

NAIRAS model and application to space weather effects for suborbital flights

Christopher J. Mertens

**NASA Langley Research Center
Hampton, VA**

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Suborbital flights: space weather impacts and modeling

Annapolis, Maryland

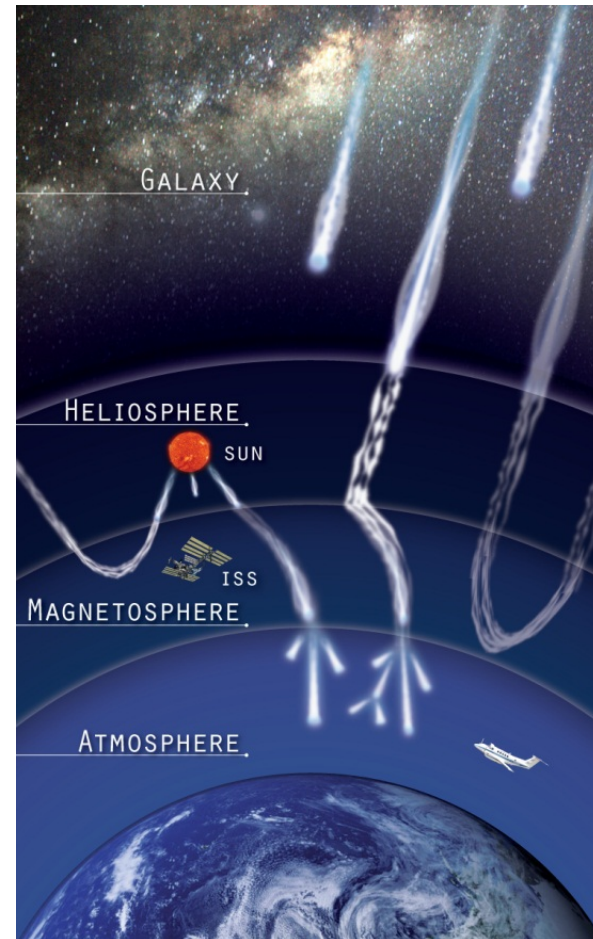
April 03, 2014

Outline

- **NAIRAS Model Overview**
 - Publically Available Products
 - Profile Information
- **NAIRAS Model Components**
 - GCR Model
 - SEP Spectral Fitting
 - Geomagnetic Shielding
 - Vertical Cutoffs
 - Directional Effects
 - Inner Proton Belt Model
- **RaD-X Balloon Mission**
- **Wrap-Up**

NAIRAS Model

- NASA LaRC's Nowcast of Atmospheric Ionizing Radiation for Aviation Safety (NAIRAS) Model:
 - Real-time physics-based, global model
 - Real-time inclusion of GCR and SEP radiation
 - Real-time solar-magnetospheric effects on radiation
 - Real-time meteorological data used



Public Web site: <http://sol.spacenvironment.net/~nairas/> (or google NAIRAS)

NAIRAS Model Predictions During March 2012 Solar Storm Events

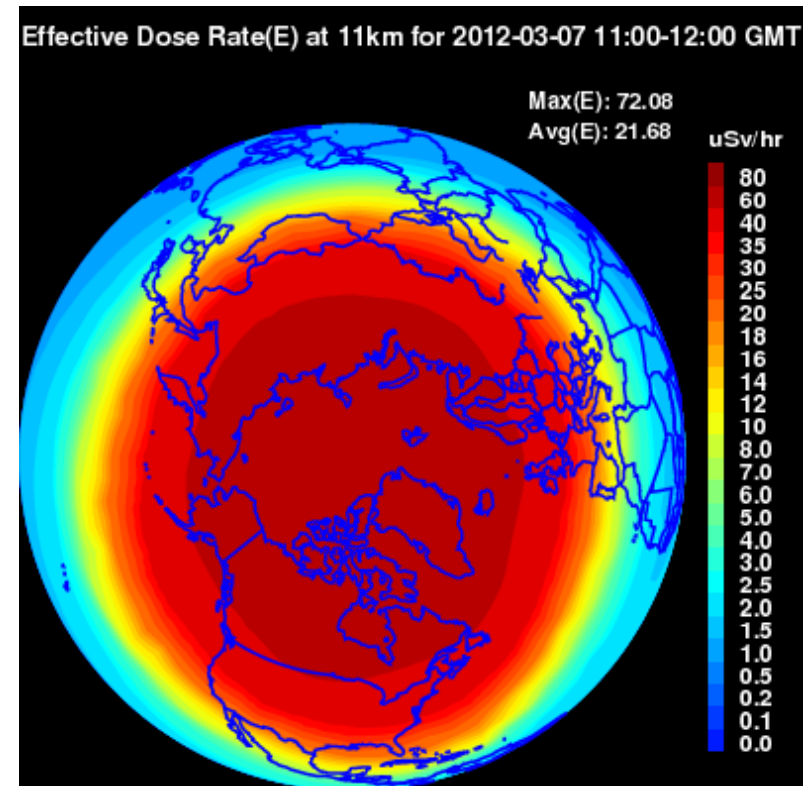
Effective Dose Rate¹(E) for 2012-03-07 11:00-12:00 GMT

5 km (16,000 feet) Radiative Dose Rate (uSv/hr)								
lat	90S-60S	60S-40S	40S-20S	20S-0	0-20N	20N-40N	40N-60N	60-90N
avg	12.01	8.51	3.77	1.06	0.88	2.58	8.21	11.74
max	13.67	12.35	8.81	3.03	3.64	8.71	12.45	13.61
11 km (35,000 feet) Radiative Dose Rate (uSv/hr)								
lat	90S-60S	60S-40S	40S-20S	20S-0	0-20N	20N-40N	40N-60N	60-90N
avg	63.79	43.36	15.16	2.69	2.14	9.45	41.37	66.03
max	69.41	66.18	47.22	10.15	13.08	45.58	67.98	72.08
15 km (49,000 feet) Radiative Dose Rate (uSv/hr)								
lat	90S-60S	60S-40S	40S-20S	20S-0	0-20N	20N-40N	40N-60N	60-90N
avg	133.91	84.70	22.91	3.21	2.51	13.33	76.84	142.45
max	144.86	140.66	94.98	13.45	17.59	90.36	147.34	152.19

Representative High-Latitude Flights

2012-03-07 11:00-12:00 GMT						
Flight Name	Time	Rate ¹	Dose ¹	Safety Signal		
	hours	uSv/hr	mSv	Aircrew ²	Public ³	Prenatal ⁴
London,GBR - New York,USA	5.50	51.52	0.283	Yellow	Green	Yellow
Chicago,USA - Stockholm,SWE	8.50	63.09	0.536	Yellow	Yellow	Red
Chicago,USA - Munich,DEU	8.50	56.35	0.479	Yellow	Yellow	Red
Chicago,USA - Beijing,CHN	13.50	60.34	0.815	Yellow	Red	Red

