



Air Force Research Laboratory



Integrity ★ Service ★ Excellence

WSA Model Update

**Community Coordinated
Modeling Center Workshop**

April 12, 2016

Annapolis, Maryland

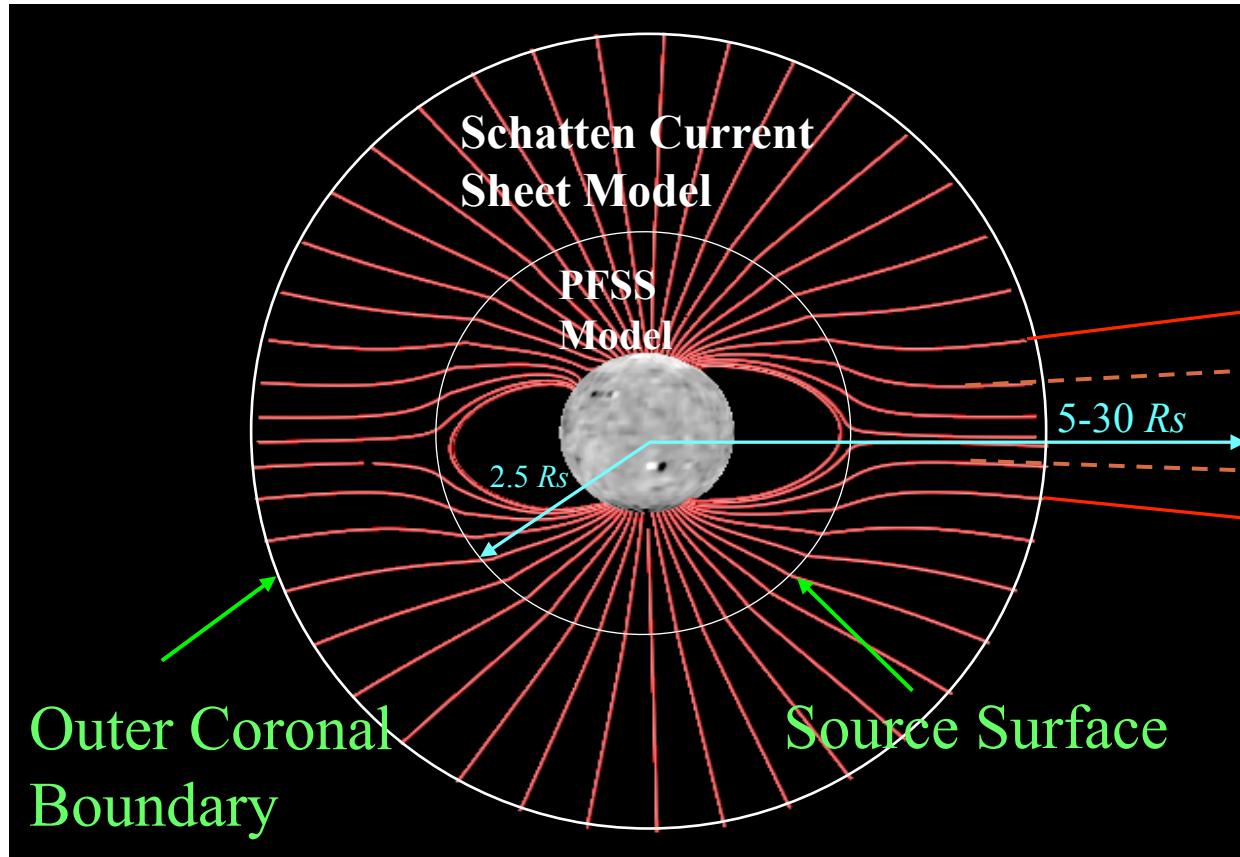
**C. Nick Arge¹, David MacKenzie²,
& Kathleen Shurkin³**

1. AFRL/Space Vehicles Directorate, Kirtland AFB, NM
2. Atmospheric and Environmental Research (AER)
3. Boston College





Wang-Sheeley-Arge (WSA) Coronal & Solar Wind Model



Solar Wind Models such as:

- 1) WSA 1D Kinematic
- 2) ENLIL
- 3) LFM-Helio
- 4) MS-FLUKSS
- 5) HAF

(5-30Rs to 1AU)

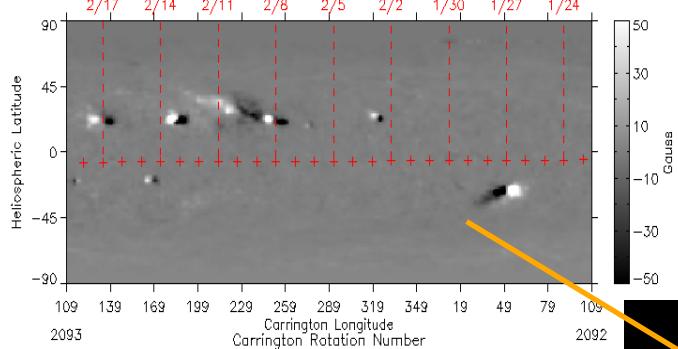
- Wang-Sheeley-Arge (WSA) model - combined empirical and physics based model of the corona and solar wind.
- Improved version of the original Wang & Sheeley model developed at NRL.
- Operational at NOAA/NCEP and NASA/CCMC.



WSA PFSS+SCS Coronal Solution

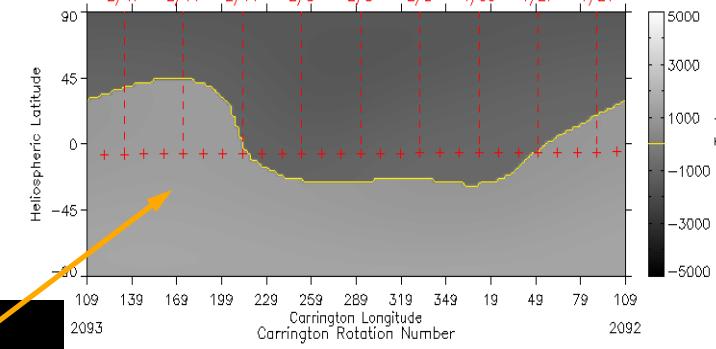


MODEL INPUT: Observed Photospheric Field



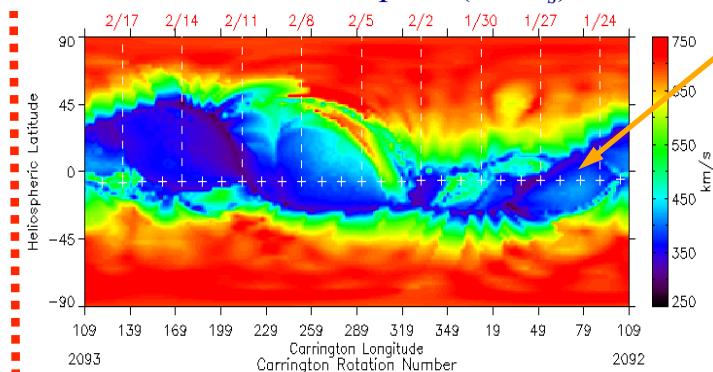
MODEL OUTPUT

Field at Outer Coronal Boundary ($5.0 R_s$)

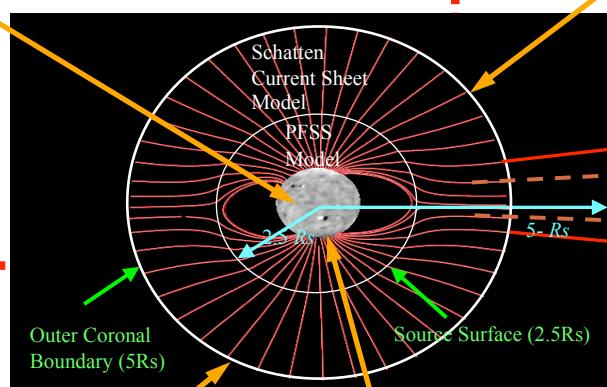


MODEL OUTPUT

Predicted Solar Wind Speed ($5.0 R_s$)

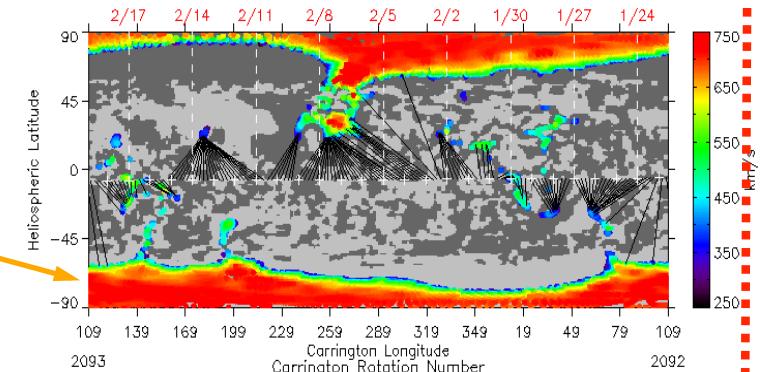


MODEL OUTPUT



Solar Wind Model
(e.g., WSA 1D Kinematic model, Enlil,
LFM-Helio, MS-FLUKSS & HAF)
($5\text{-}30R_s$ to 1AU)

