

# NRL SAMI2/3 IONOSPHERE MODELS AT CCMC AND FUTURE UPGRADES

J.D. Huba

Plasma Physics Division  
Naval Research Laboratory  
Washington, DC

CCMC Workshop  
April 2014  
Annapolis MD

Acknowledge: J. Krall, G. Lu, M. Swisdak and G. Joyce (deceased)  
Research supported by NRL 6.1 Base Funds and NASA

- magnetic field: IGRF-like
- interhemispheric
- **Nonorthogonal, nonuniform fixed grid**
- seven (7) ion species (**all ions are equal**):  
 $H^+$ ,  $He^+$ ,  $N^+$ ,  $O^+$ ,  $N_2^+$ ,  $NO^+$ , and  $O_2^+$ 
  - solve continuity and momentum for all 7 species
  - solve temperature for  $H^+$ ,  $He^+$ ,  $O^+$ , and  $e^-$
- Plasma motion
  - $\mathbf{E} \times \mathbf{B}$  drift perpendicular to  $\mathbf{B}$
  - **Ion inertia included parallel to  $\mathbf{B}$**
- neutral species: NRLMSISE00 and HWM93
- chemistry: 21 reactions + recombination
- photoionization: Daytime (EUVAC) and nighttime

- ion continuity

$$\frac{\partial n_i}{\partial t} + \nabla \cdot (n_i \mathbf{V}_i) = P_i - L_i n_i$$

- ion velocity

$$\begin{aligned} \frac{\partial \mathbf{V}_i}{\partial t} + \mathbf{V}_i \cdot \nabla \mathbf{V}_i = & -\frac{1}{\rho_i} \nabla P_i + \frac{e}{m_i} \mathbf{E} + \frac{e}{m_i c} \mathbf{V}_i \times \mathbf{B} + \mathbf{g} \\ & - \nu_{in} (\mathbf{V}_i - \mathbf{V}_n) - \sum_j \nu_{ij} (\mathbf{V}_i - \mathbf{V}_j) \end{aligned}$$

- ion temperature

$$\frac{\partial T_i}{\partial t} + \mathbf{V}_i \cdot \nabla T_i + \frac{2}{3} T_i \nabla \cdot \mathbf{V}_i + \frac{2}{3} \frac{1}{n_i k} \nabla \cdot \mathbf{Q}_i = Q_{in} + Q_{ij} + Q_{ie}$$

- electron momentum

$$0 = -\frac{1}{n_e m_e} b_s \frac{\partial P_e}{\partial s} - \frac{e}{m_e} E_s$$

- electron temperature

$$\frac{\partial T_e}{\partial t} - \frac{2}{3} \frac{1}{n_e k} b_s^2 \frac{\partial}{\partial s} \kappa_e \frac{\partial T_e}{\partial s} = Q_{en} + Q_{ei} + Q_{phe}$$

- NRLMSISE-00 (Picone et al)
  - neutral composition H, He, O, N, N<sub>2</sub>, O<sub>2</sub>
  - neutral temperature  $T_n$
- HWM93 (Hedin)
  - neutral wind  $V_n$  (meridional/zonal)
- electric field (Fejer/Scherliess)
  - $E \times B$  drift  $V_E$  (vertical at magnetic equator)

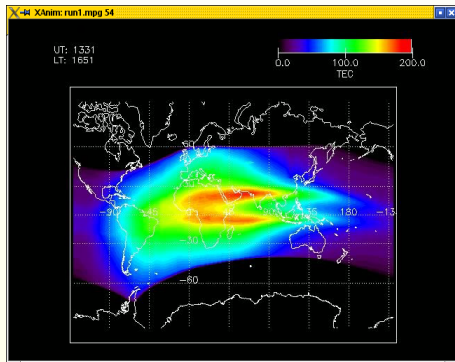
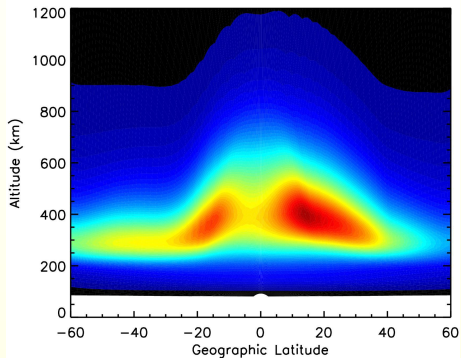
# EXAMPLE OUTPUT

SAMI2 (electron density contour); SAMI3 (TEC)