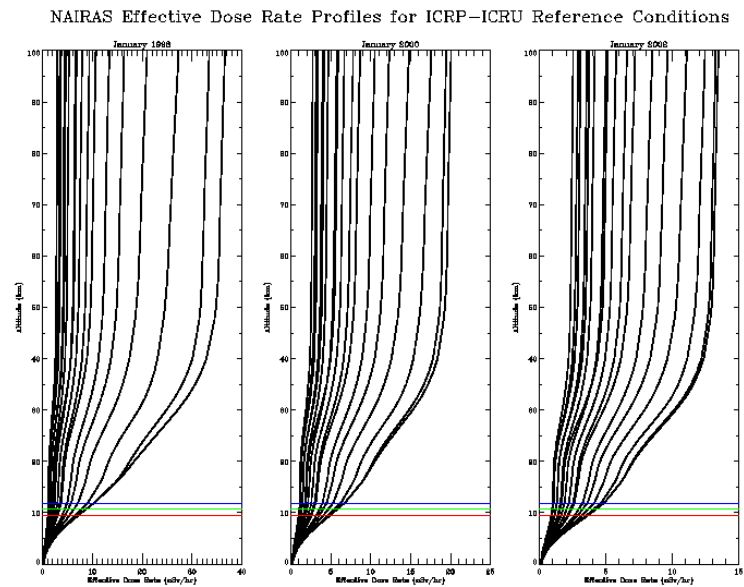
Signal	Aircrew ⁵ Max_Annual(800hrs)	Public ⁶ one_trip	Prenatal ⁶ one_trip
Green	0-6.0mSv	0-0.330mSv	0-0.167mSv
Yellow	6.0-12.0mSv	0.330-0.670mSv	0.167-0.333mSv
Red	>12.0mSv	>0.670mSv	>0.333mSv



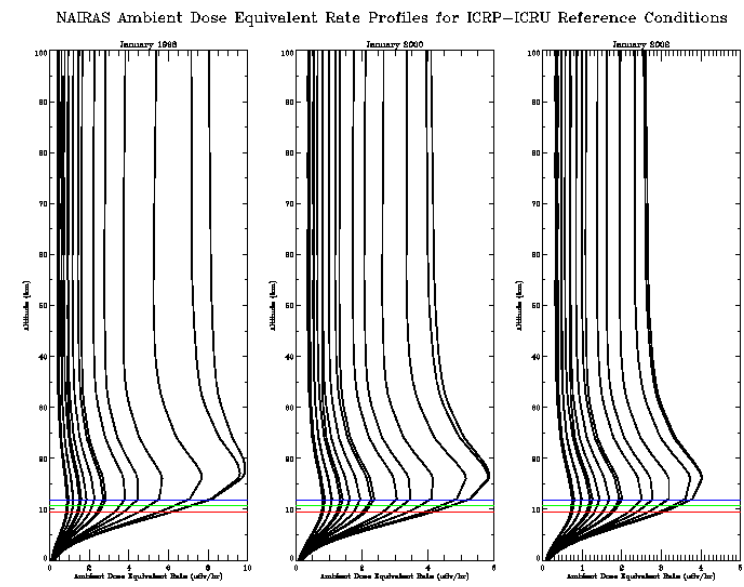
Public Web site: <http://sol.spacenvironment.net/~nairas/> (or google NAIRAS)

NAIRAS GCR Dose Profiles

(Mertens et al., 2013)



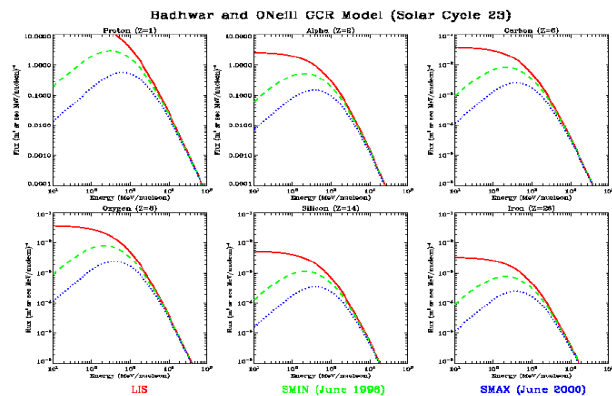
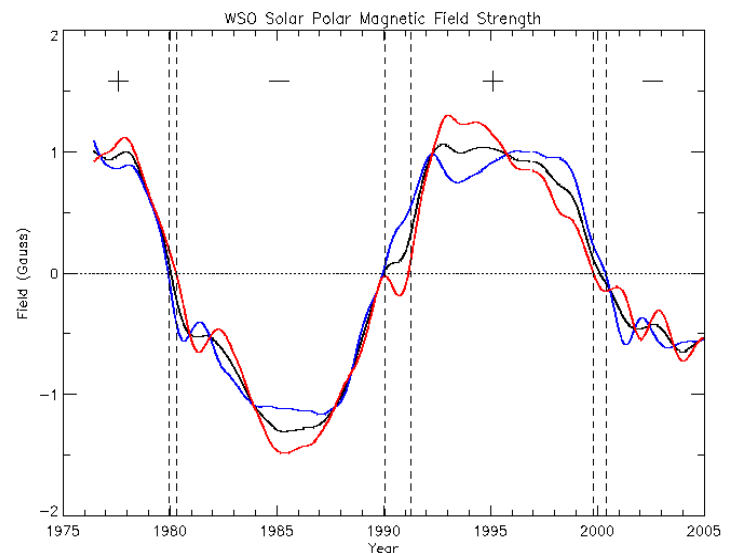
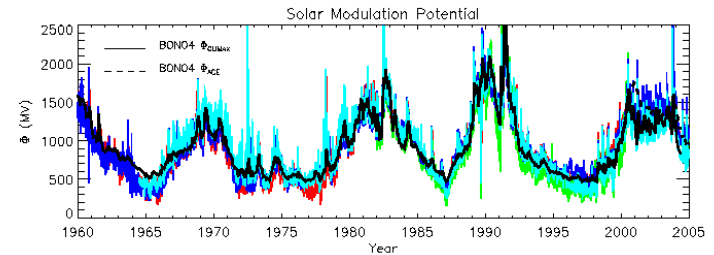
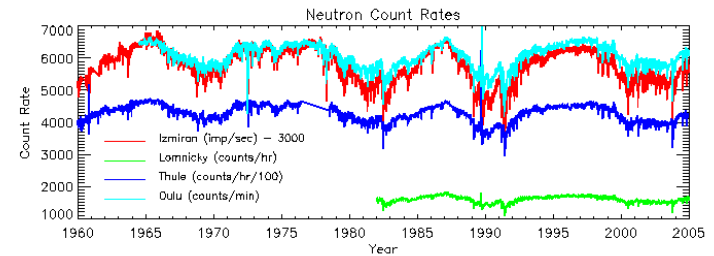
Effective Dose Rate



