

The WINDMI Model at the CCMC

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$$L \frac{dl}{dt} = V_{\text{sw}}(t) - V + M \frac{dl_1}{dt} \quad \text{Low-Dimensional dynamical system:} \quad (1)$$

$$C \frac{dV}{dt} = I - I_1 - I_{\text{ps}} - \Sigma V \quad \text{WINDMI} \quad (2)$$

$$\frac{3}{2} \frac{dp}{dt} = \frac{\Sigma V^2}{\Omega_{\text{cps}}} - u_0 p K_{\parallel}^{1/2} \Theta(u) - \frac{p V A_{\text{eff}}}{\Omega_{\text{cps}} B_{\text{tr}} L_y} - \frac{3p}{2\tau_E} \quad (3)$$

$$\frac{dK_{\parallel}}{dt} = I_{\text{ps}} V - \frac{K_{\parallel}}{\tau_{\parallel}} \quad (4)$$

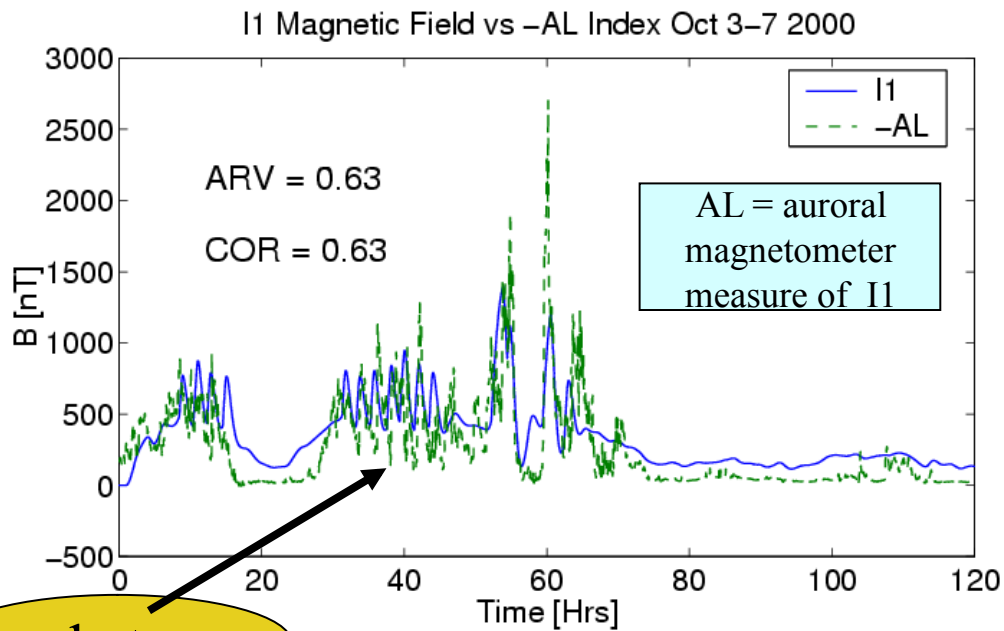
$$L_I \frac{dl_1}{dt} = V - V_I + M \frac{dl}{dt} \quad (5)$$

$$C_I \frac{dV_I}{dt} = I_1 - I_2 - \Sigma_I V_I \quad (6)$$

$$L_2 \frac{dl_2}{dt} = V_I - (R_{\text{prc}} + R_{A2}) I_2 \quad (7)$$

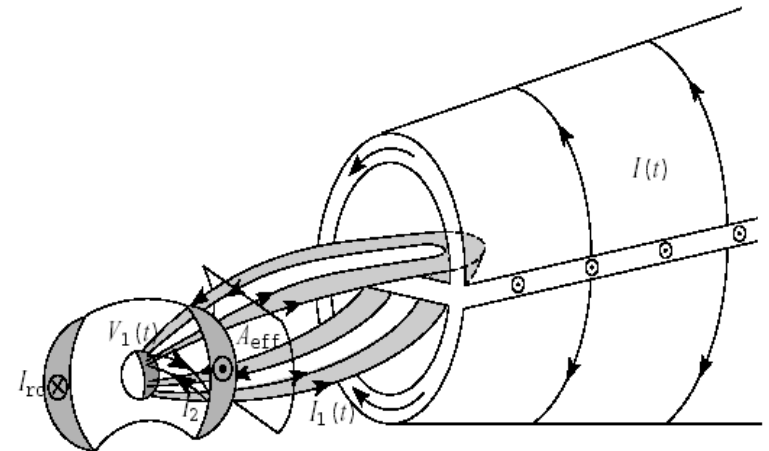
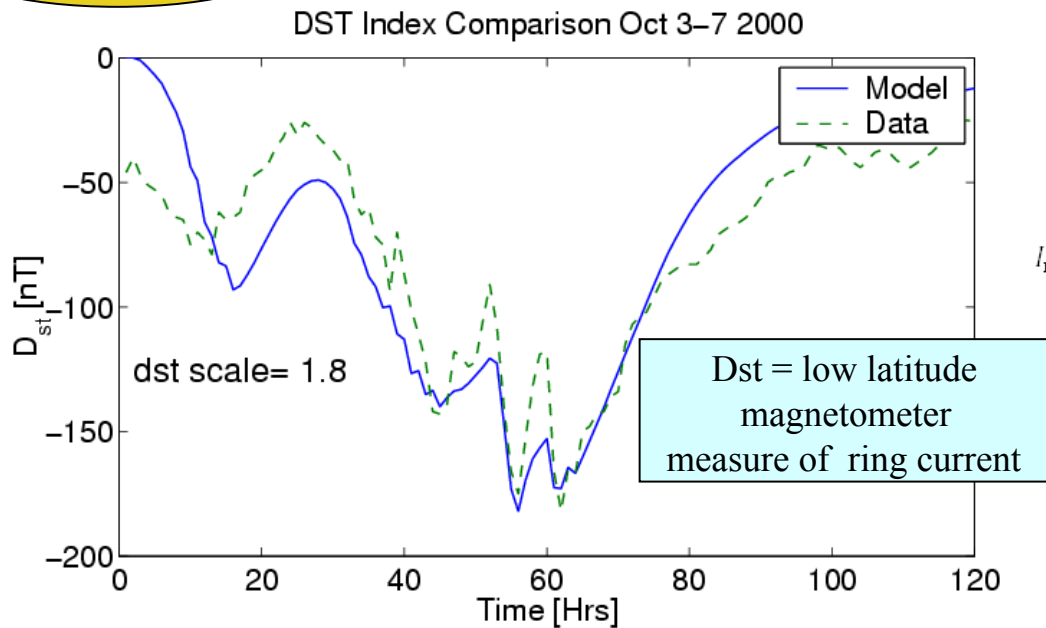
$$\frac{dW_{\text{rc}}}{dt} = R_{\text{prc}} I_2^2 + \frac{p V A_{\text{eff}}}{B_{\text{tr}} L_y} - \frac{W_{\text{rc}}}{\tau_{\text{rc}}} \quad (8)$$

Derived by projection of pde's on to key basis functions: not simple dynamics. attractors, bifurcations, chaos from strange attractors. Limits to predictability of Weather systems



WINDMI-RC Simulation for 4-6 Oct 2000 GEM Storm

8 substorms



Spencer et al, Mays et al
JGRs 2007, 2010, 2012,
Patra 2011

Events and Validation Publications

3–7 October 2000 AL and Dst for [Spencer-JGR-2007]
15–24 April 2002 with 8 substorms $-Al > 1000$ [Mays]
August 2001 Dst tail current contribution [Spencer]
25 February 2008 isolated moderate substorm $-Al=70$ [Patra]
4-8 Apr 2010 Dst – 80nT and $-AL > 1300, 900$
5-12 August 2011 Dst min. -113 nT AL min. -2000 nT
9-14 September 2011 Dst min. -64 nT AL min. -1300 nT
17-19 September 2011 Dst min. -58 nT AL min. -1200 nT
26-30 September 2011 Dst min. -103 nT AL min. -1600 nT