Derived Coronal Holes ($1.0 R_s$)





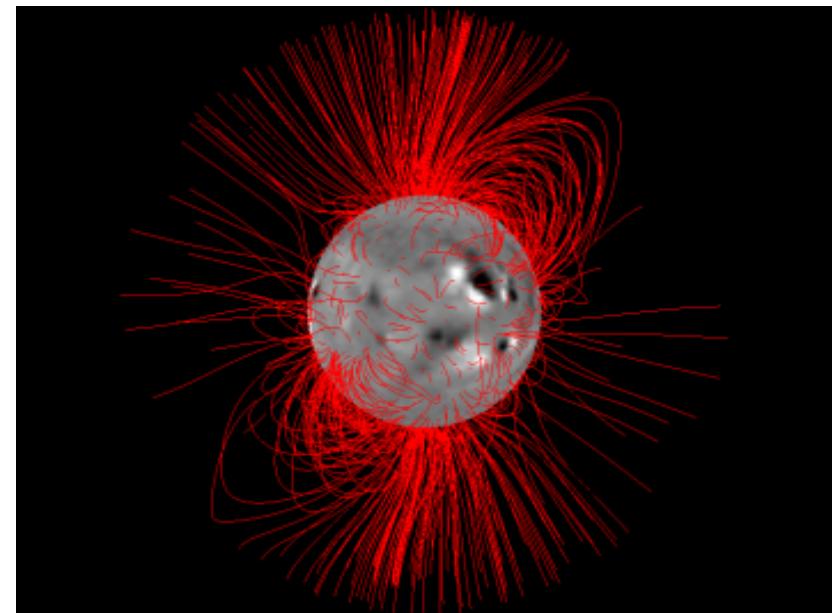
WSA 4.0 Available Soon!



Key Features:

1. Field line tracing parallelized
2. Compatible with
 - NSO (VMS, GONG), WSO & MWO
 - ADAPT (VSM, GONG, HMI)
 - Understands multi-realization input files
3. Can be run in the following modes:
 - PFSS
 - Coupled PFSS+Schatten Current Sheet (SCS)
 - Traditional or Improved interface between PFSS & SCS models.
 - Improved: minimizes “kinking” at interface.
4. IDL & Perl scripts replaced with Python
5. Forecasts solar wind speed and IMF polarity at
 - L1, STEREO A & B, Ulysses, & all inner planets
 - Easy to add other positions/satellites
6. Retuning empirical solar wind speed relationship
7. New field line tracing package

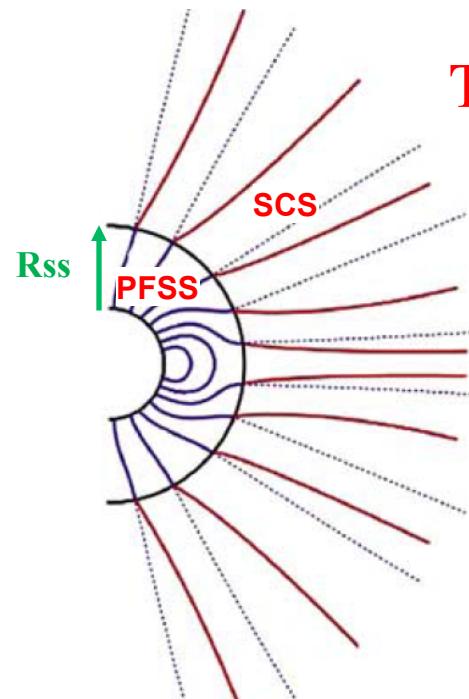
**Wang-Sheeley-Arge (WSA)
(Coronal Model)**



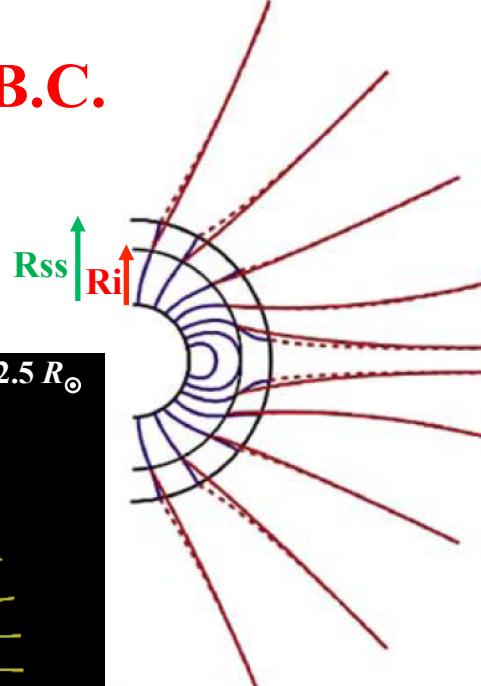
Input: Global maps of photospheric magnetic field
(e.g., ADAPT)



Improved Interface B.C. Between PFSS & SCS Models

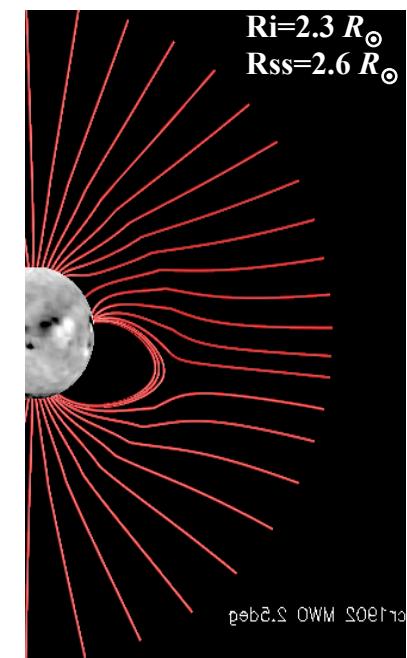
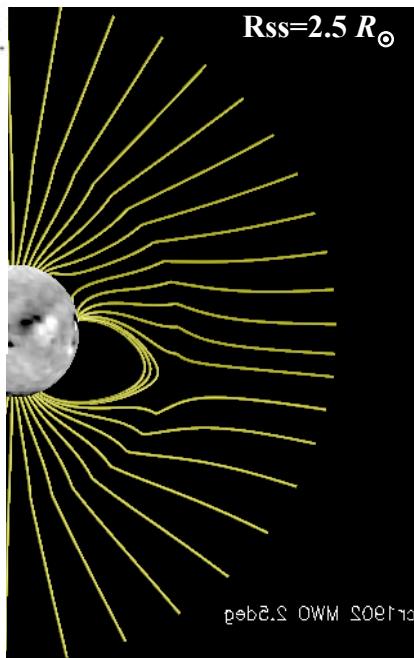


Traditional B.C.



Improved B.C.

WSA 4.0 allows user to choose between standard interface B.C. (left) and improved one (right)