Ambient Dose Equivalent Rate

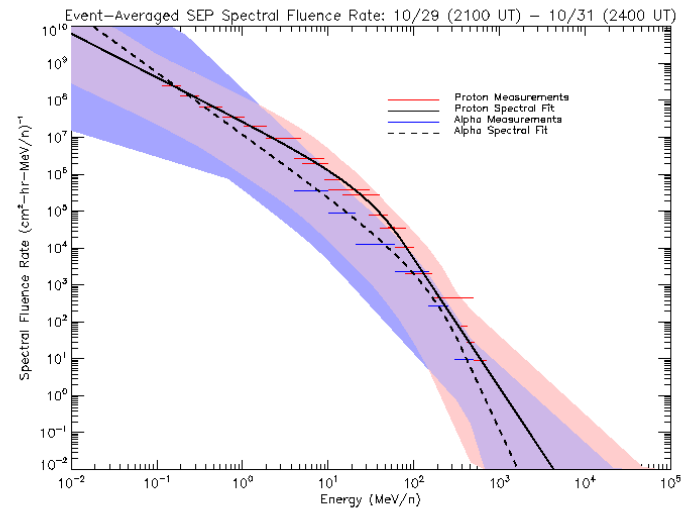
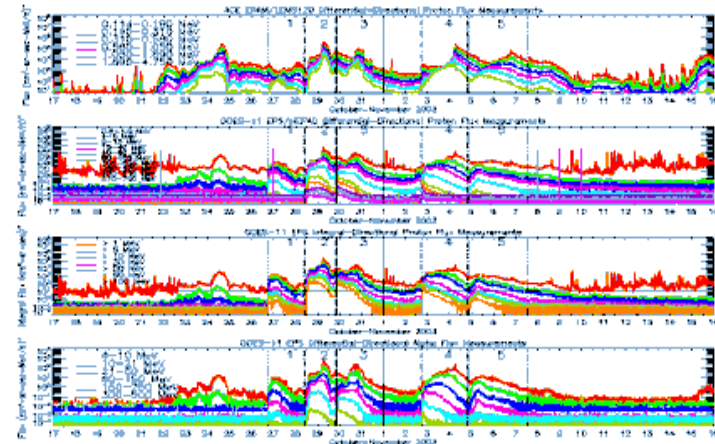
GCR Fluence Spectra

- Ground-based neutron monitor count rates related to incident GCR flux
- Use high-latitude, real-time neutron monitor count rates to derive heliospheric potential (HP)
- Fit real-time neutron count rates to Badhwar & O'Neill 2010 GCR HP
 - Steady-state Fokker-Planck transport to 1 AU
 - HP embedded in diffusion coefficient
 - Validated by NASA/ACE

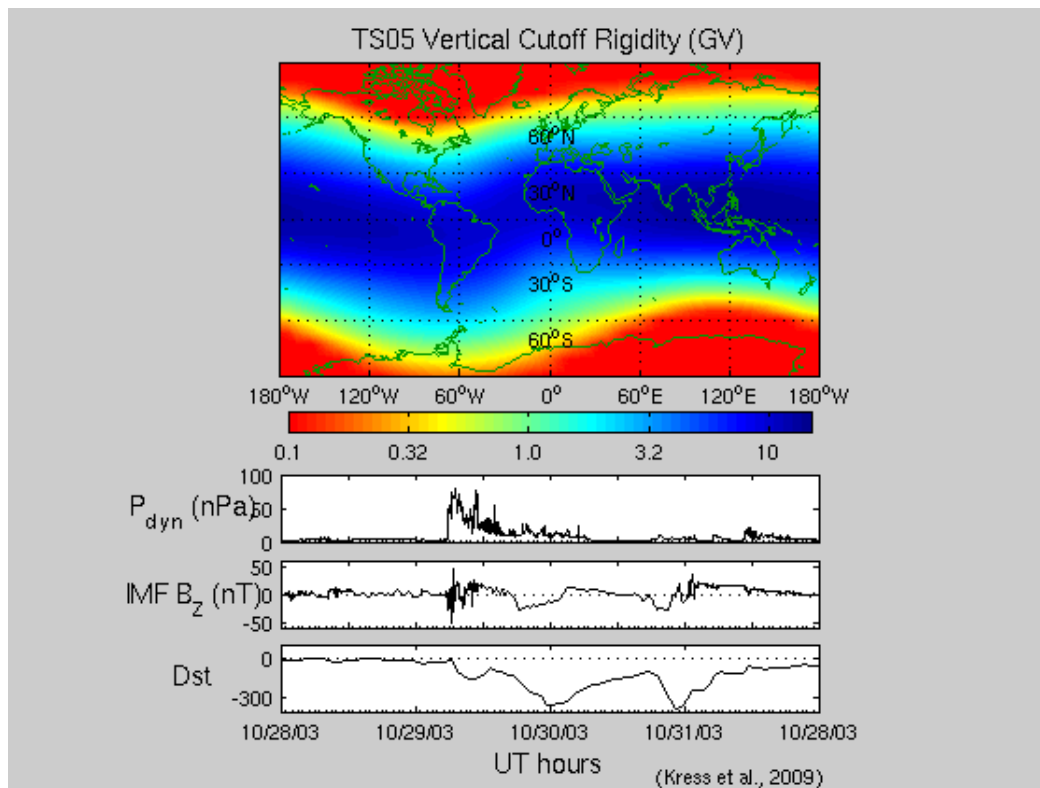


SEP Fluence Spectra

- GOES and ACE observes proton/alpha fluxes which we need to derive the incident fluence rates and spectral characteristics
- Fit measurements to analytical spectral fluence distributions
 - Require differential spectral flux measurements to decrease monotonically with increasing energy
 - Fit a single power-law, Ellison-Ramaty, double power-law, and Weibull distributions
 - Select spectral distribution with the minimum χ^2

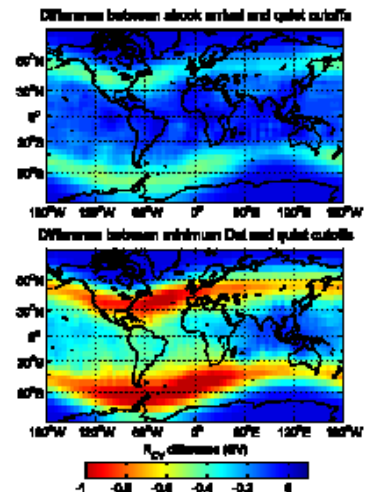


Geomagnetic Shielding

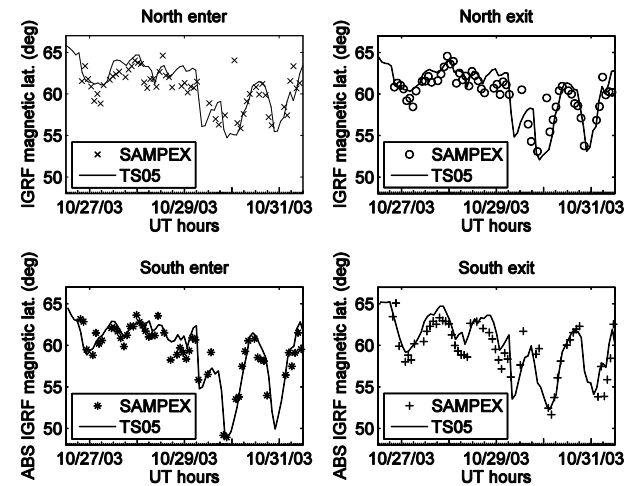


- Severe geomagnetic storms suppress geomagnetic shielding allowing SEPs access to mid-latitudes.
 - Due primarily to a build up of the ring current
 - Shock arrival can also be significant
- Particles with rigidities below the a cutoff value cannot access that point in space
 - We compute these using the TS05 storm magnetic field model and a particle tracing code
 - During the Halloween storms we find 1 and 0.5 GV suppression during main phase and shock arrival, respectively