Publications:

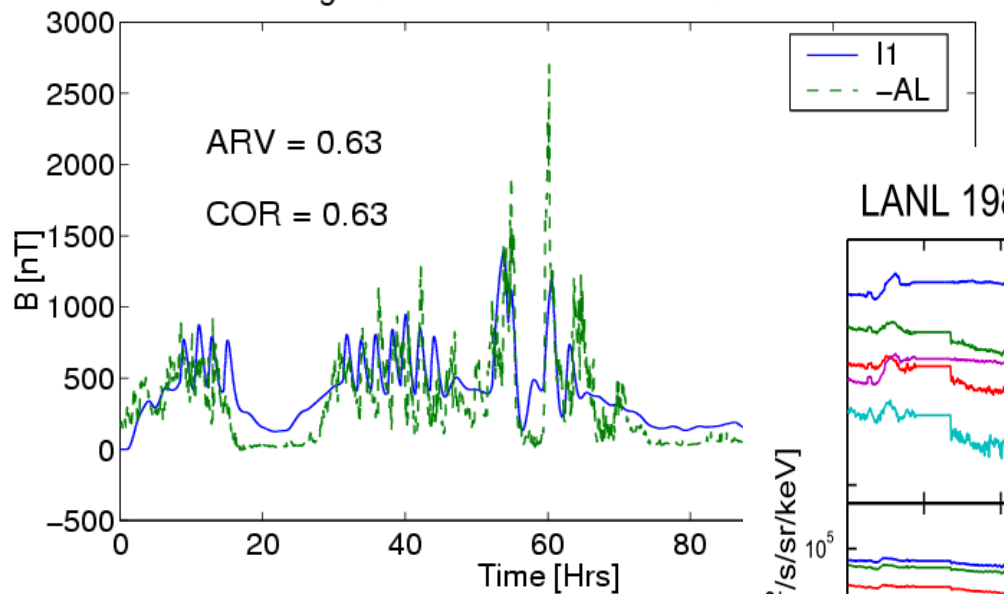
Spencer et al. JGR 2007, 2010, 2011

Mays et al. Space Weather 2009

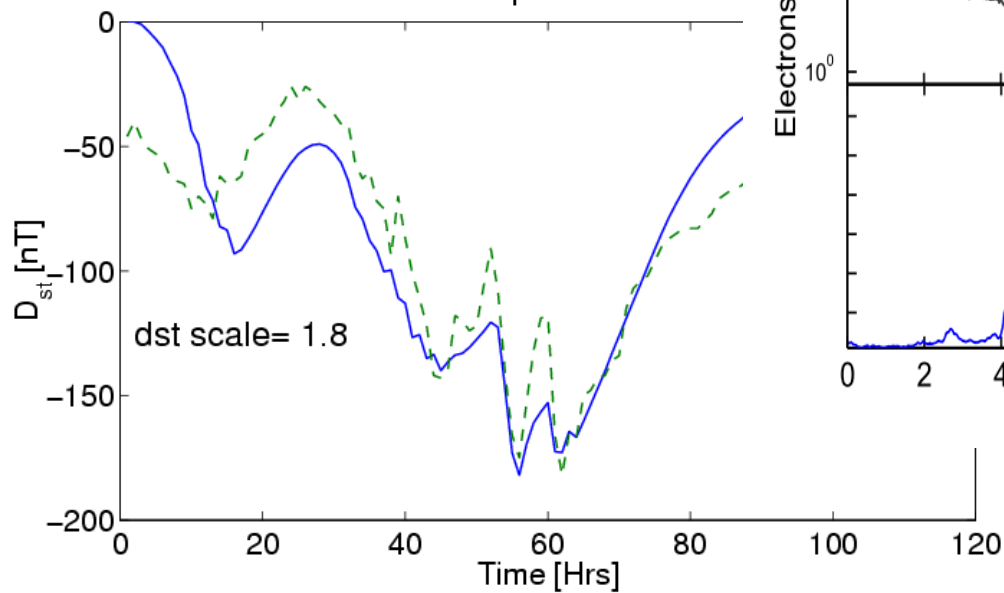
Patra et al. JGR 2011

Spencer et al JGR 2012

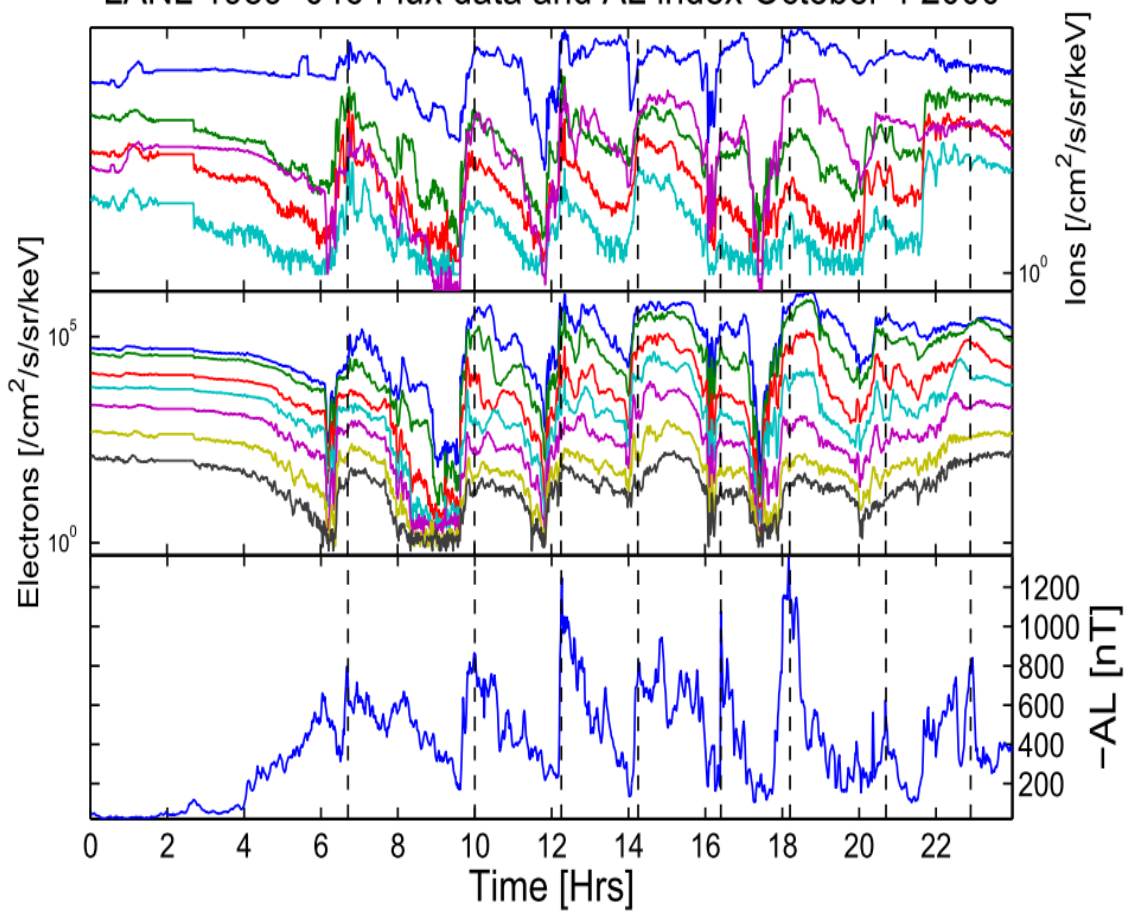
I1 Magnetic Field vs -AL Index Oct 3-7 2000



