnetosphere
- Local time dependence
  - Becomes significant with an increase of the geocentric distance
- Modeling
  - Global modeling of the entire magnetosphere is important



## **Backward Particle Tracking**

Small and Large scale energetic particle trajectory integration

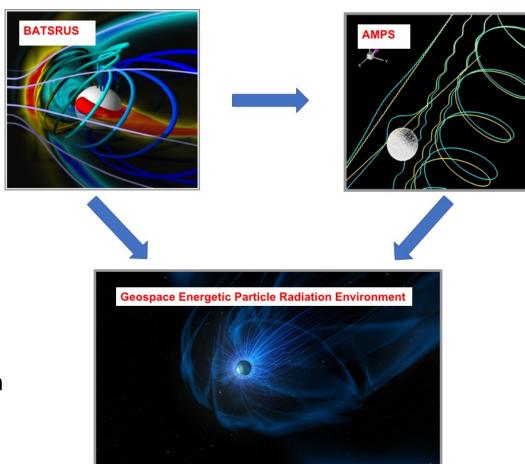
Radiation environment in the entire Earth's magnetosphere



# Modeling Geospace Radiation Environment

1. Transport of the energetic particles in a realistic geomagnetic field

- 2. Results at the top of the atmosphere
  - **Cutoff rigidity**
  - Energy spectrum
  - Flux
- 3. Practical applications
  - Radiation environment in Earth's atmosphere, at LEO, MEO, and GEO



#### Conclusion

- Adaptive Mesh Particle Simulator
  - Self-sufficient Direct Simulation Monte Carlo model
  - A component of SWMF
  - A particle tracking tool in CCMC
- Application
  - Tracking of the charged particle trajectories, and sampling of distribution functions
  - Kinetic modeling of the charge particle dynamics moving in ambient plasma
  - Geospace radiation environment
- Available in the CCMC in a test mode