Halloween 2003 Geomagnetic Storm



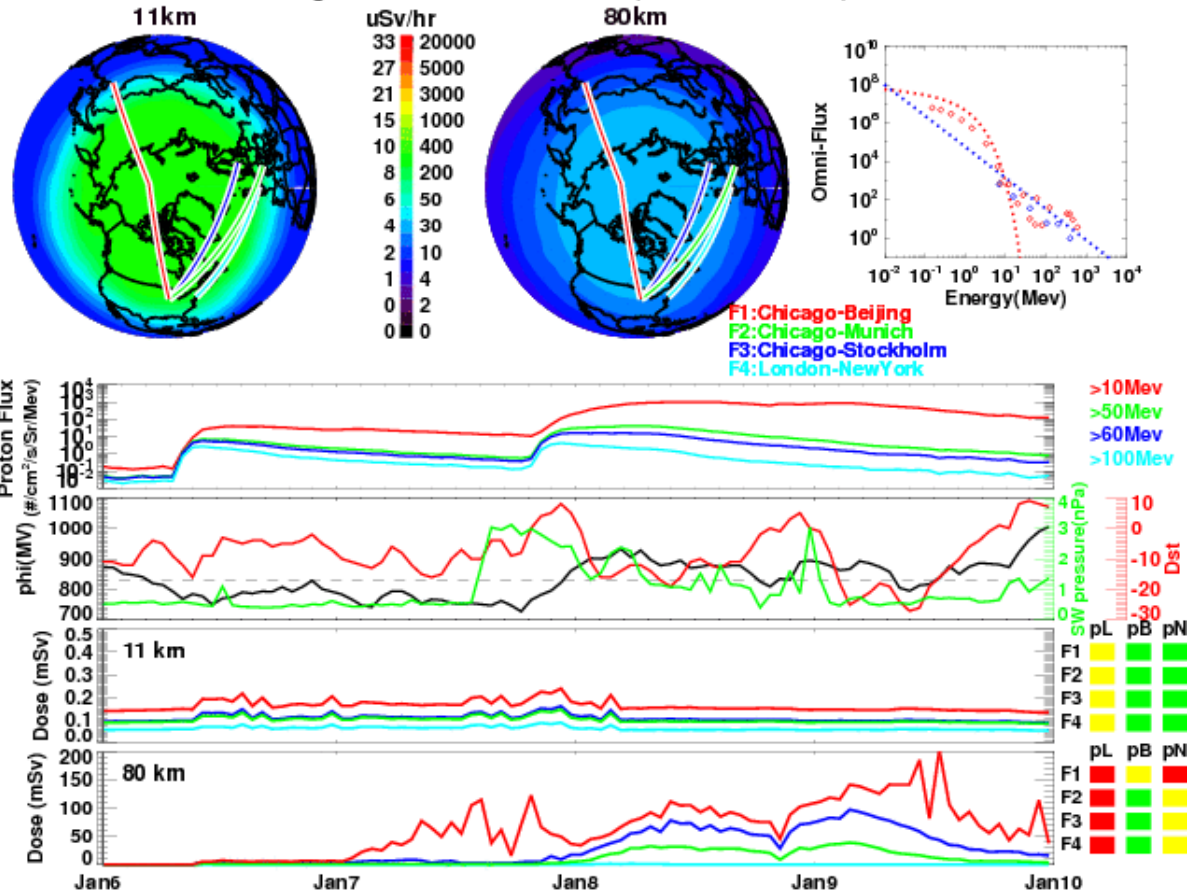
Cutoff Latitude of ~ 20 MeV protons



- TS05 cutoff latitude accuracy: ~ 2 -degrees
- TS05 cutoff accuracy significantly better LFM/MHD code
- 2-degree cutoff uncertainty in latitude can translate into factor ~ 2 or greater aircraft dose uncertainty