UT 03:15 SAMI3/NRLMSIS

LT 15:21 Longitude 181

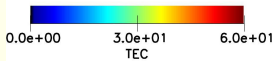
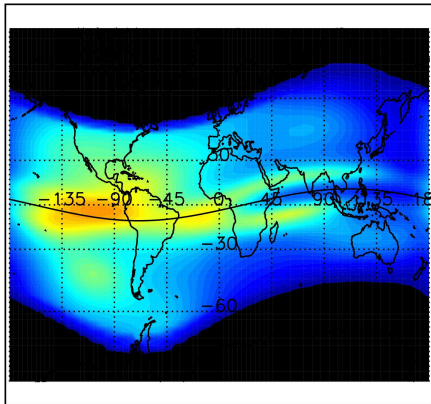
0.0e+00 2.0e+06 4.0e+06  
Electron Density ( $\text{cm}^{-3}$ )



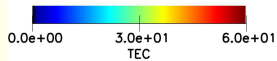
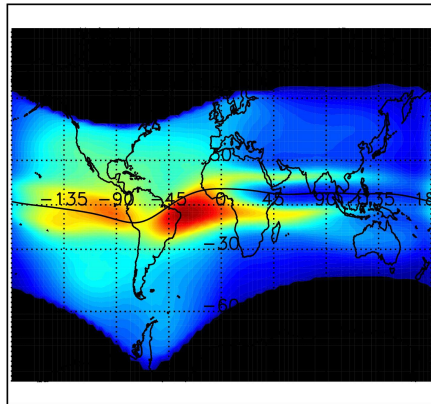
# UPGRADES

- neutral wind dynamo electric field
- IGRF (with Gang Lu)
- SAMI3/Volland-Stern/Weimer
- SAMI3/RCM
- flux corrected transport (4th order, partial donor cell method)

TILTED: HWM93



APEX: HWM93



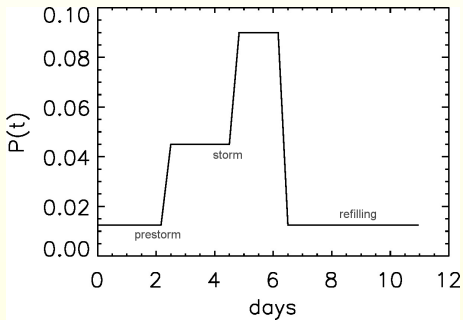


- corotation potential

$$\Phi_{\text{cor}} = -92/r(\text{kV})$$

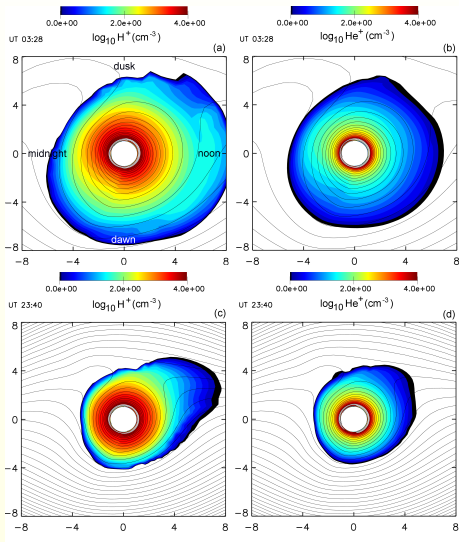
- volland/stern/maynard/chen potential ( $K_p = 6$ )

$$\Phi_{\text{vsmc}} = Ar^2 \sin(\phi) \quad A = \frac{P(t)}{(1. - 0.159K_p + 0.0093K_p^2)^3} (\text{kV})$$