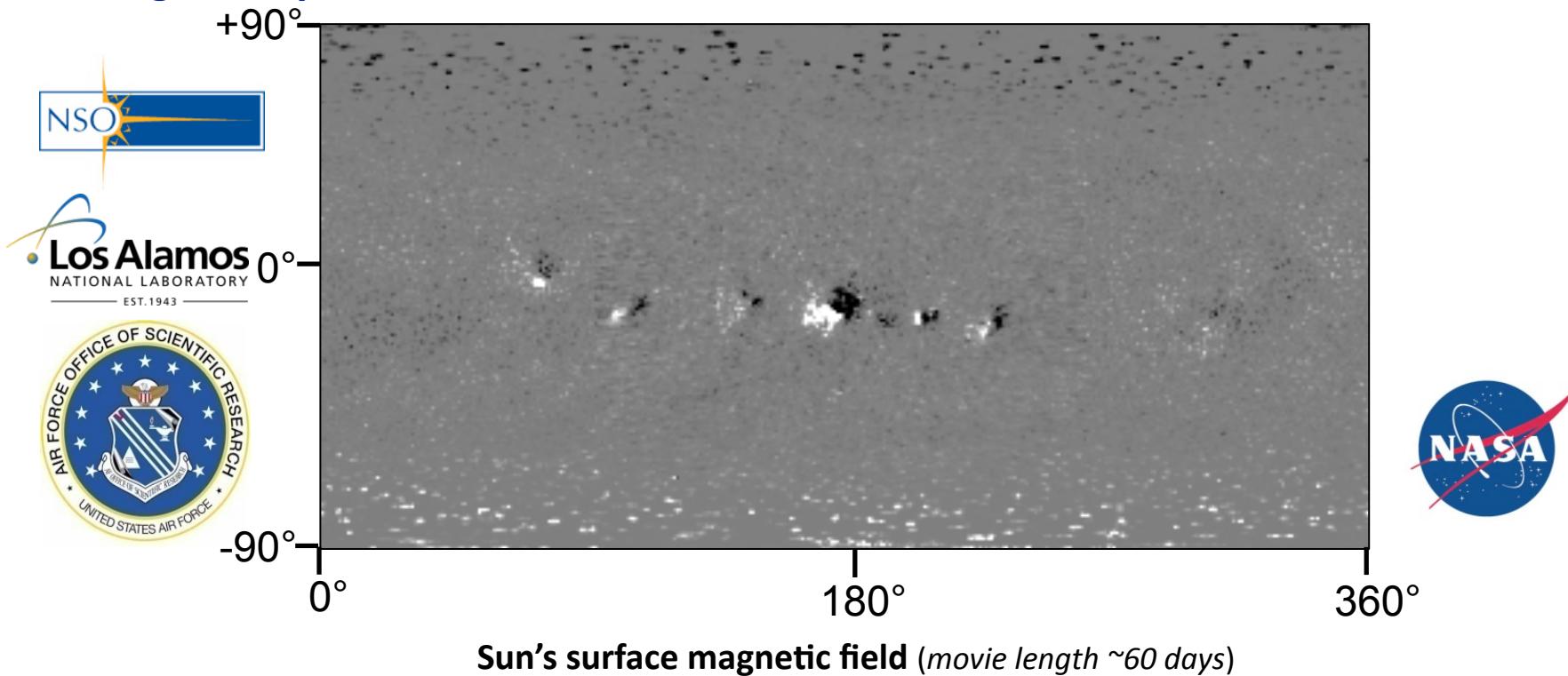
McGregor et al., JGR, 2008



Air Force Data Assimilative Photospheric Flux Transport (ADAPT) Model



1. Evolves solar magnetic flux using well understood transport processes where measurements are not available.
2. Updates modeled flux with new observations using *data assimilation methods*
 - Rigorously takes into account model & observational uncertainties.



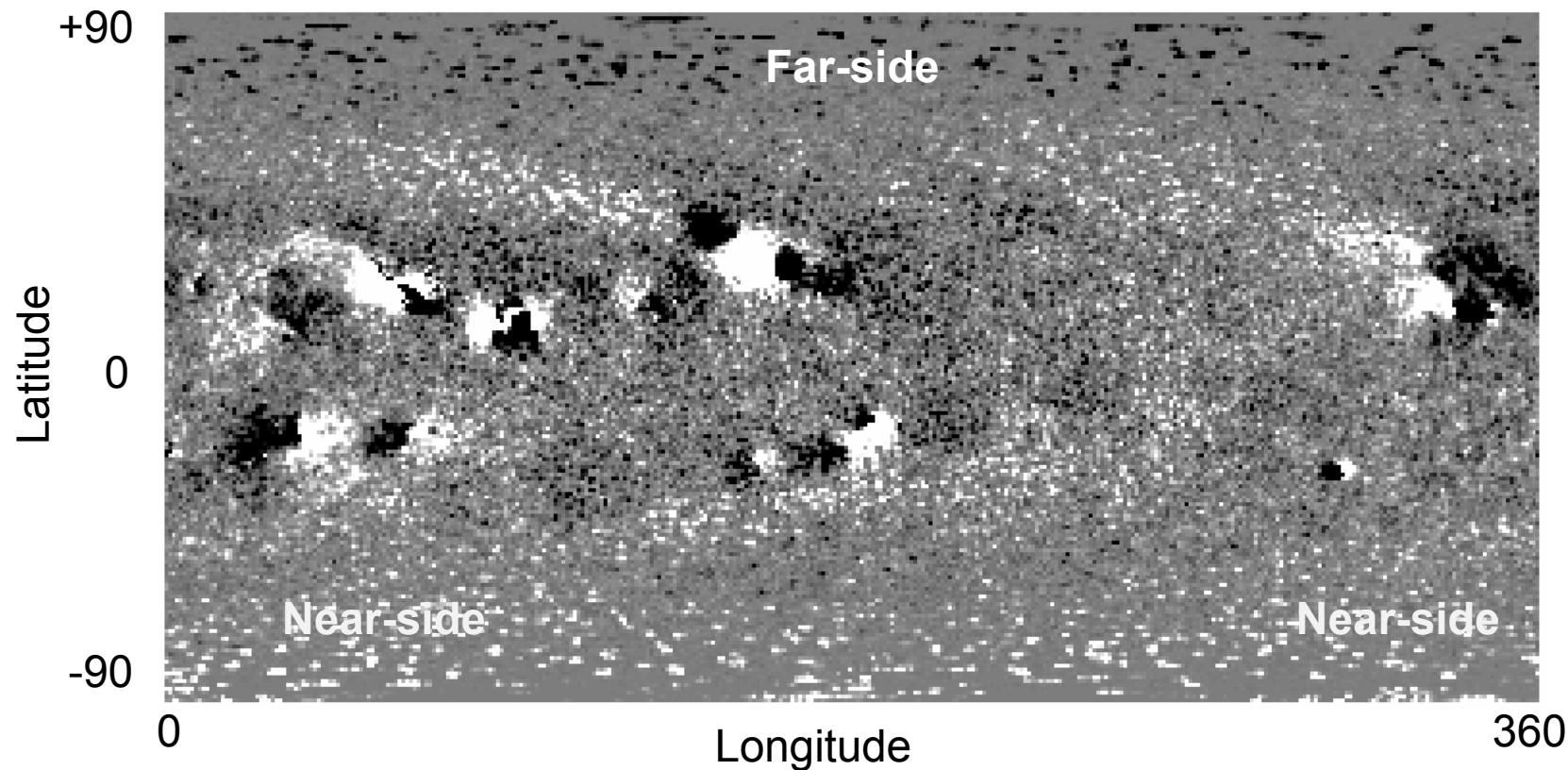
Provides more realistic estimates of the instantaneous global photospheric magnetic field distribution than those provided by traditional synoptic maps.