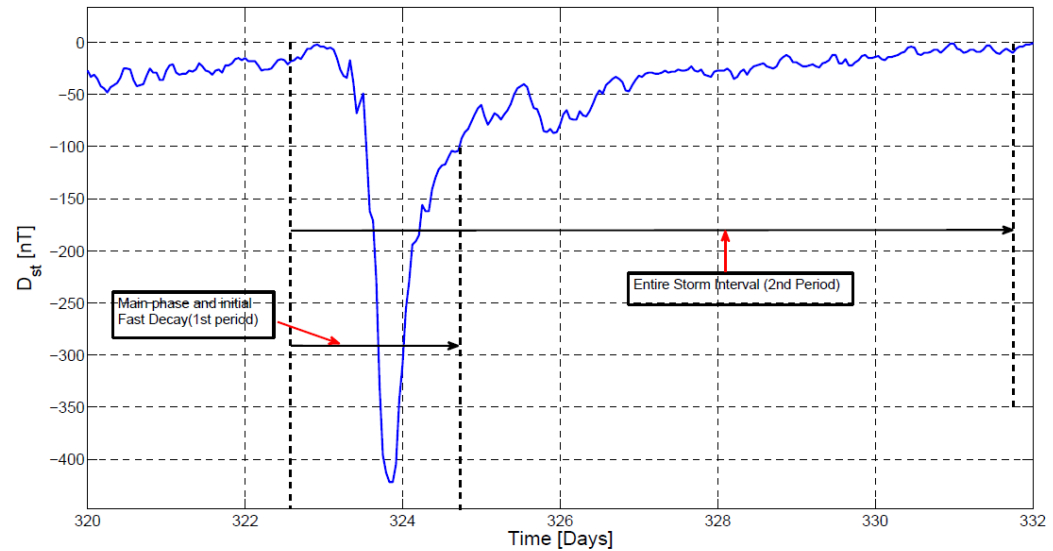
DST Index Comparison Oct 3-7 2000



LANL 1989-046 Flux data and AL index October 4 2000



Typical two phase decay of a geomagnetic storm as indicated by the corresponding Dst Index.

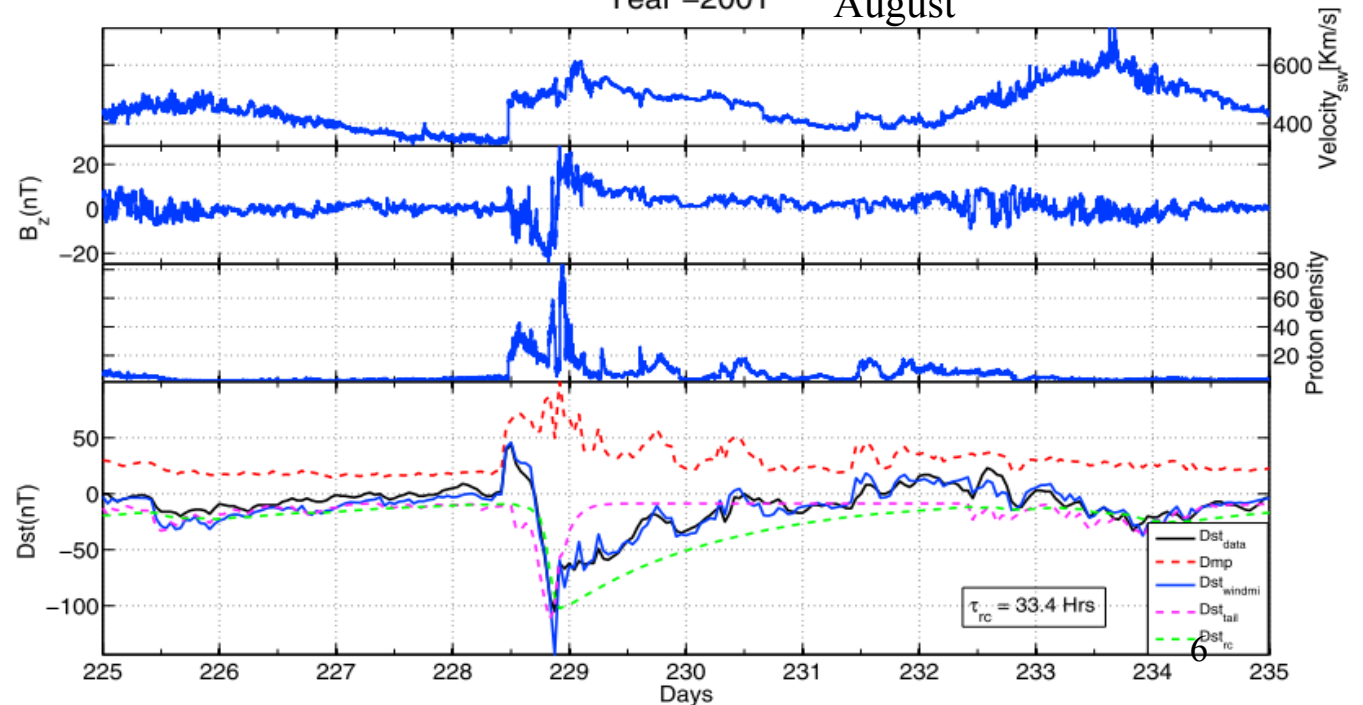


WINDMI Storms
 [Spencer et al JGR 2010]

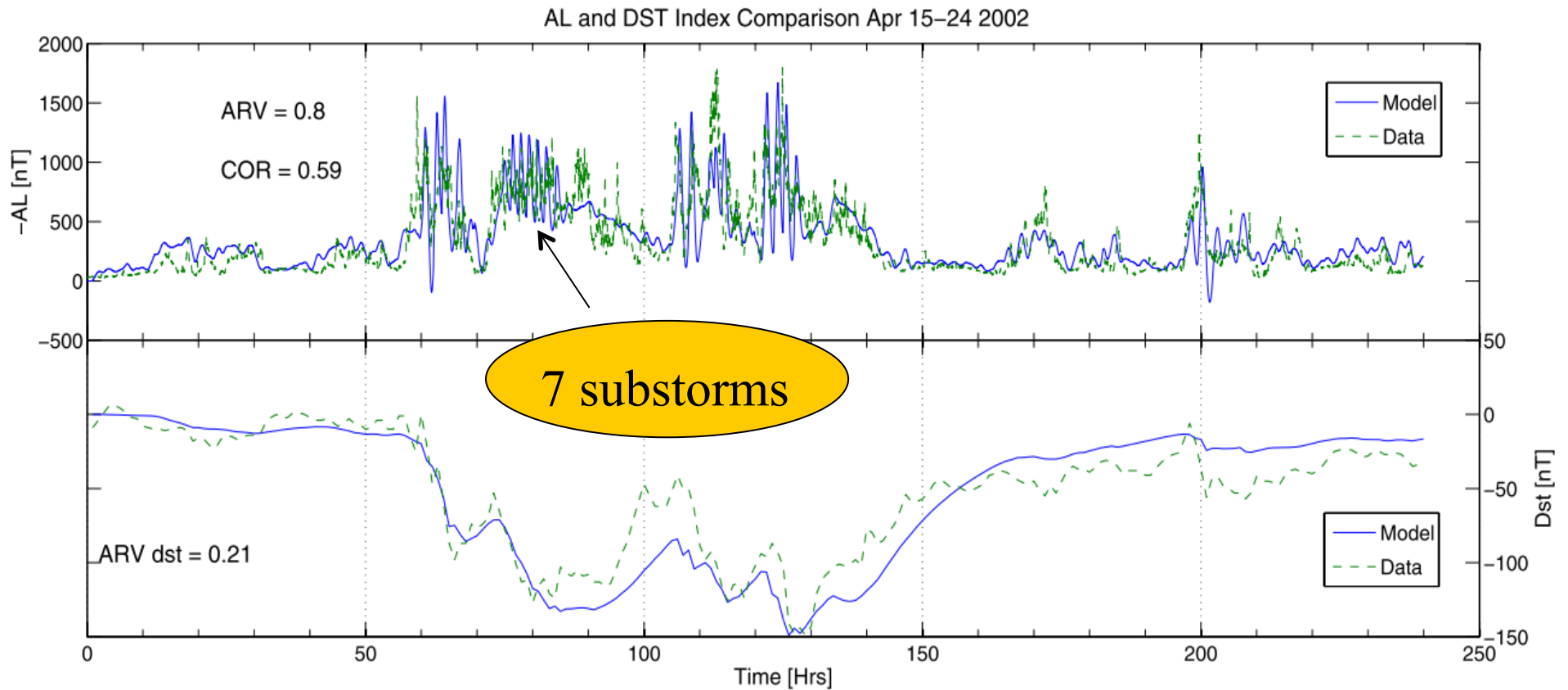
Windmi model calculation of Dst after including contributions from magnetospheric currents.