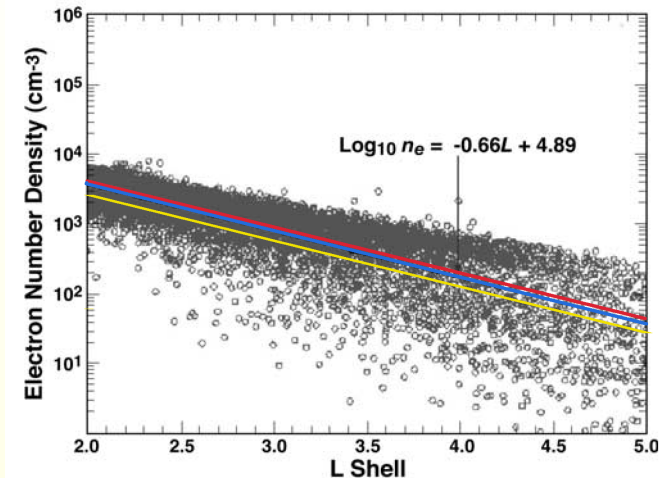
# H<sup>+</sup>/He<sup>+</sup> CONTOURS

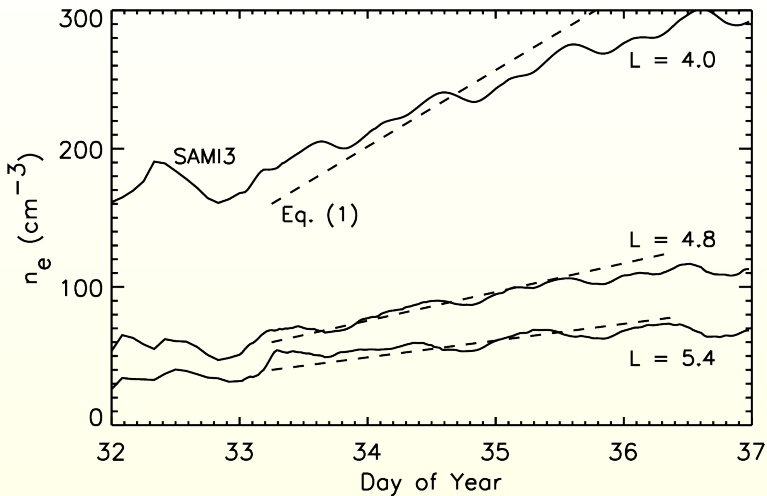
hydrogen / helium ion density contours at t = 52 hrs and 96 hrs in equatorial plane



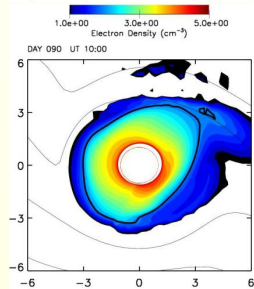
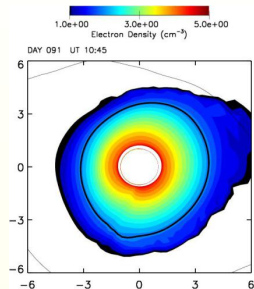
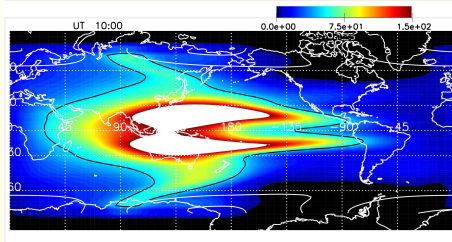
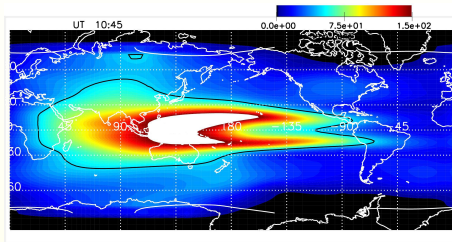
# ELECTRON DENSITY AND SAMI3 DATA

line plots of electron density vs  $L$  shell; empirical relationship (Berube et al., JGR, 2005)



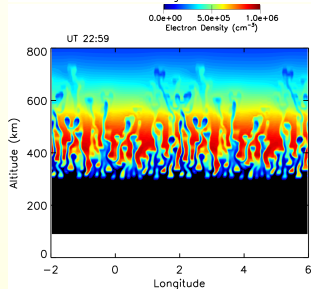
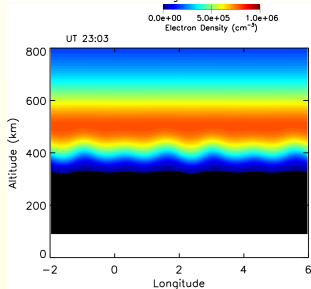
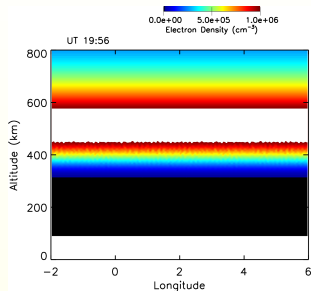
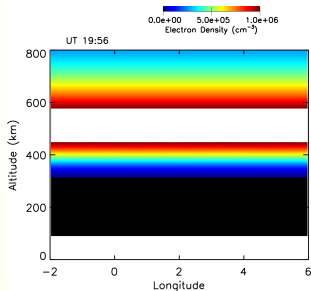