ADAPT Solution Ensemble

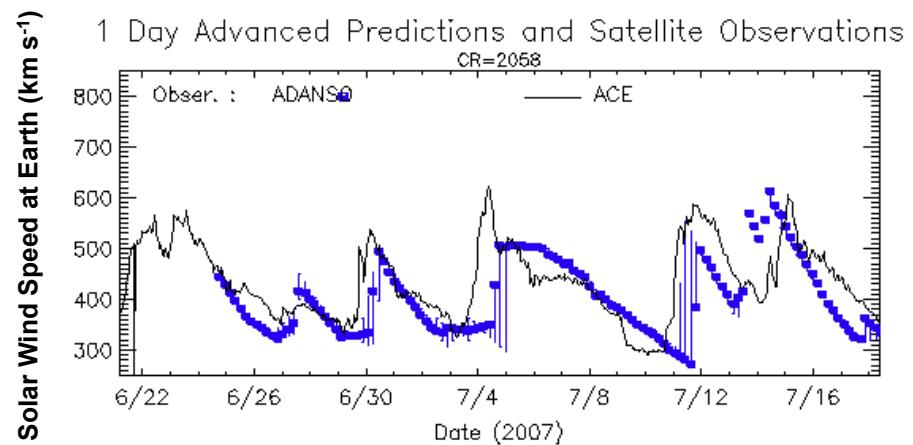
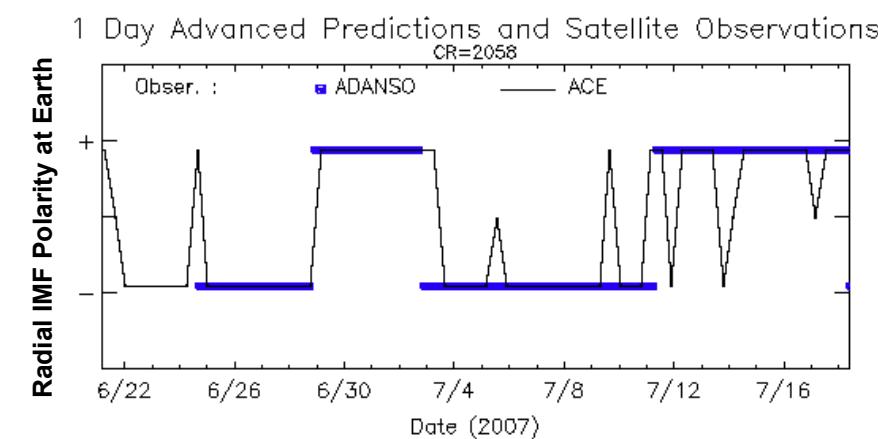
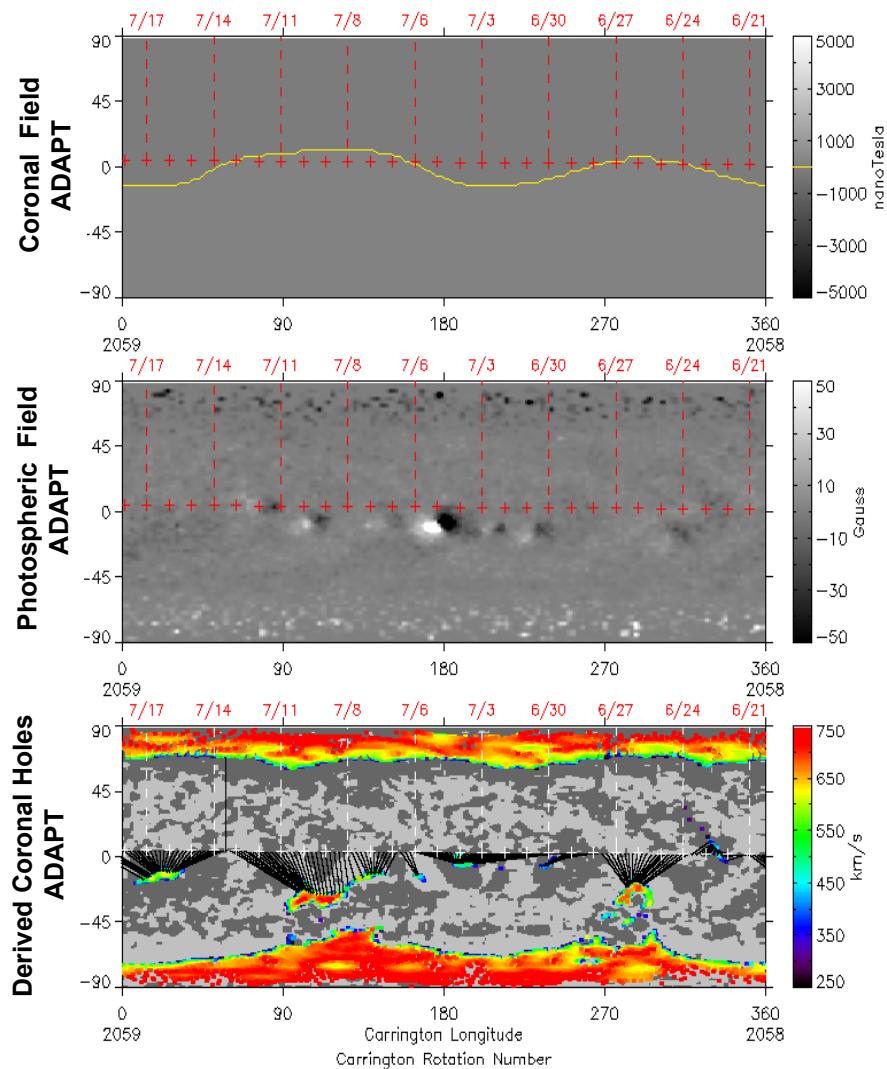


ADAPT generates an ensemble of global photospheric magnetic field maps for each given moment in time



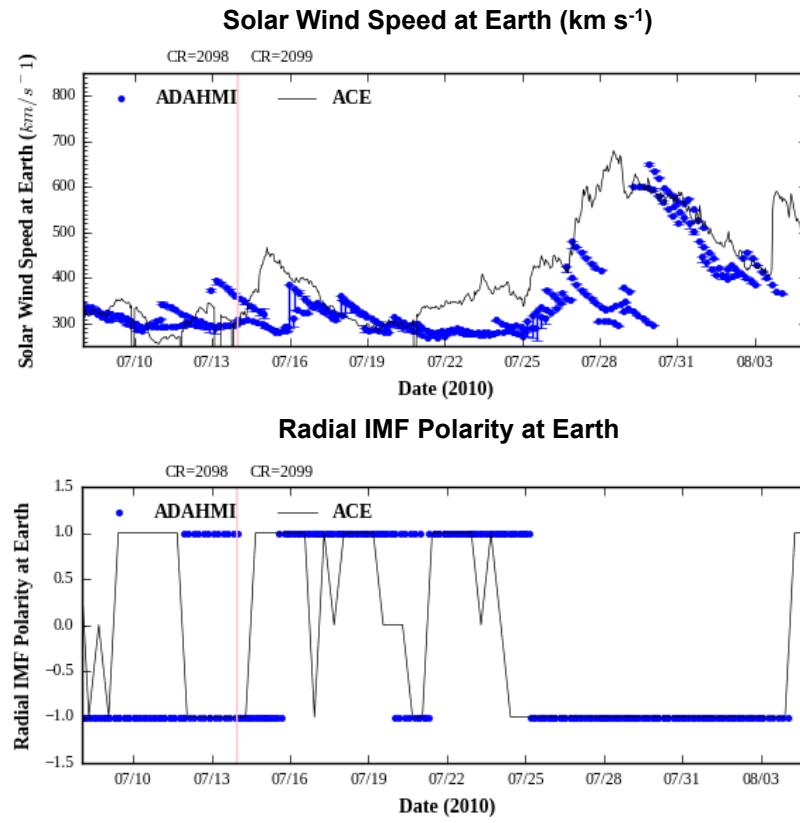
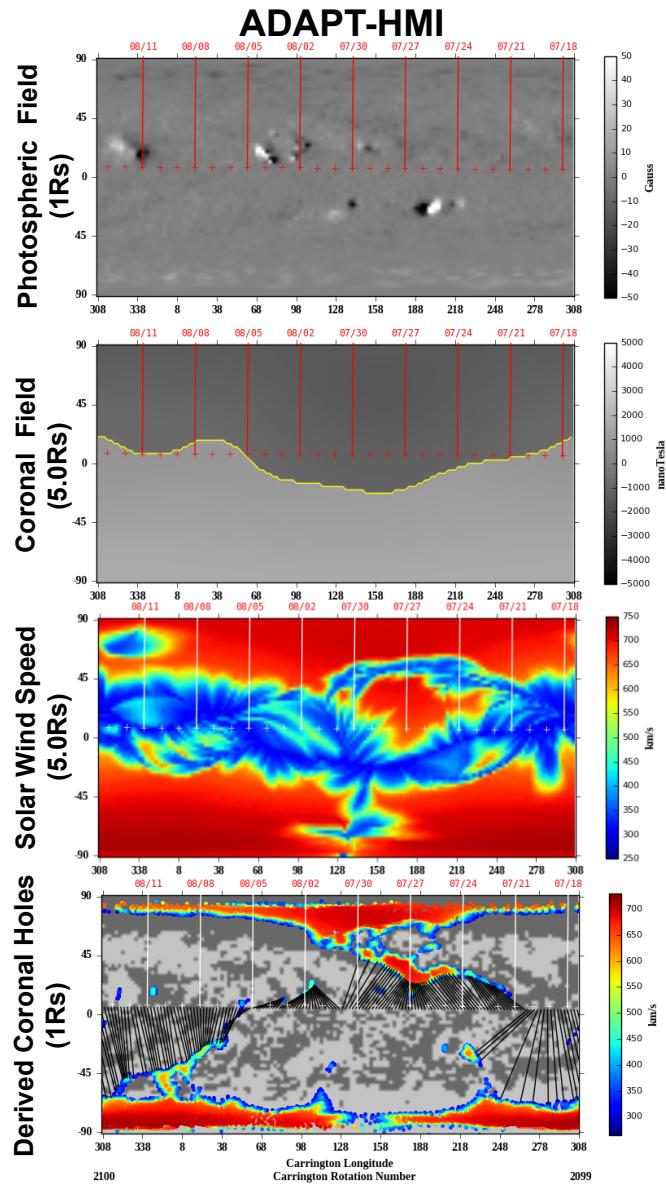


WSA Coronal & Solar Wind Solutions using the 12 ADAPT Realizations for June 21, 2007 (Start of CR2058)