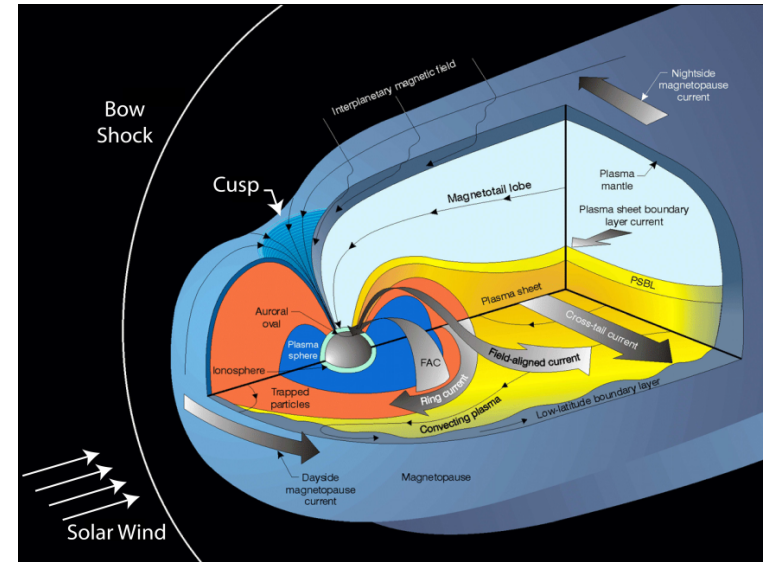
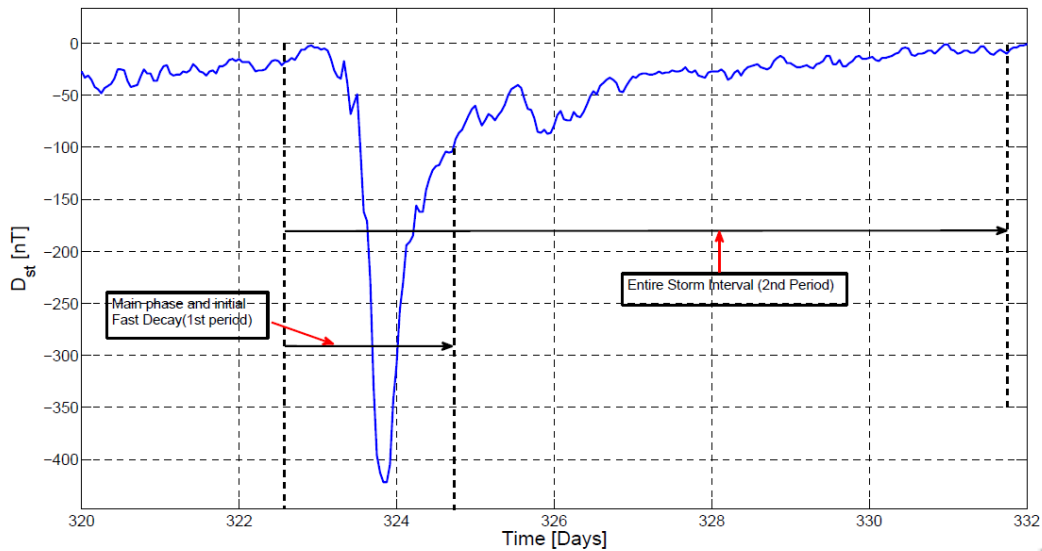
$$Dst_{windmi} = Dst_{rc} + Dst_{mp} + Dst_t$$

Year -2001 August



The 15–24 April 2002 prediction using VswN(Newell) with optimized parameters from equal weighting of AL and Dst.





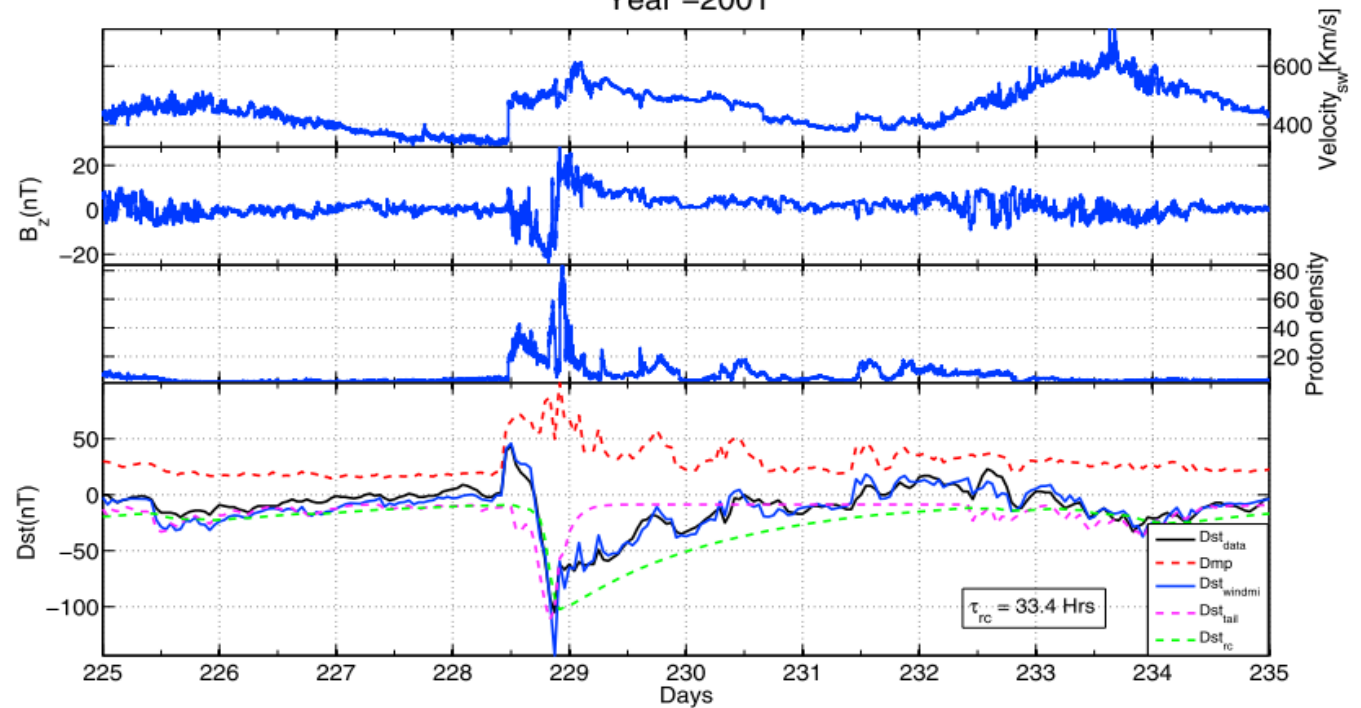
$$Dst_{windmi} = Dst_{rc} + Dst_{mp} + Dst_t$$