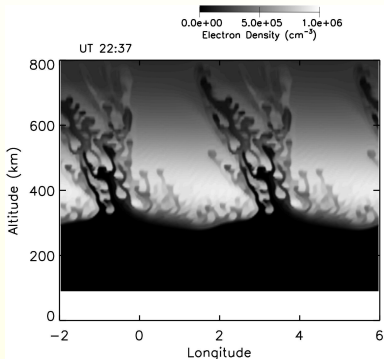
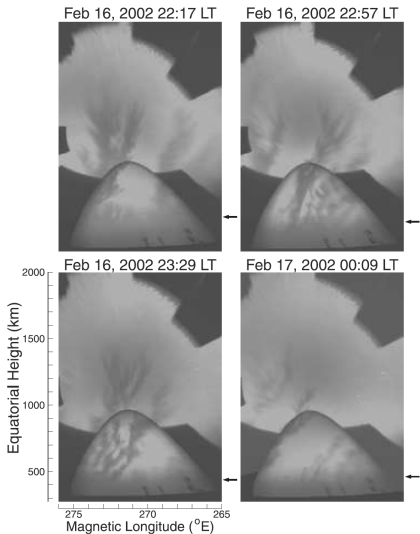
$$dn_e/dt = 3.81(6.8/L)^{4.94} \text{cm}^{-3} \text{day}^{-1} \quad (1)$$



# LOW/HIGH ORDER SAMI3

sami3/esf (uniform 100 m/s zonal wind)

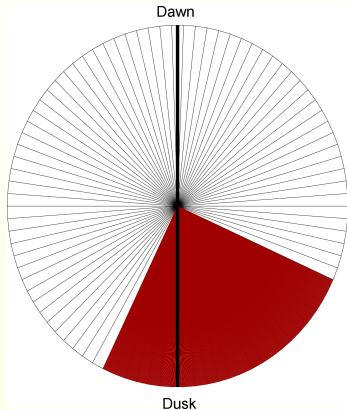




# GLOBAL SOLUTION

incorporate a high-resolution grid in a global model, i.e., SAMI3

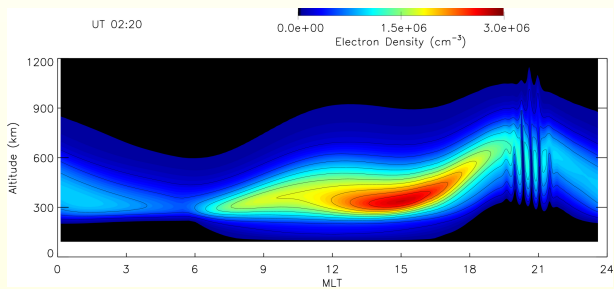
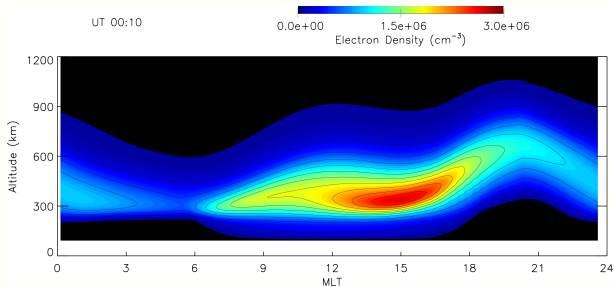
- reference frame: copernican (sun-fixed: rotating earth)
- coarse mesh: 90 grid points
- zonal resolution  $\sim 500$  km
- high resolution mesh: 956 grid points between  $\sim 16:30$  MLT -  $22:30$  MLT
- zonal resolution  $\sim .0625^\circ$  or  $\sim 7$  km





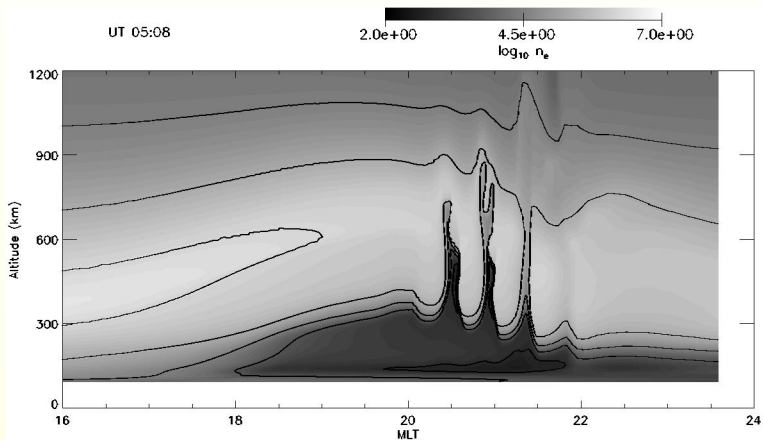
# FIRST GLOBAL MODEL OF ESF

Huba and Joyce, *GRL*, 2010



# APPLICATION TO SAMI3

hi-res global code (under development)



# FUTURE UPGRADES AT CCMC: SAMI3

modeling the near-earth space environment

- global low- to mid-latitude ionosphere model with IGRF magnetic field
- global low- to high-latitude ionosphere model with Volland/Stern/Maynard/Chen and Weimer potentials
- global low- to high-latitude ionosphere model with RCM potential