WSA 4.0 ADAPT-HMI Coronal Solutions & Solar Wind Solutions



Created 2016 February 29 1704 UTC

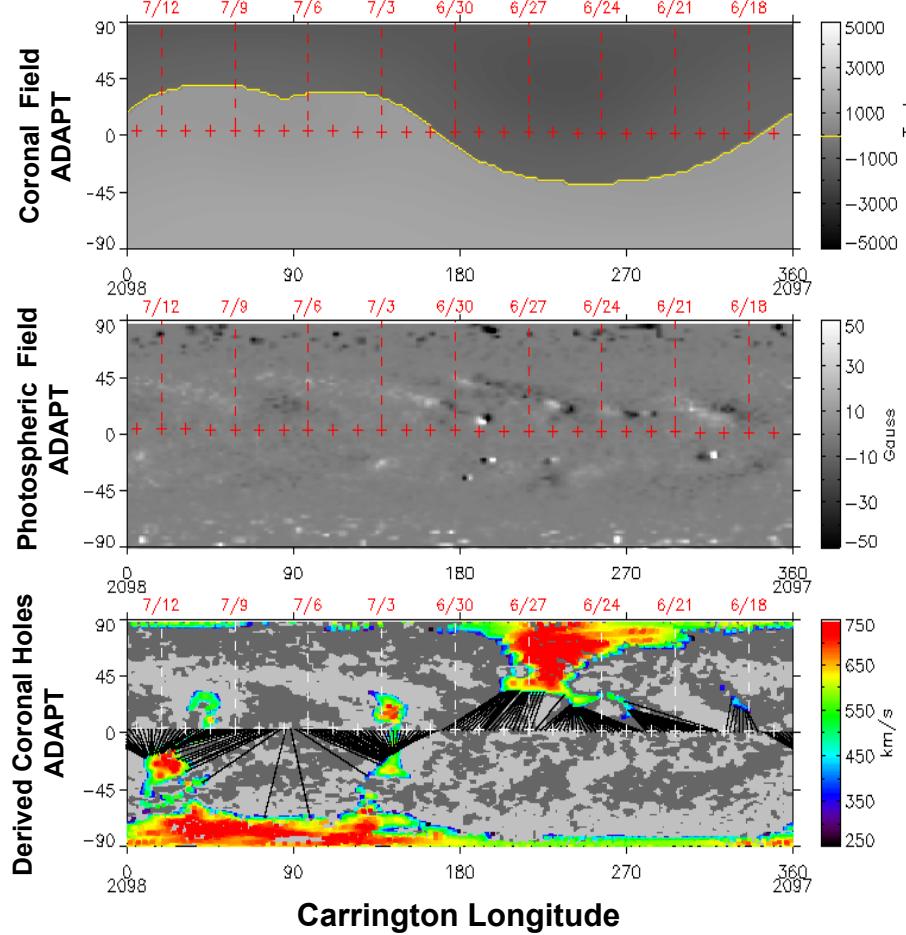
- *ADAPT-HMI maps fully integrated into WSA 4.0*
- *Tuning the solar wind empirical relationship*



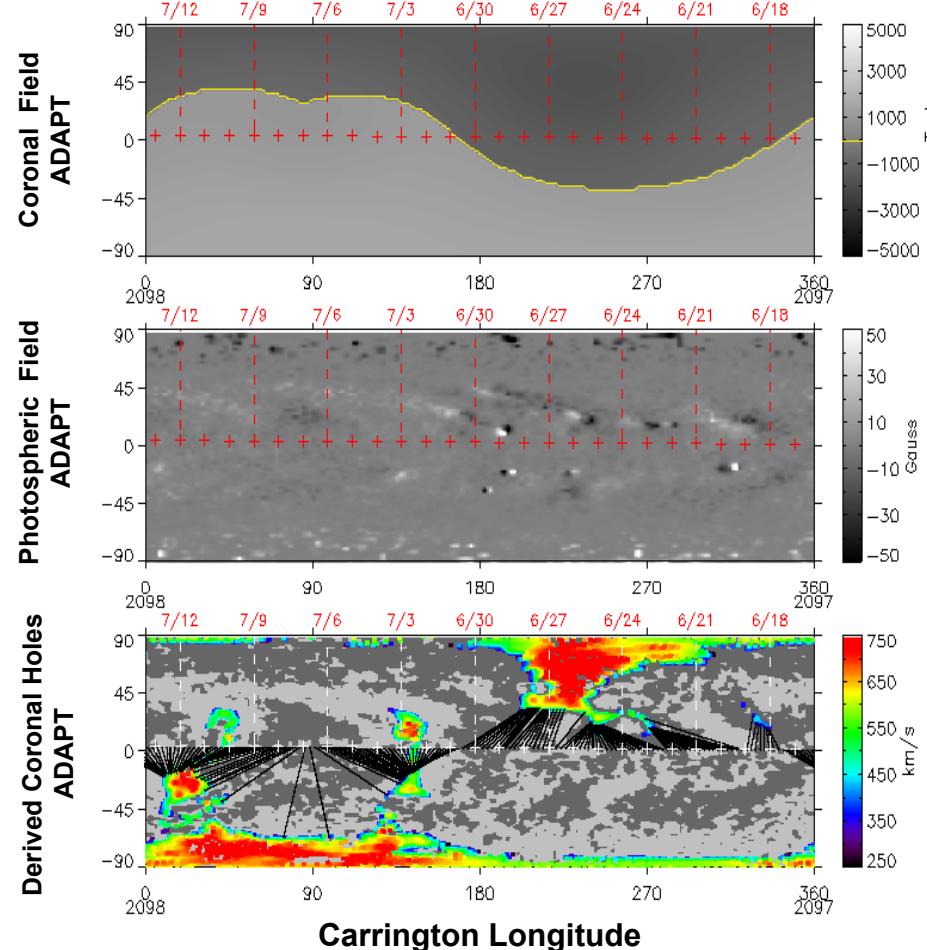
Time Evolution of the Photosphere & Coronal (With & Without Far-Side Active Region Inserted)



Without Far-Side Active Region Inserted



With Far-Side Active Region Inserted



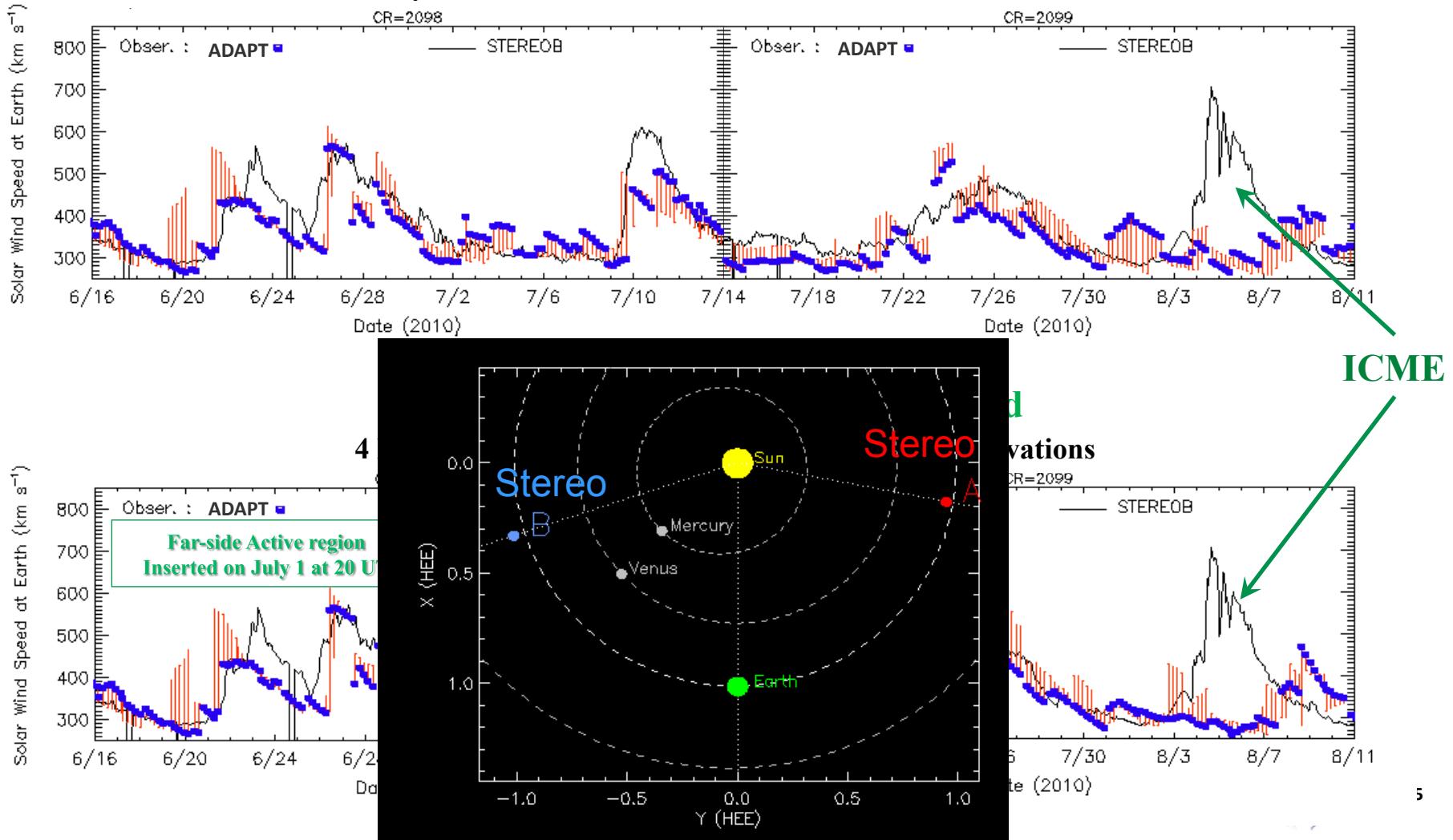


WSA Solar Wind Speed vs Observations at STEREO B (With & Without Far-Side Active Region Inserted)