Year -2001

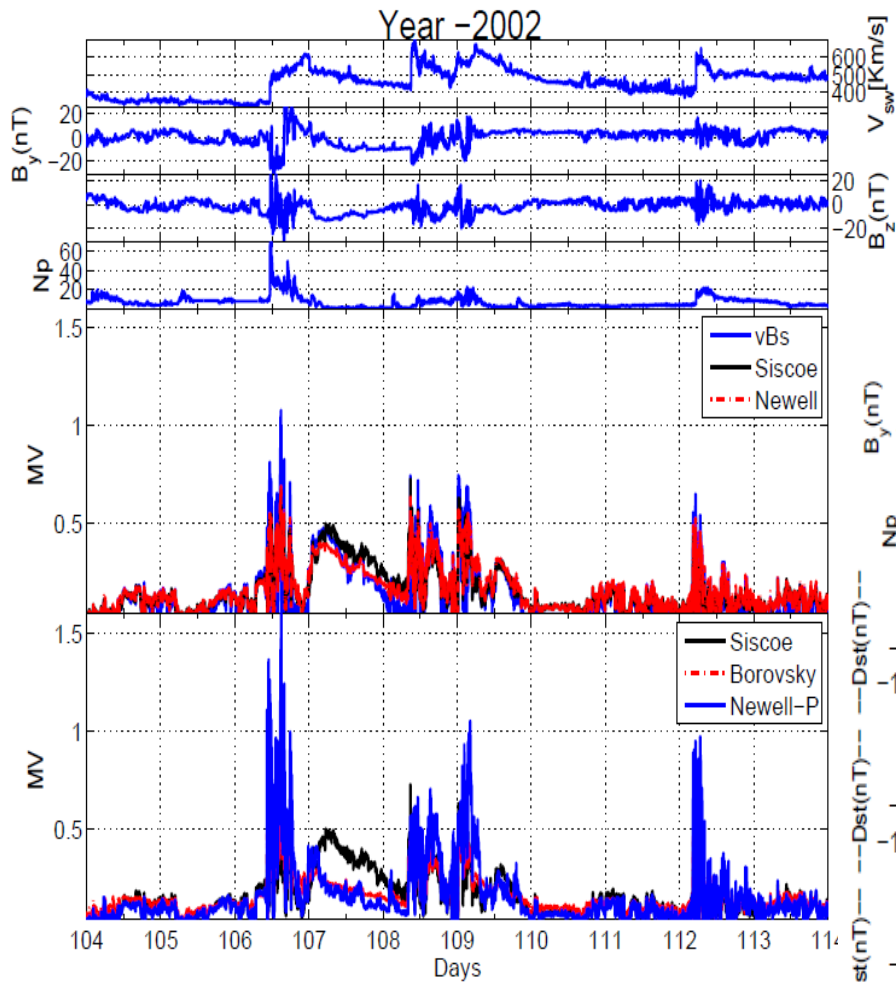
Top: Typical two phase decay of a geomagnetic storm as indicated by the corresponding Dst Index.

Top Right: Magnetospheric currents

Right: Windmi model estimation of Dst after including contributions from magnetospheric currents.