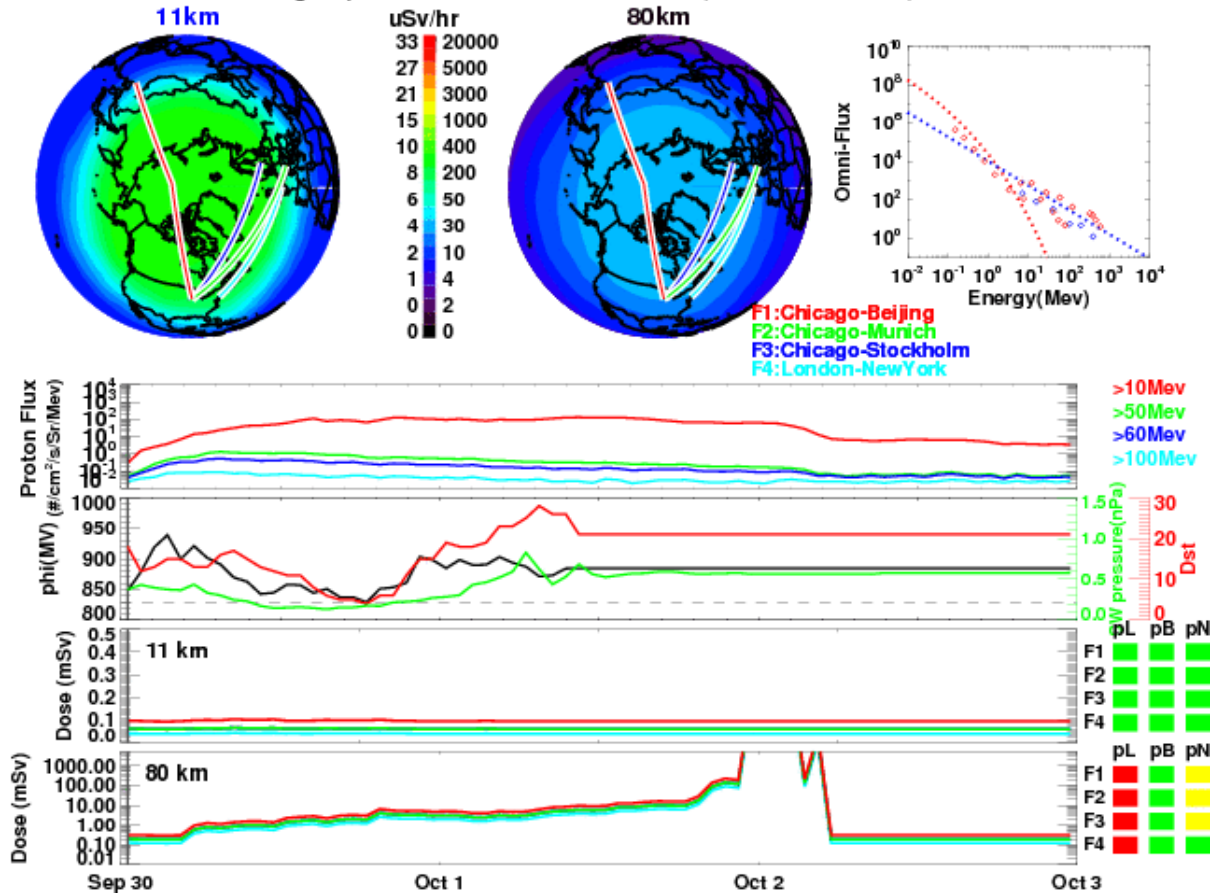
January 2014 Solar Storm Period

Effective Dose during Jan6-9, 2014 Solar Storm (01-06-2014, 00UT)

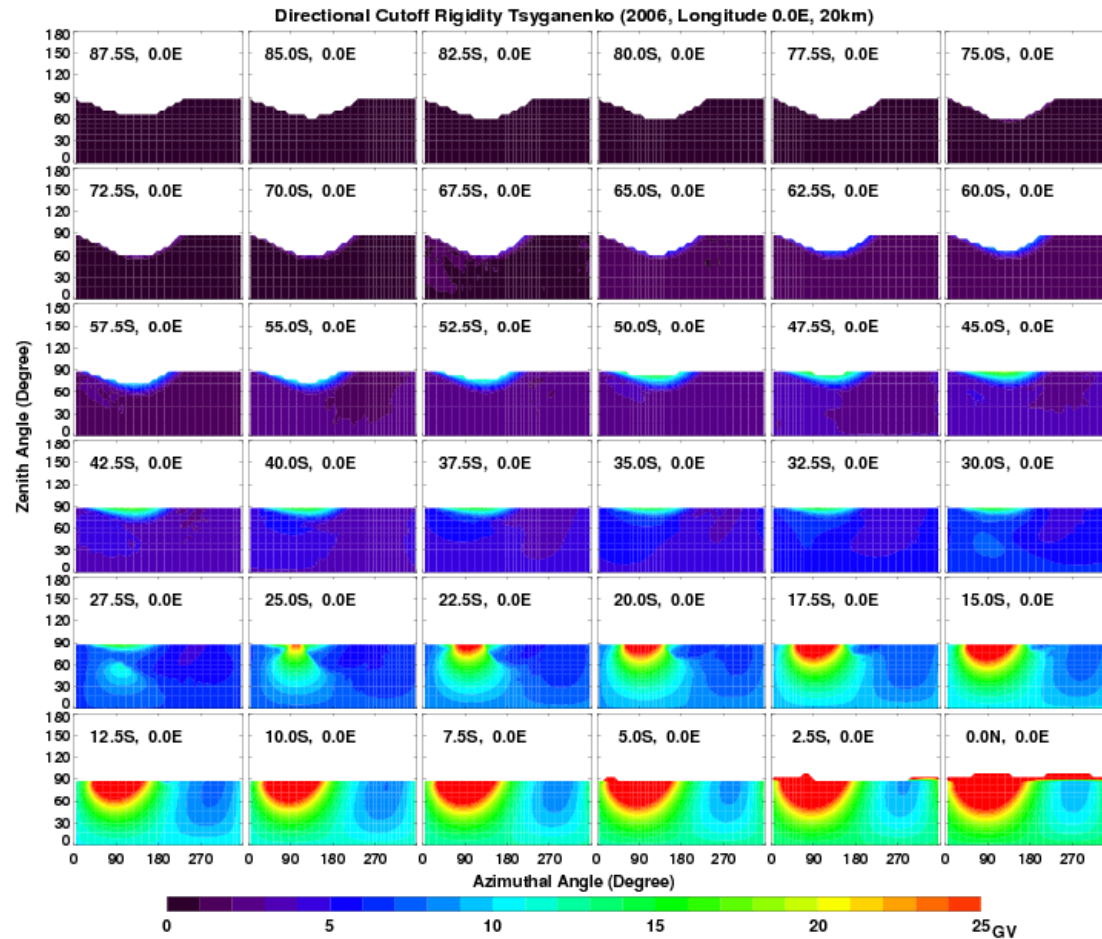


September-October 2013 Solar Storm Period

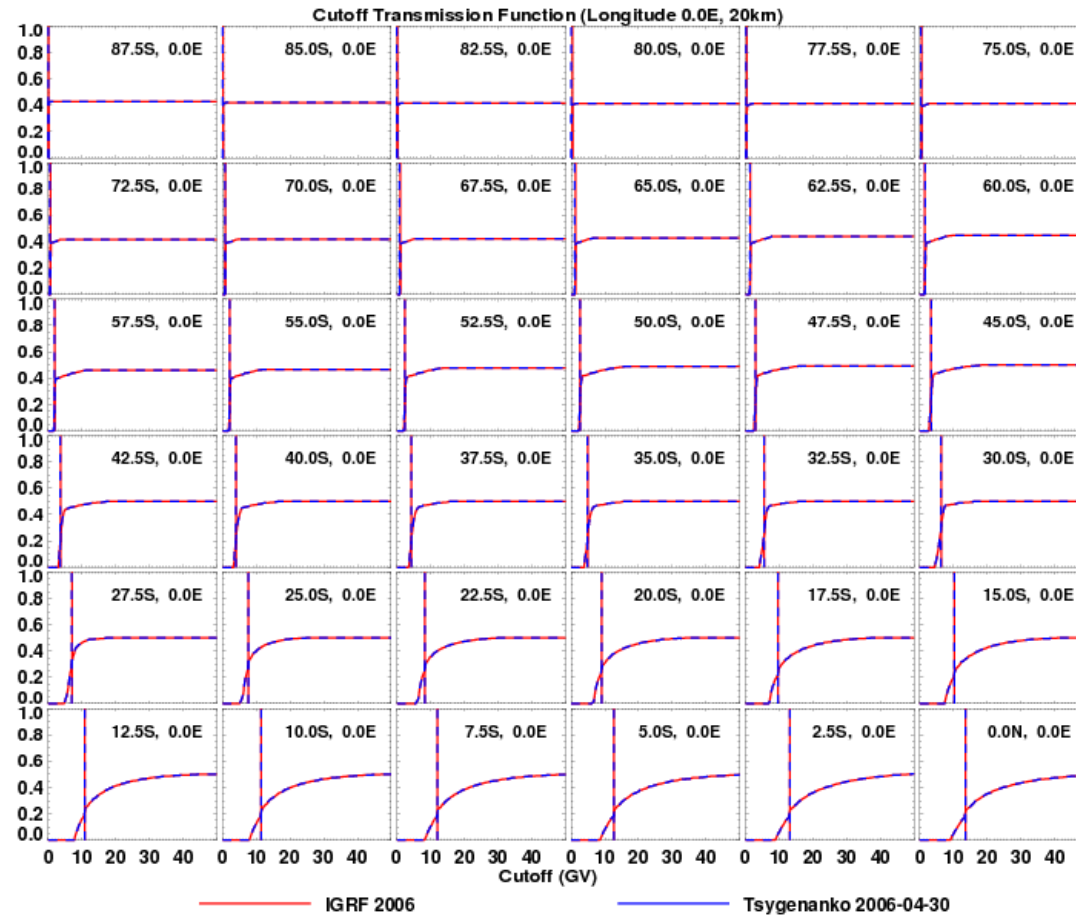
Effective Dose During Sep 31 - Oct 2, 2013 Solar Storm (09-30-2013, 00UT)