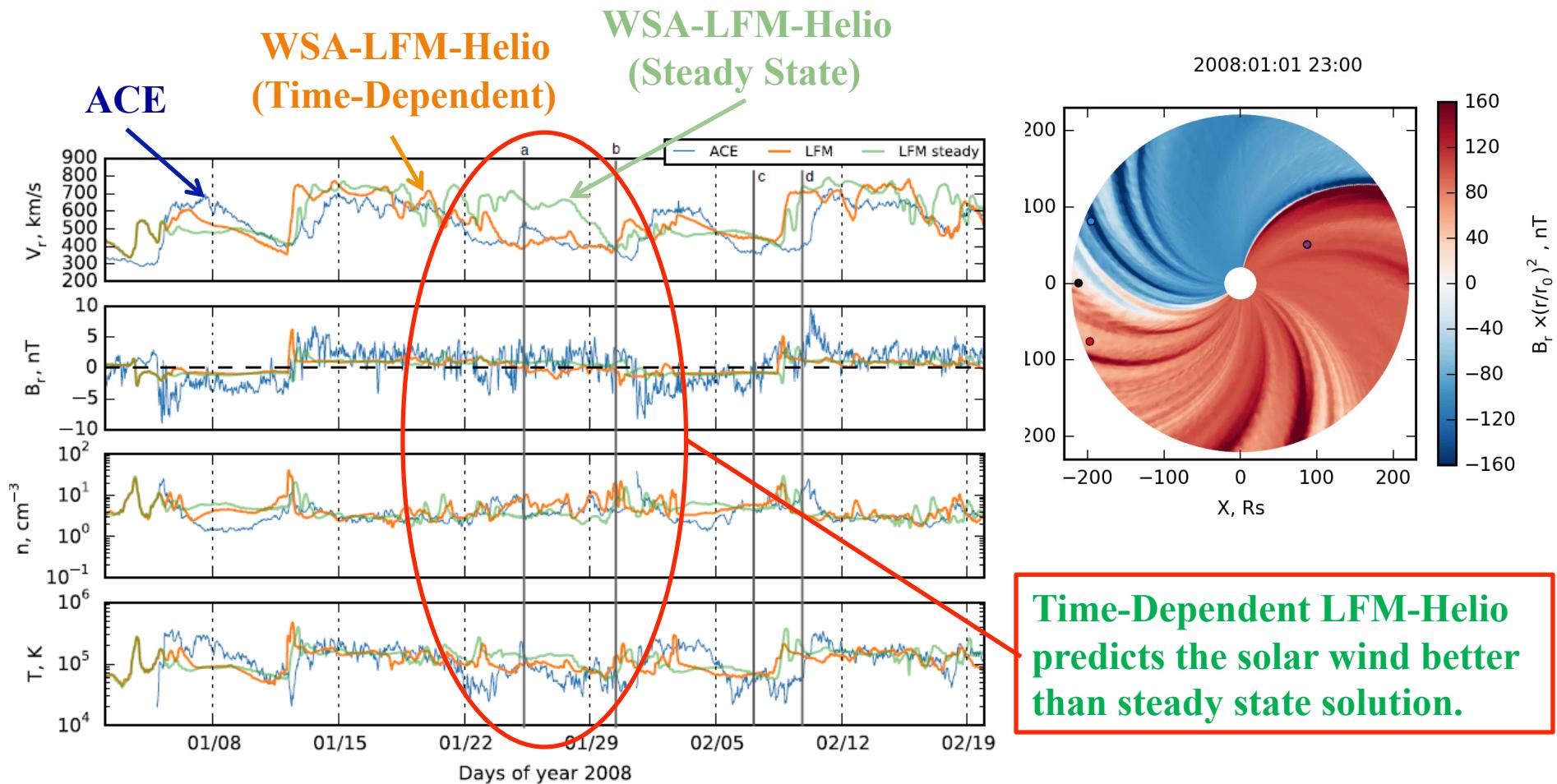
Without Far-Side Active Region Inserted

4 Day Advanced Predictions and STEREO B Observations





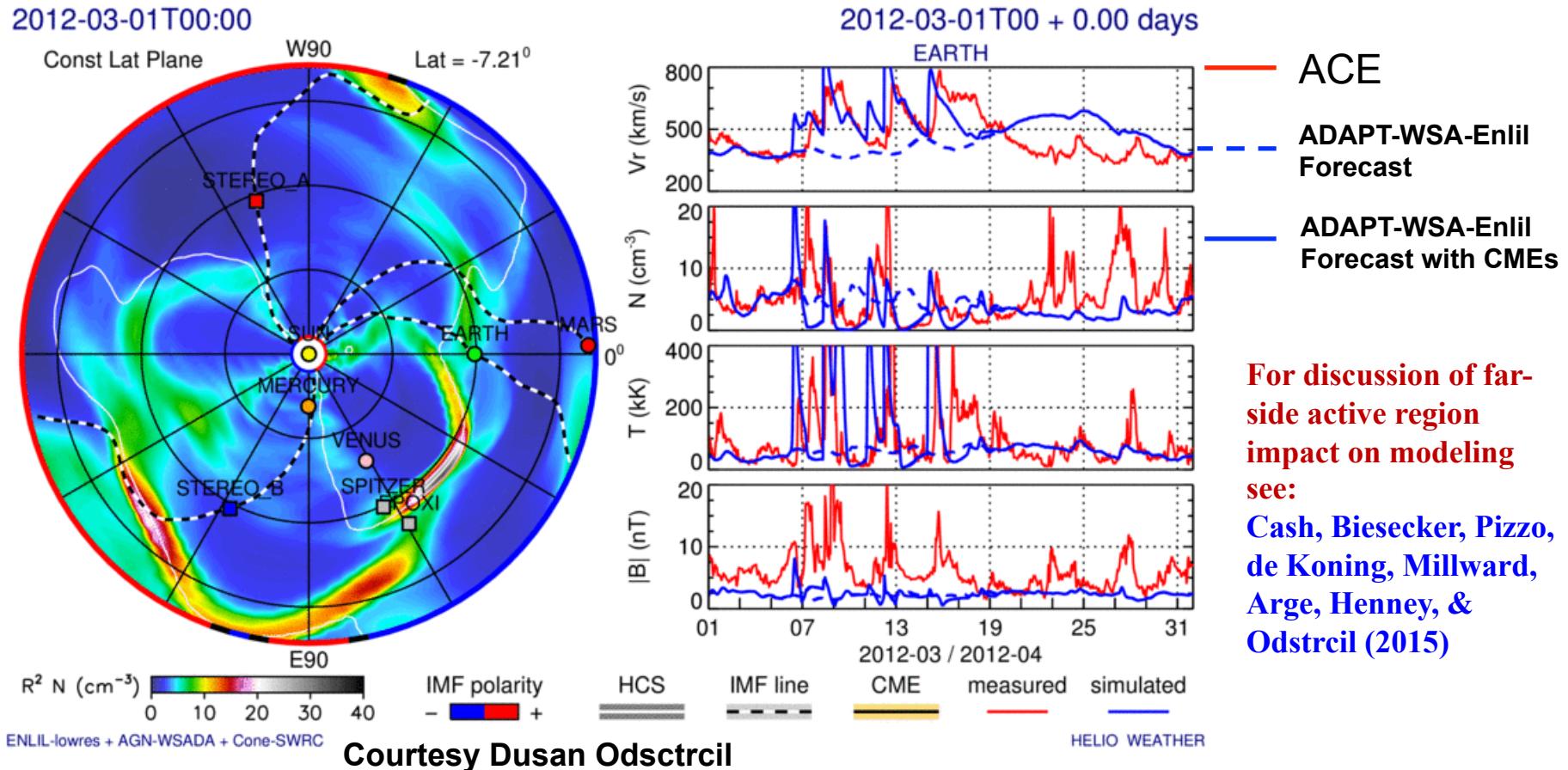
ADAPT Driven WSA-LFM+Helio



Merkin, Lyon, Lario, Arge, & Henney, JGR, 2016 (Accepted)