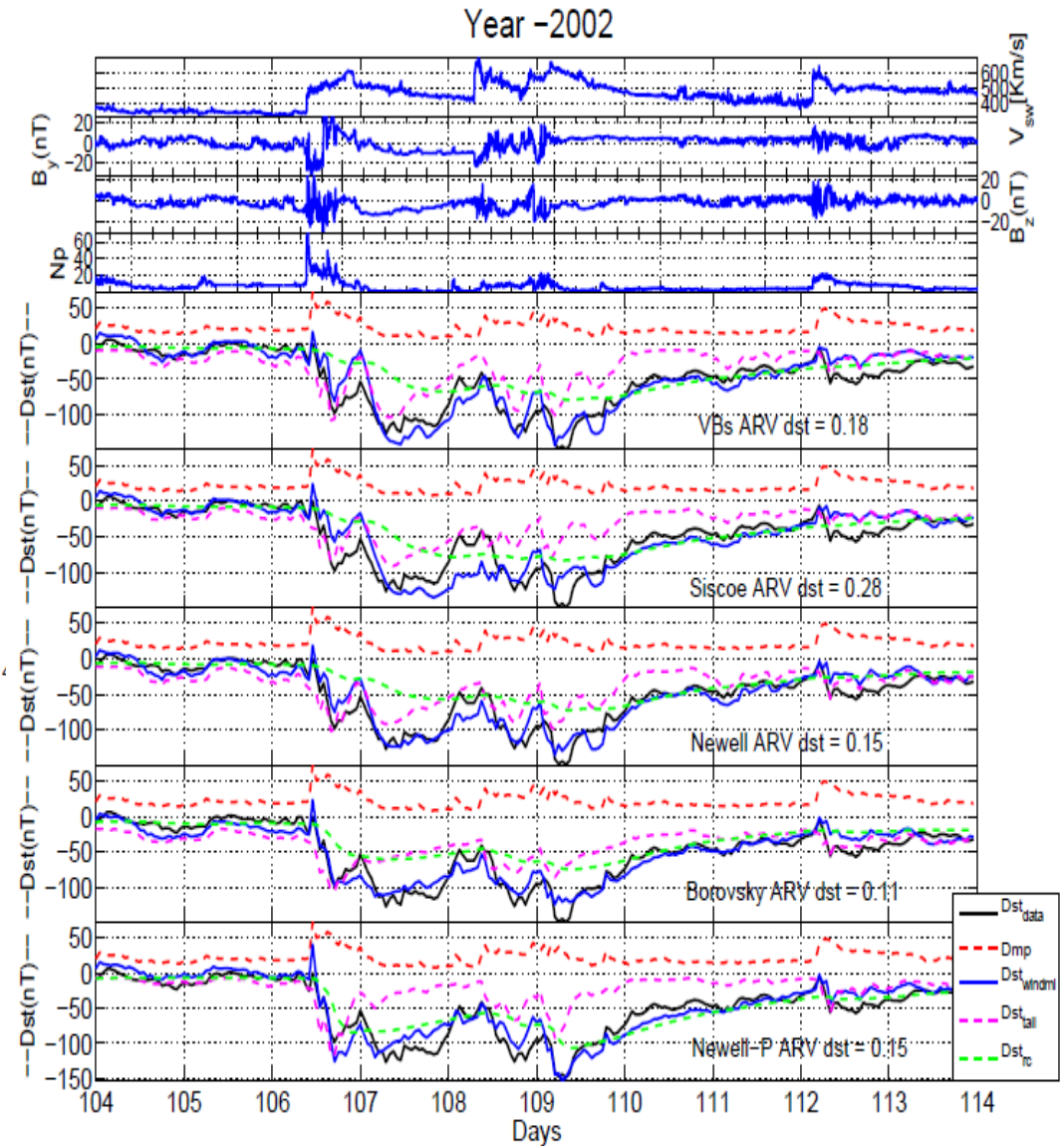


Generalized to include spectrum of SW Couplers



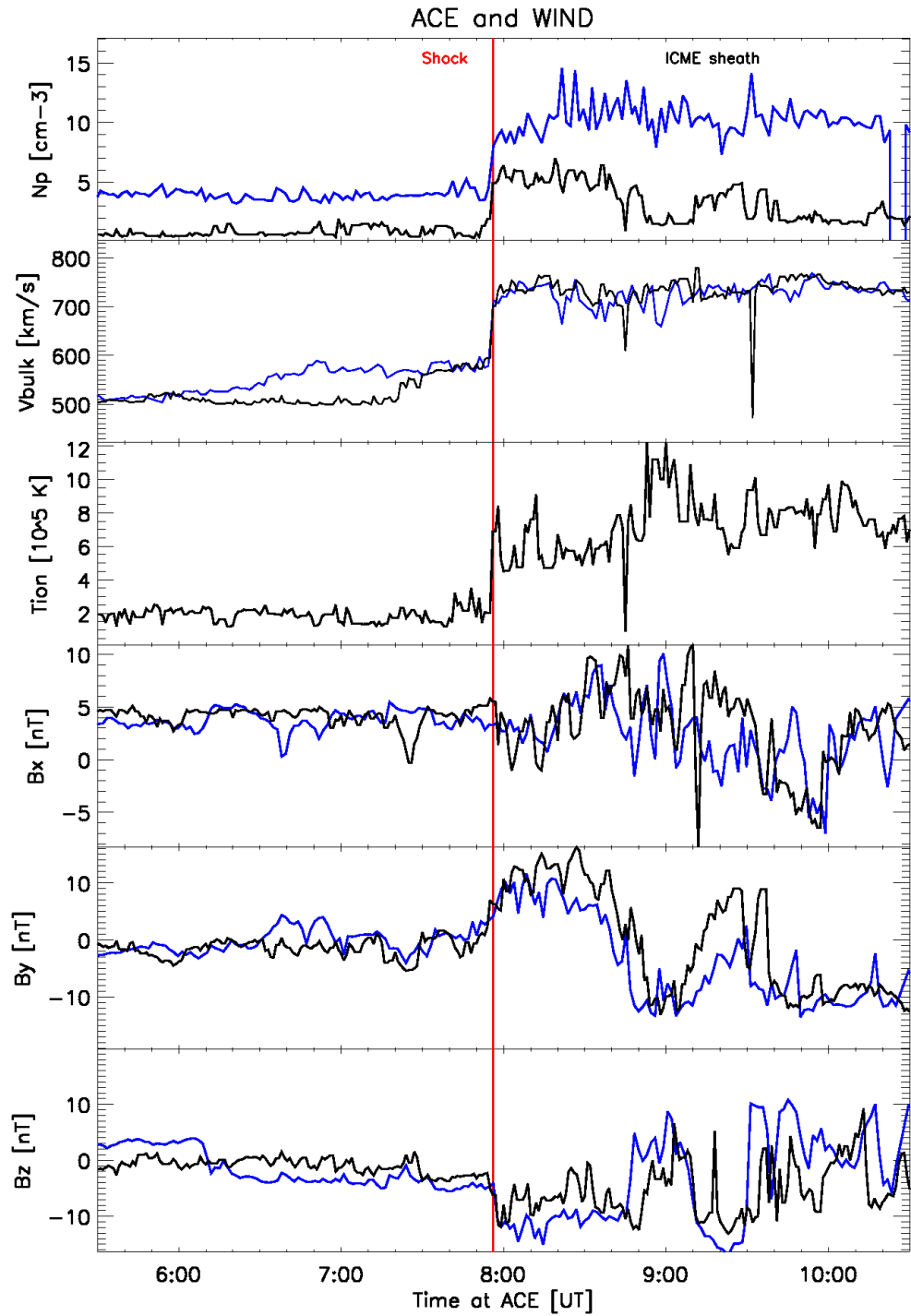
Top: Different coupling functions for a category II storm.

Right: Contributions from different magnetospheric currents to the Dst index using various coupling functions. Notice the different relative contributions from the tail and ring currents.

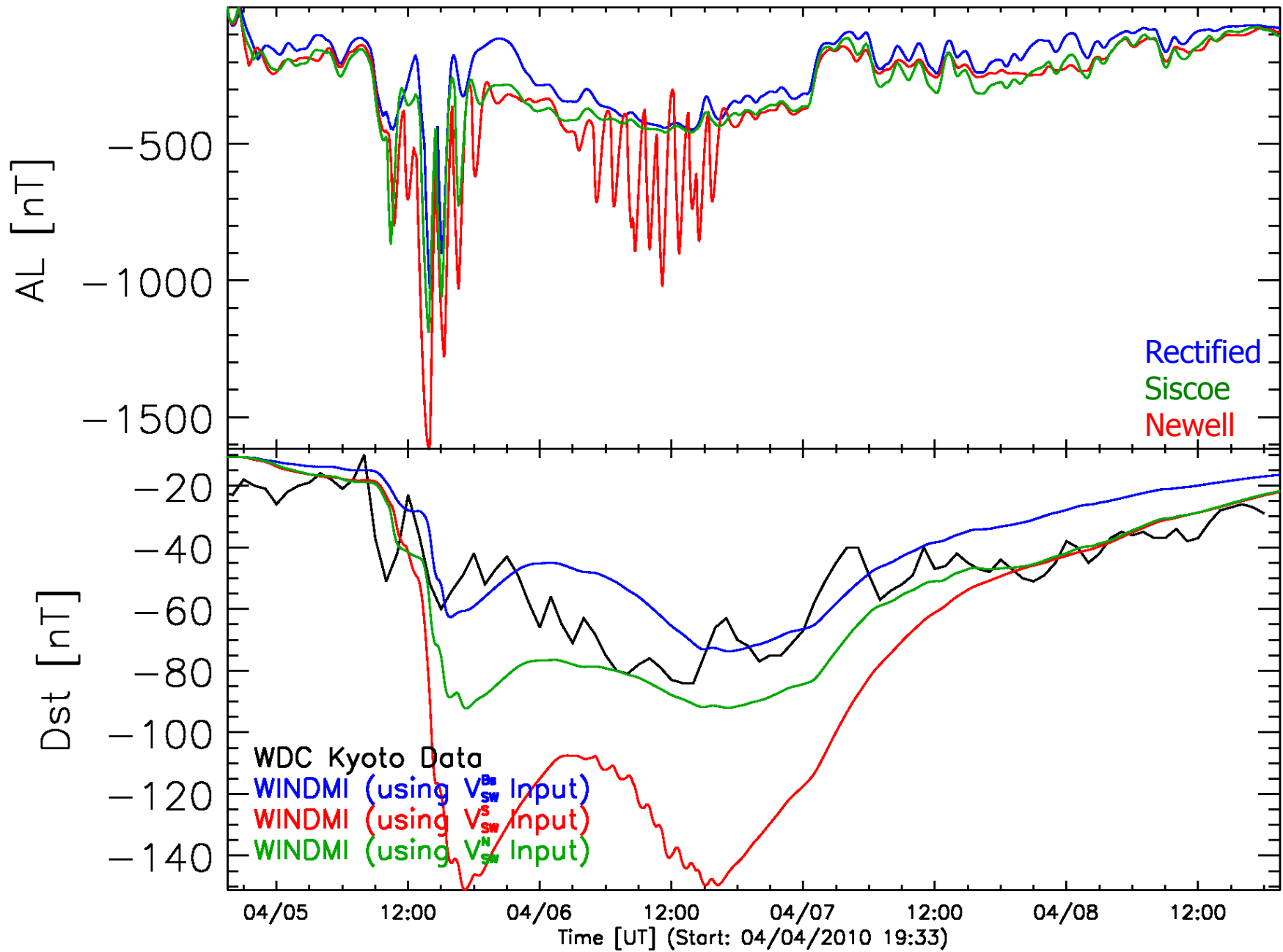


04 April 2010

Real-time ACE and WIND
solar wind data
comparison.