TS05 Directional Geomagnetic Cutoffs Quiet-Time

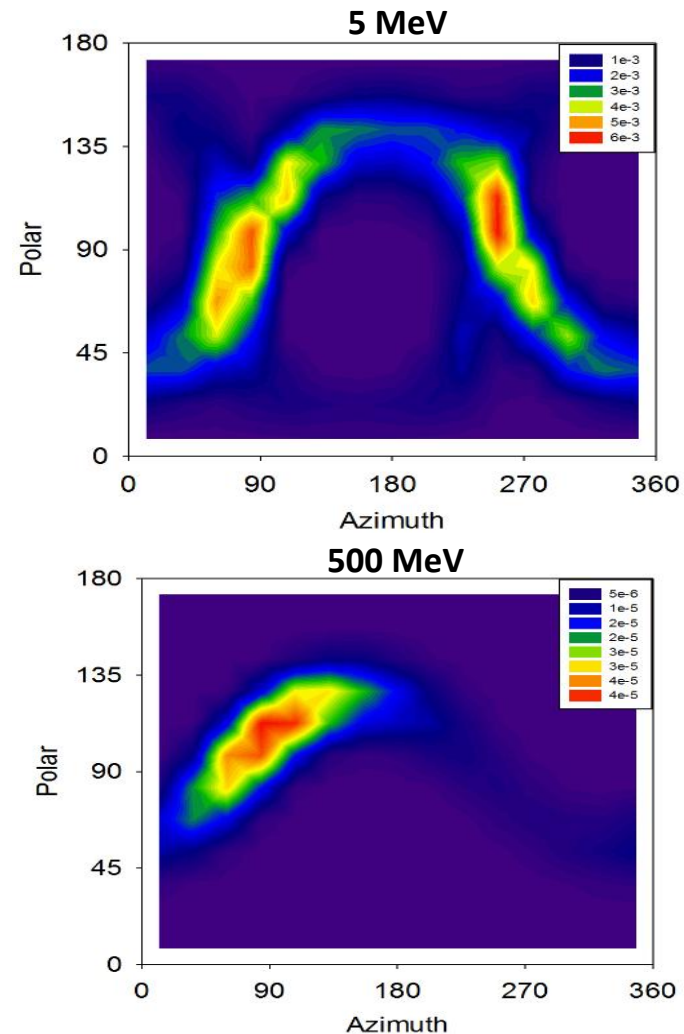


Quiet-Time Geomagnetic Transmission TS05 & IGRF



Inner Proton Radiation Belt

- Under quiet conditions, inner zone protons with energies > 100 MeV represent most the important contributions to radiation exposure in LEO
- AP-8 model extended to include temporal, spatial and directional effects
 - SAA north-west temporal/spatial drift
 - Solar cycle variation using F10.7 and NM data
 - Trapped protons sample large altitude region in LEO, which introduces prominent east-west asymmetry in atmospheric loss rate, requiring a vector trapped proton model
- Same extensions to AP-8 will be applied to AP-9



Directional dependency of the trapped protons as viewed in the center of the SAA at 400 km. Contour units in $(\text{MeV}\cdot\text{cm}^2\cdot\text{sr}\cdot\text{s})^{-1}$



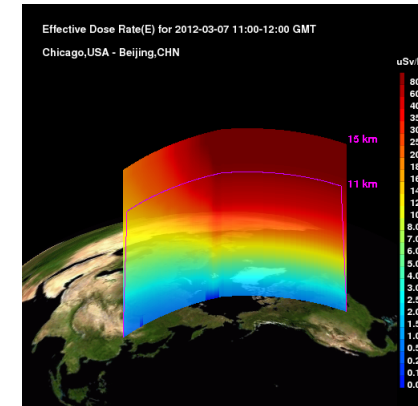
RaD-X : Radiation Dosimetry Experiment



Science Goals and Objectives

- Improve predictive atmos. ionizing dose tools by characterizing energy deposition by cosmic ray primaries
 - Measure dosimetric quantities in upper atmos
 - Separate cosmic ray primary contributions
- Identify and characterize low-cost atmos. radiation dose measurement solutions for continuous, global monitoring

RaD-X Measures Radiobiological Dose and CR Primary Proton and HZE Contributions



Mission and Instrument Parameters

- Platform: High-Altitude Balloon
- Launch Site: Fort Sumner, NM
- Mission Duration: 24-hours
- Temporal Sampling: 1-5 minutes
- Launch Readiness Date: September 2015
- Instruments: (1) TEPC, (2) TID detector, (3) LET spectrometer, and (4) microdosimeter emulator
- Measurement uncertainty: < 30 %
- All instrument components at TRL 6 or higher

Science Team and Partners

NASA Langley
NASA ARC
NASA GSFC/WFF
Prairie View A & M University (PVAMU)
Center for Radiation Engineering
and Science for Space Exploration (CRESSE)
University of Virginia
Space Environment Technologies, Inc.