“Time-Dependent” ADAPT-WSA-Enlil



Moving toward time-dependent ADAPT-WSA-Enlil solar wind forecast capability



Issues Facing Coronal & Solar Modeling



- Solar community moving toward *true* time-dependent modeling of corona & solar wind
- The global solar photospheric magnetic field distribution serves as primary input to *nearly all* coronal and solar wind models.
- Several *issues* with photospheric magnetic field input driver maps
 - Poles not observed for months at a time
 - Coronal models very sensitive to the polar fields
 - Can Introduce artificial monopoles into maps
 - Traditional magnetograms of the solar far-side not available
 - Helioseismology helping
 - Solar Orbiter will provide some far-side coverage
 - Far-side flux emergence can
 - Change *global* coronal solution
 - Introduce artifacts into the models (e.g., “sudden” appearance of active regions as they rotate on to the solar near-side)
 - Introduce artificial monopoles into maps
- Flux transport models (e.g., ADAPT) ideally suited to drive solar models time-dependently
 - Required because routine global observations of Sun’s magnetic field not available
 - Provide time sequences of the global photospheric magnetic field at cadences required by models.

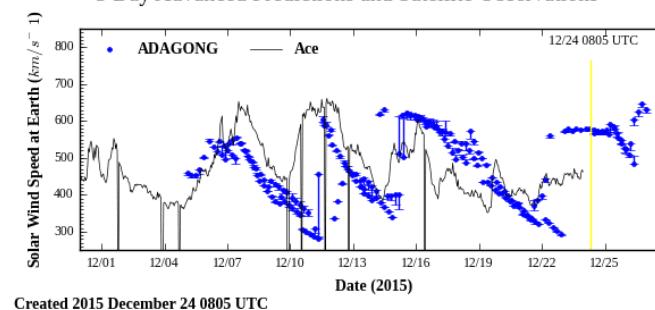
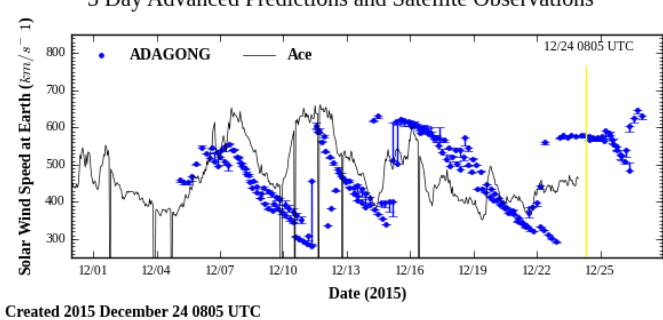
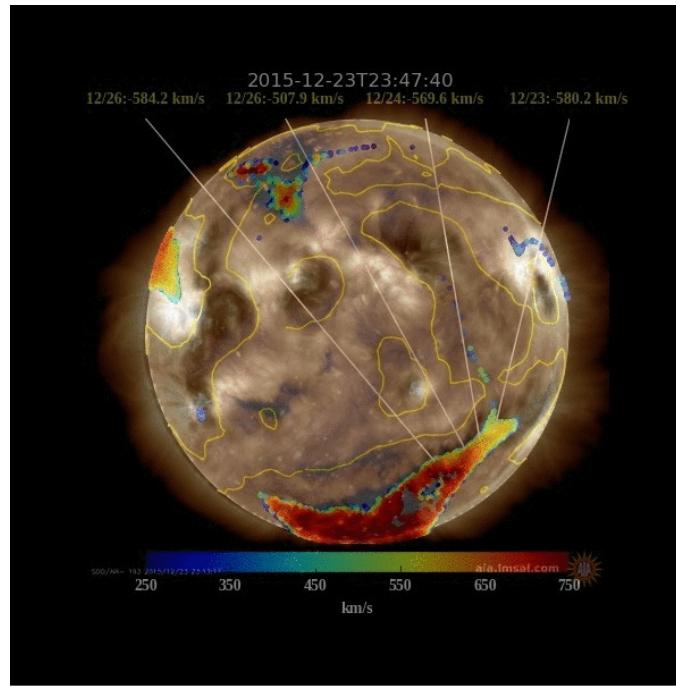
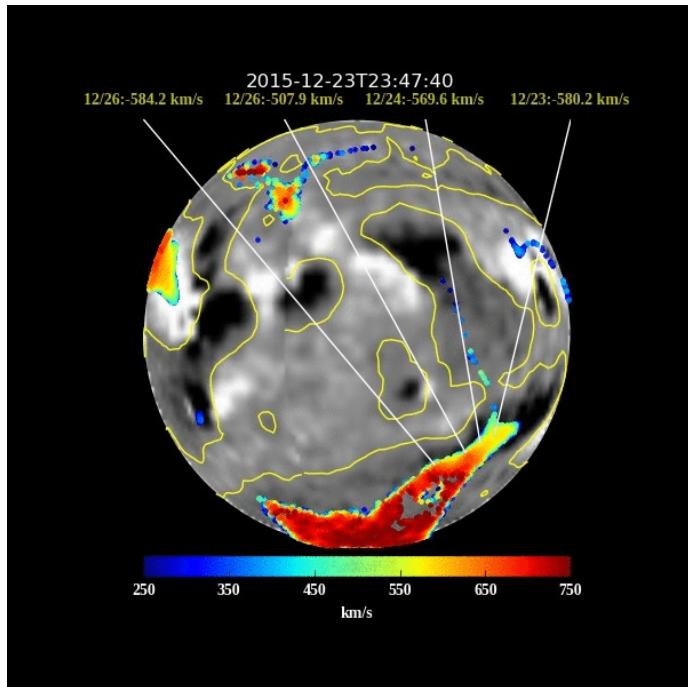


Summary

- Wang-Sheeley-Arge (WSA) model - combined empirical and physics based model of the corona and solar wind.
 - Improved version of the original Wang and Sheeley model.
 - Operational at NOAA/NCEP and NASA/CCMC.
- Key Features:
 1. Field line tracing parallelized
 2. Compatible with most input photospheric field maps including ADAPT.
 3. Can be run in PFSS or coupled PFSS+Schatten Current Sheet (SCS) modes.
 4. Optional improved interface between PFSS & SCS models can be employed
 - minimizes “kinking” at interface.
 5. IDL & Perl scripts replaced with Python
 6. Can forecast solar wind speed and IMF polarity at inner planets & most satellites
 - Easy now to add other satellites.
 7. In the process of retuning empirical solar wind velocity relationship.
 8. New field line tracing package.
 9. Available soon!
- Coronal & solar wind modelers face new challenges
 1. when moving to time-dependent modeling.
 2. contending with high cadence, high resolution, multi-realization input maps.

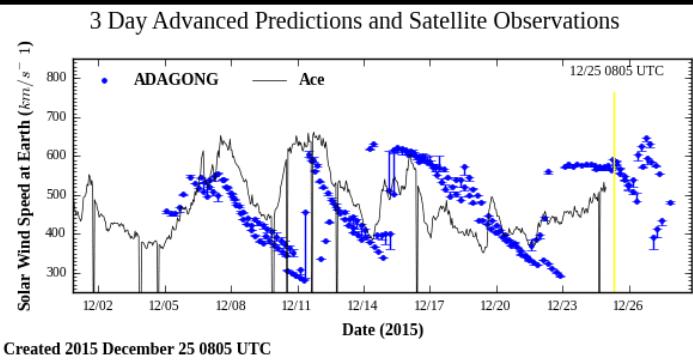
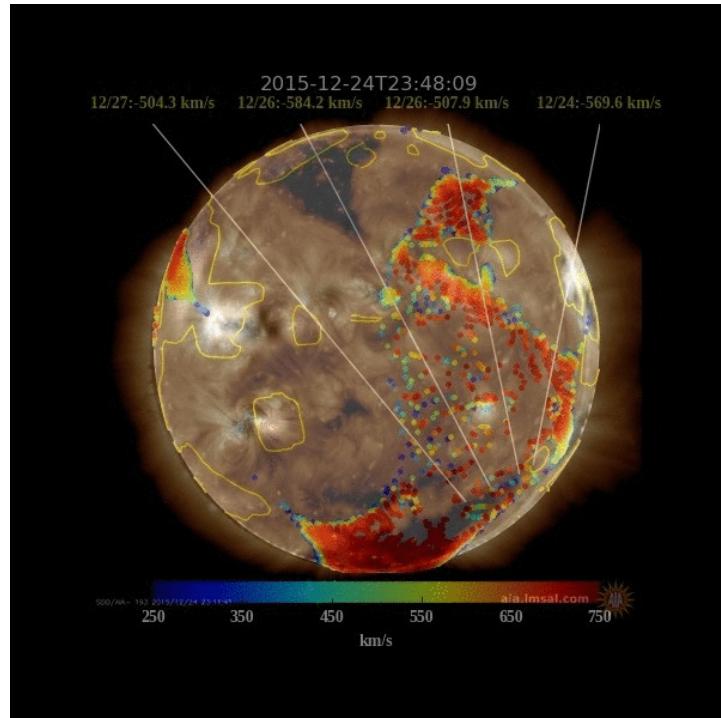