WINDMI Results (04/04-04/08/2010)



The Newell function performs best in predicting the *AL*;
 The Rectified function performs best in predicting the *Dst*

Feb. 2006 - Aug. 2008 Selected Events, Mean ARV		
Input	Mean <i>AL</i> ARV	Mean <i>Dst</i> ARV
<i>Rectified</i> V_{sw}^{Bs}	0.38 ± 0.21	0.37 ± 0.27
<i>Siscoe</i> V_{sw}^S	0.41 ± 0.16	0.42 ± 0.23
<i>Newell</i> V_{sw}^N	0.33 ± 0.17	0.54 ± 0.39

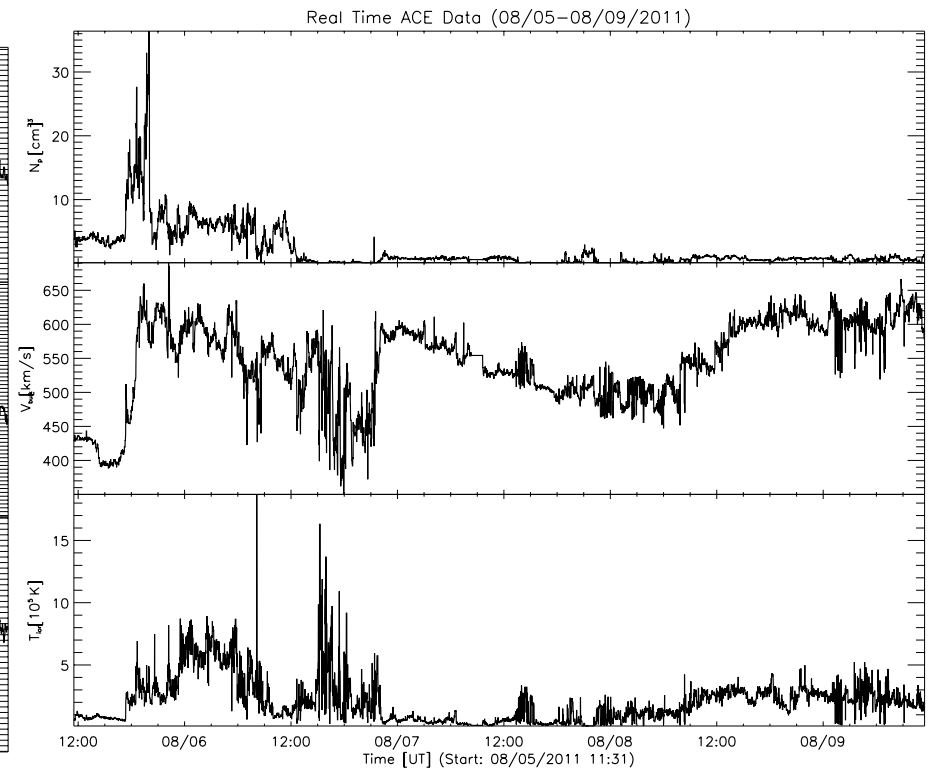
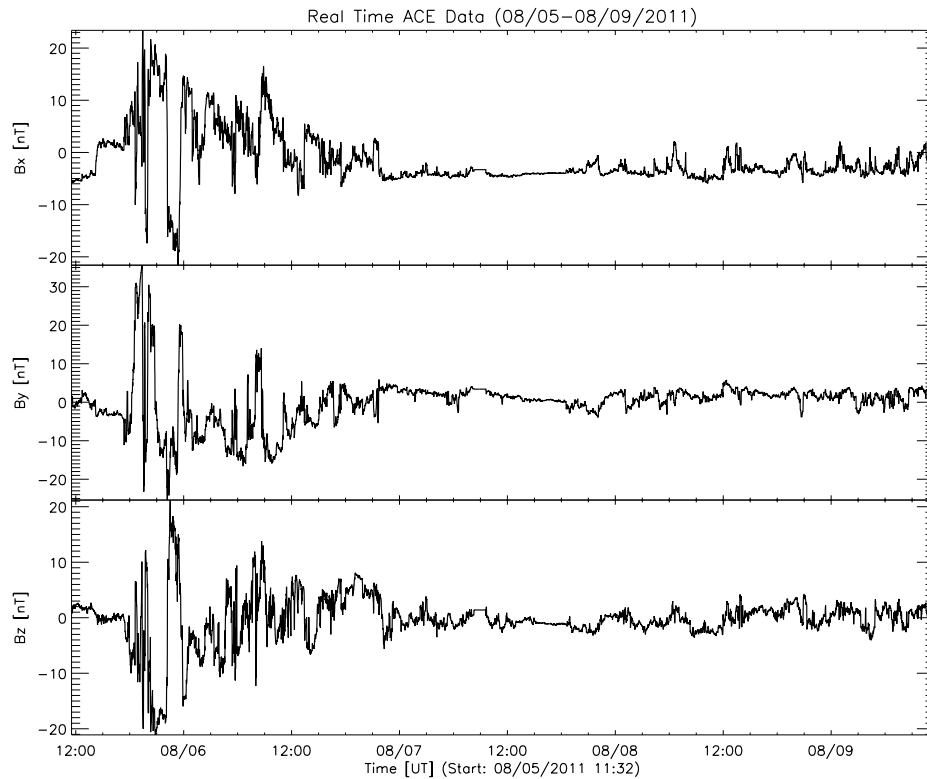
Feb. 2006 - Aug. 2008 Selected Events, Mean COR			
Input	Mean <i>AL</i> COR	Mean <i>AL</i> Direct COR	Mean <i>Dst</i> COR
<i>Rectified</i> V_{sw}^{Bs}	0.62 ± 0.13	0.40 ± 0.20	0.80 ± 0.12
<i>Siscoe</i> V_{sw}^S	0.52 ± 0.15	0.37 ± 0.18	0.77 ± 0.13
<i>Newell</i> V_{sw}^N	0.64 ± 0.12	0.42 ± 0.18	0.79 ± 0.14

Validation statistics for more recent events during solar minimum are in progress.
 (*Mays et al, Space Weather, 2009*)

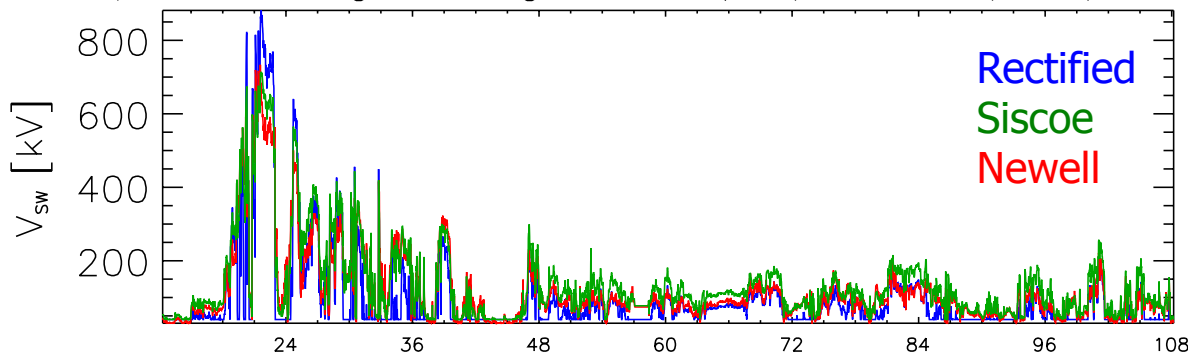
Some Examples of Recently Captured Storms