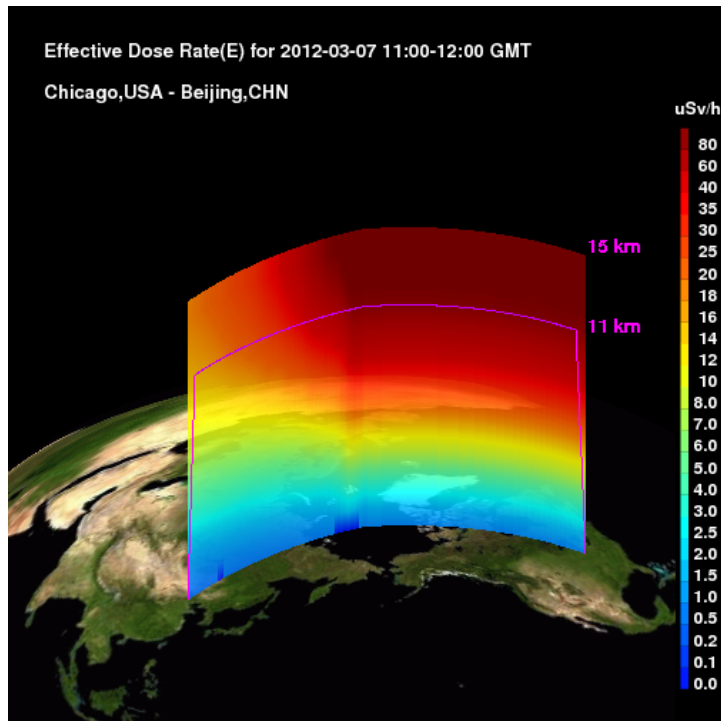
Wrap-Up

- **NAIRAS Computed Output Products**
 - **Dosimetric Quantities**
 - Effective dose, ambient dose equivalent, silicon absorbed dose, tissue absorbed dose, and tissue dose equivalent
 - **Flux Quantities**
 - Spectral flux (total and by particle), total flux (by particle and particle sum), atmospheric ionization rate, >10 MeV neutron flux (microelectronic effects)
 - **Differential flux and differential absorbed dose LET Spectra (tissue and silicon)**
- **NAIRAS Studies/Updates for LEO**
 - Continue to assess directional cutoff rigidity effects
 - Integrate inner proton belt radiation model

Backup Slides

NAIRAS Graphical Products

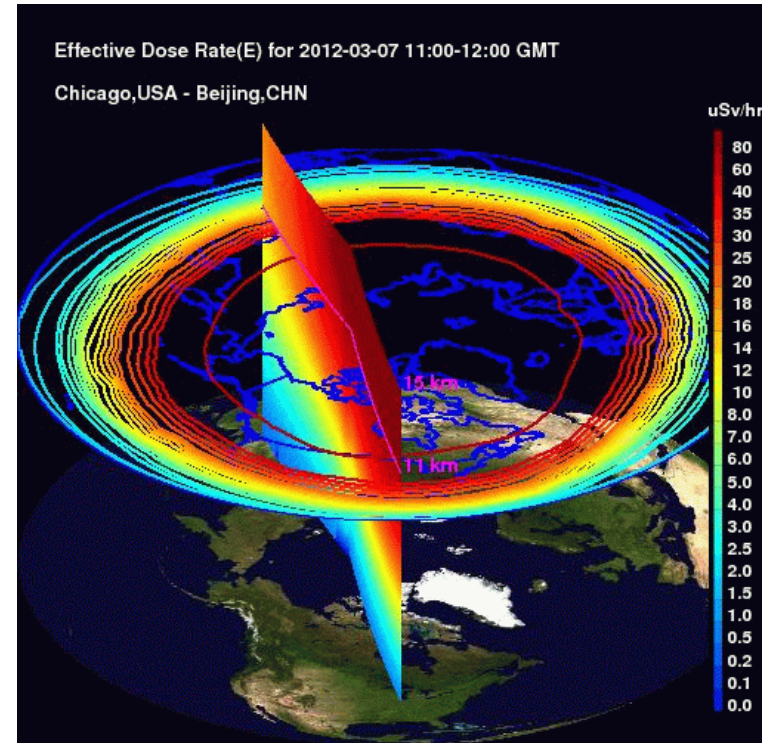
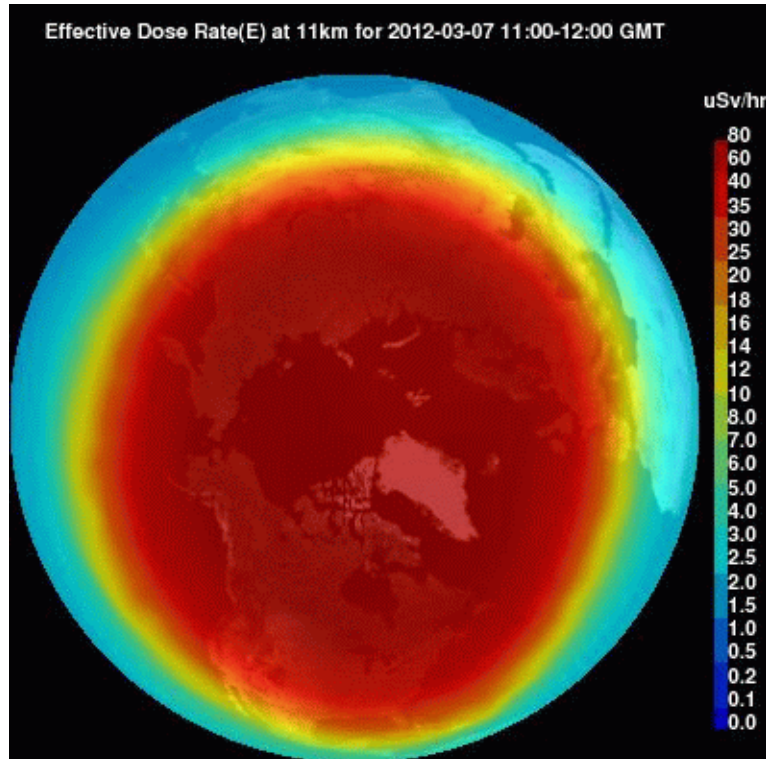
March 2012 Solar Storm Event



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NAIRAS Graphical Products

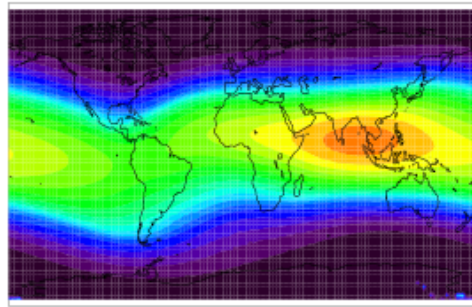
March 2012 Solar Storm Event



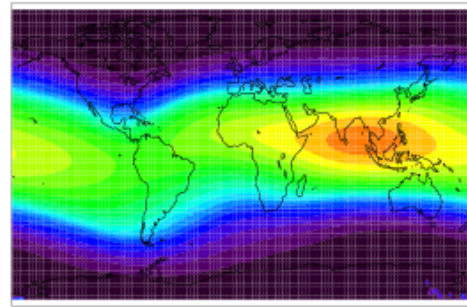
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Vertical Cutoff Rigidity (2006) TS05 & IGRF

Vertical Cutoff (IGRF, 2006)

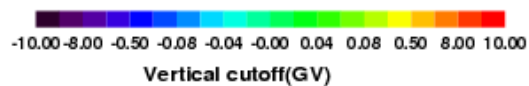
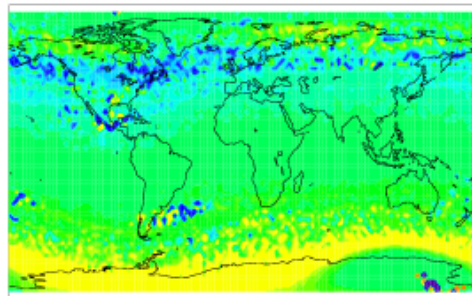


Vertical Cutoff (Tsyganenko, Apr 30, 2006, 2:30UT)