ADAPT-Driven WSA Coronal & Solar Wind Predictions

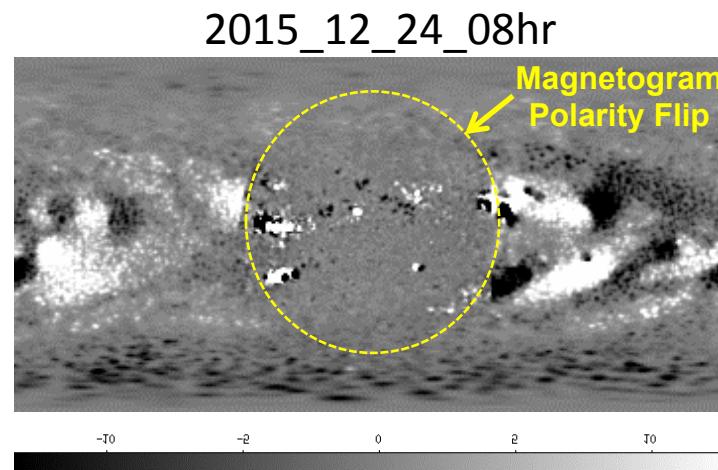
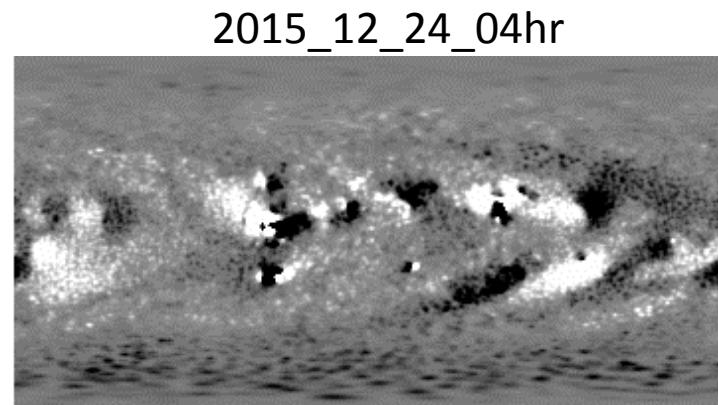




ADAPT-Driven WSA Solar Wind Predictions (What can happen when things go wrong!)

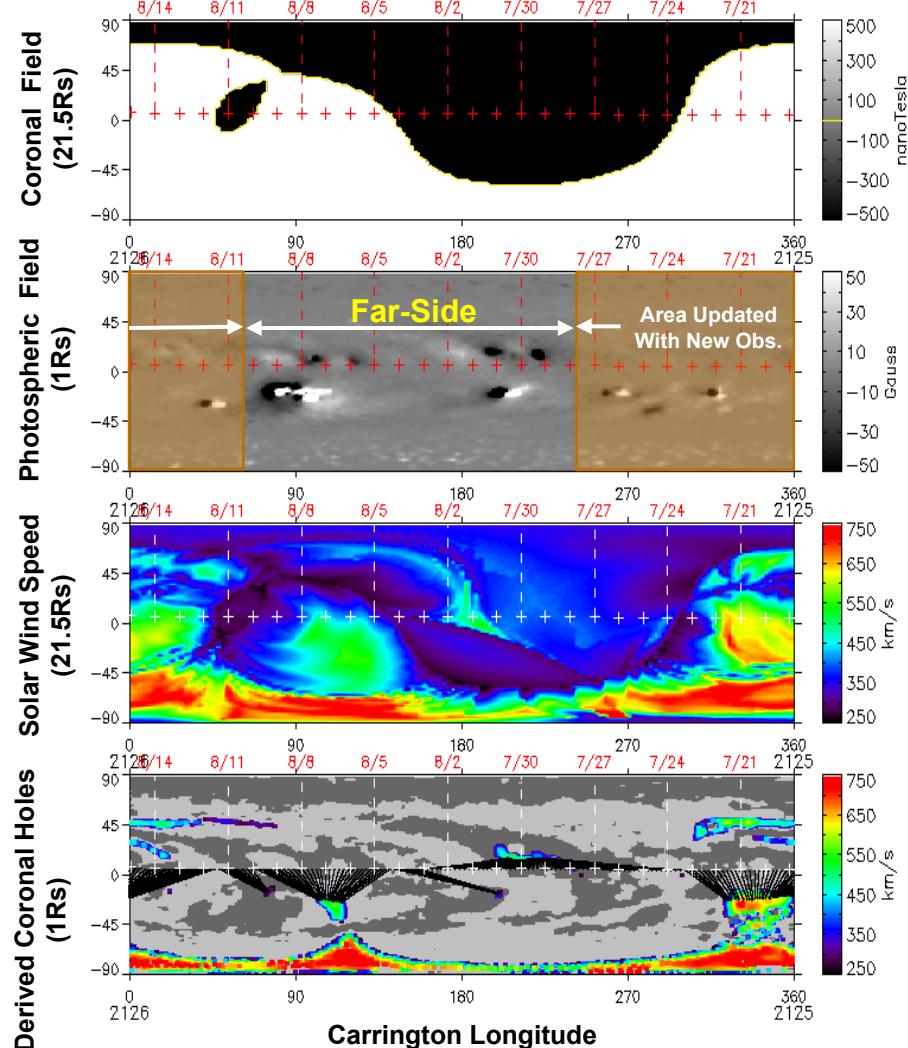


ADAPT-GONG Map

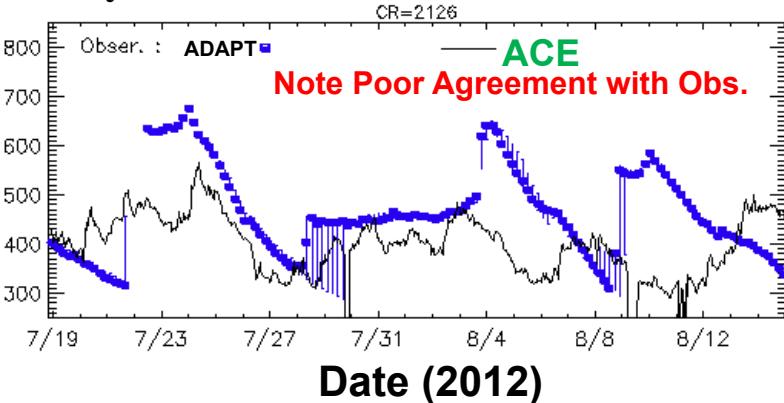




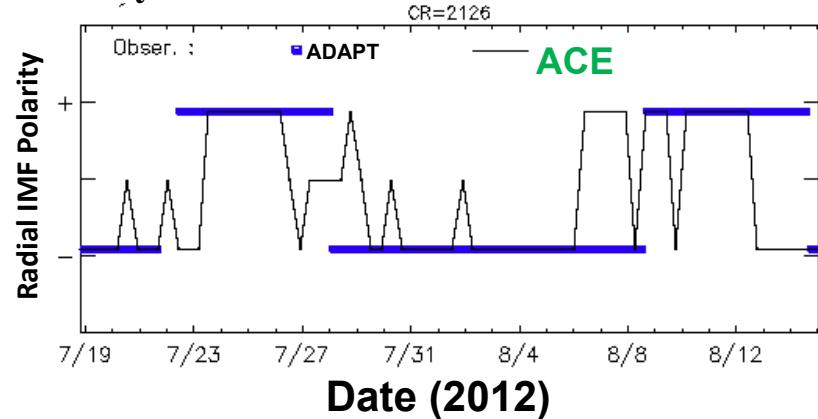
WSA Coronal and Solar Wind Predictions Using *July 21, 2012* ADAPT Map as Input to WSA



4 Day Advanced Predictions vs Observations



4 Day Advanced Predictions vs Observations





WSA Coronal and Solar Wind Predictions Using *July 25, 2012* ADAPT Map as Input to WSA