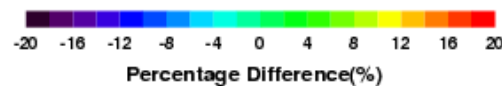
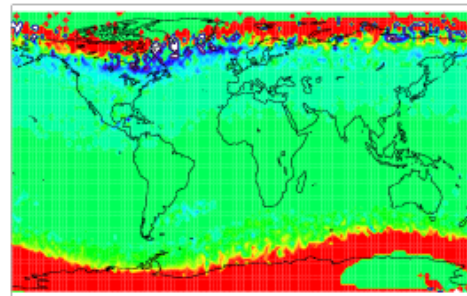


Difference between Vertical IGRF(2006) and Tsyganenkco Cutoff (Apr 30, 2006, 2:30UT)

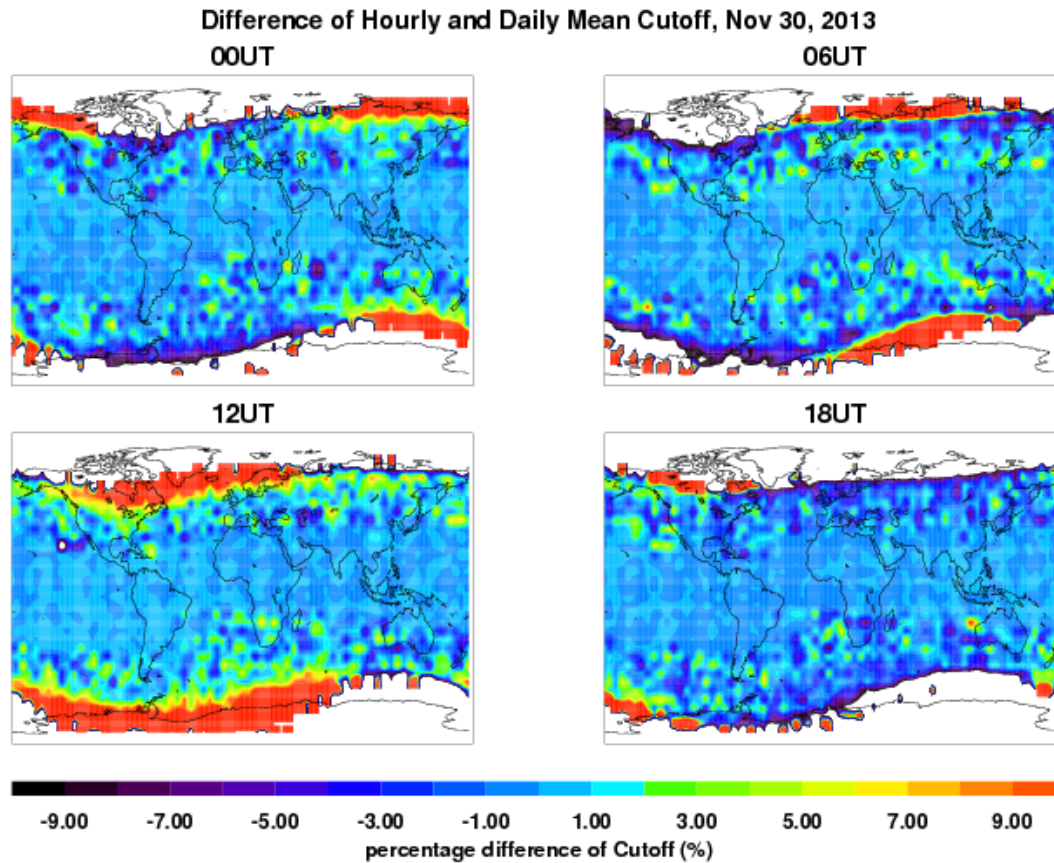
IGRF-Tsyganenko



(IGRF-Tsyganenko)/IGRF *100%

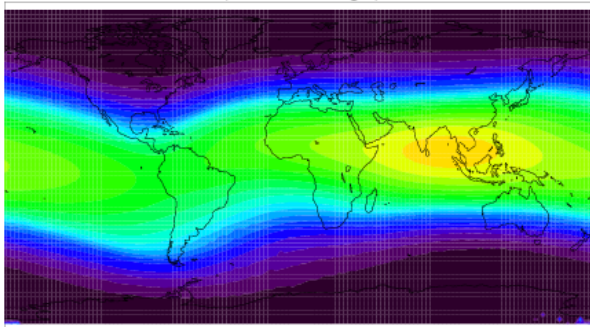


Daily Variation in TS05 Vertical Cutoffs Quiet-Time

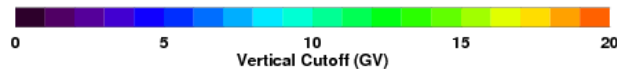
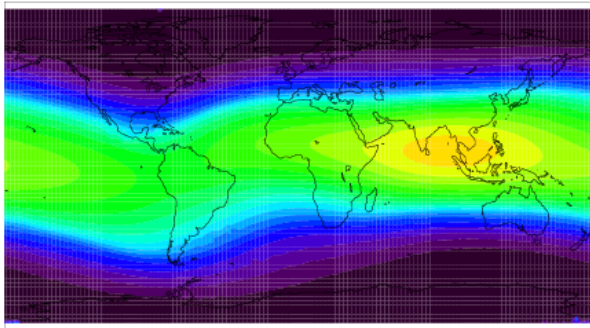


Diurnal Variation in TS05 Vertical Cutoffs Quiet-Time

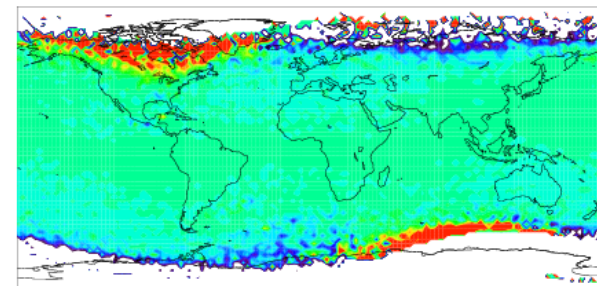
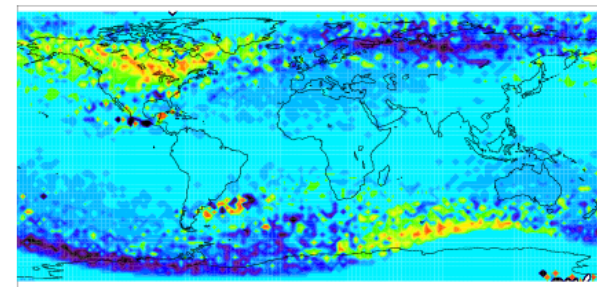
Vertical Tsyganenko Cutoff Rigidity on Mar 22, 2006 at 20km
(6UT, CST Midnight)



(18UT, CST Noon)



Vertical Tsyganenko Cutoff Difference on Mar 22, 2006 at 20km
6UT-18UT (CST_Midnight - CST_Noon)



Quiet-Time Directional Cutoffs TS05 & IGRF Differences