5-12 August 2011	Dst min. -113 nT	AL min. -2000 nT
9-14 September 2011	Dst min. -64 nT	AL min. -1300 nT
17-19 September 2011	Dst min. -58 nT	AL min. -1200 nT
26-30 September 2011	Dst min. -103 nT	AL min. -1600 nT

5-12 August 2011 Dst min -113 nT, AL min -2000 nT



Input Driving Voltage V_{sw} (08/05-08/09/2011)



From AR1261

M6.0 flare peak 2011-08-03 13:48 UT

M9.3 flare peak 2011-08-04 03:57 UT

Two CMEs merged at ACE, initiated at

2011-08-03 13:55UT (~1350km/s)

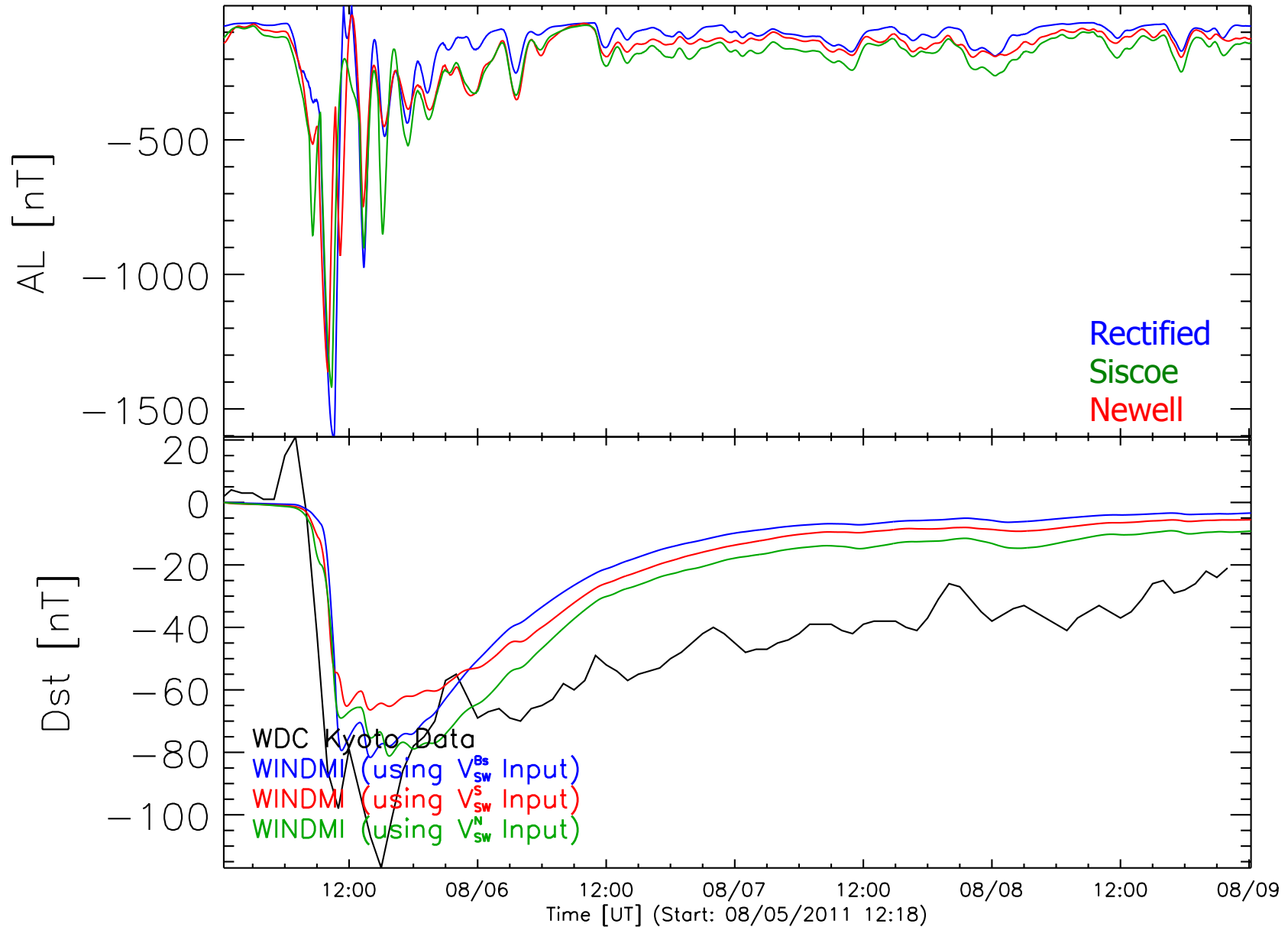
2011-08-04 04:10UT (~1950 km/s)

ACE shock at 2011-08-05 17:20 UT

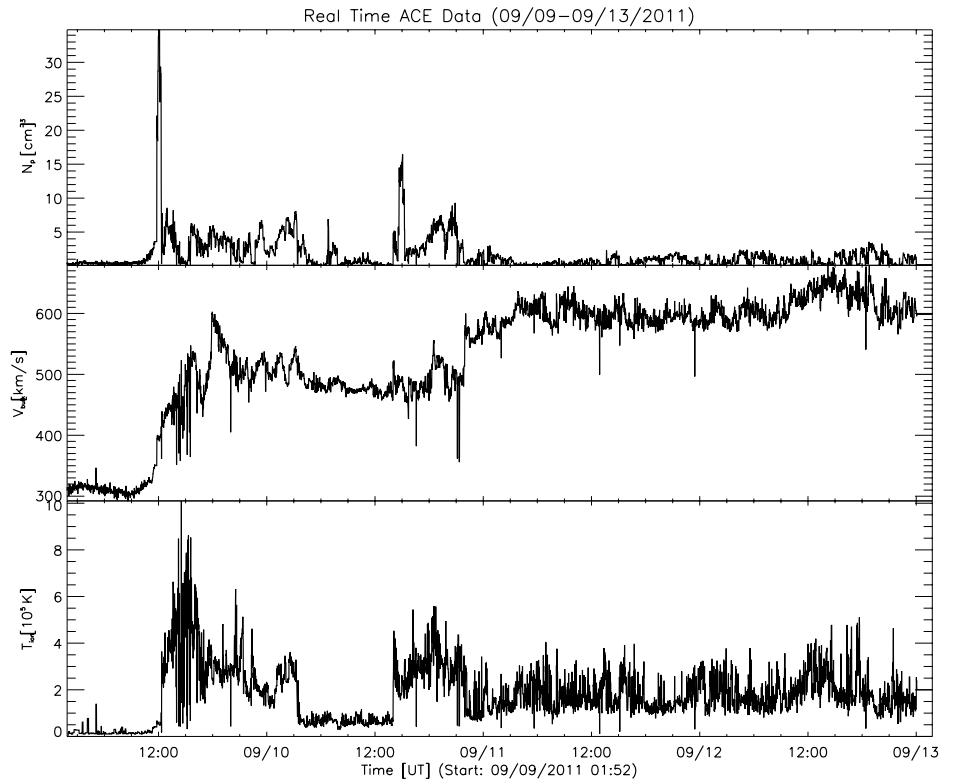
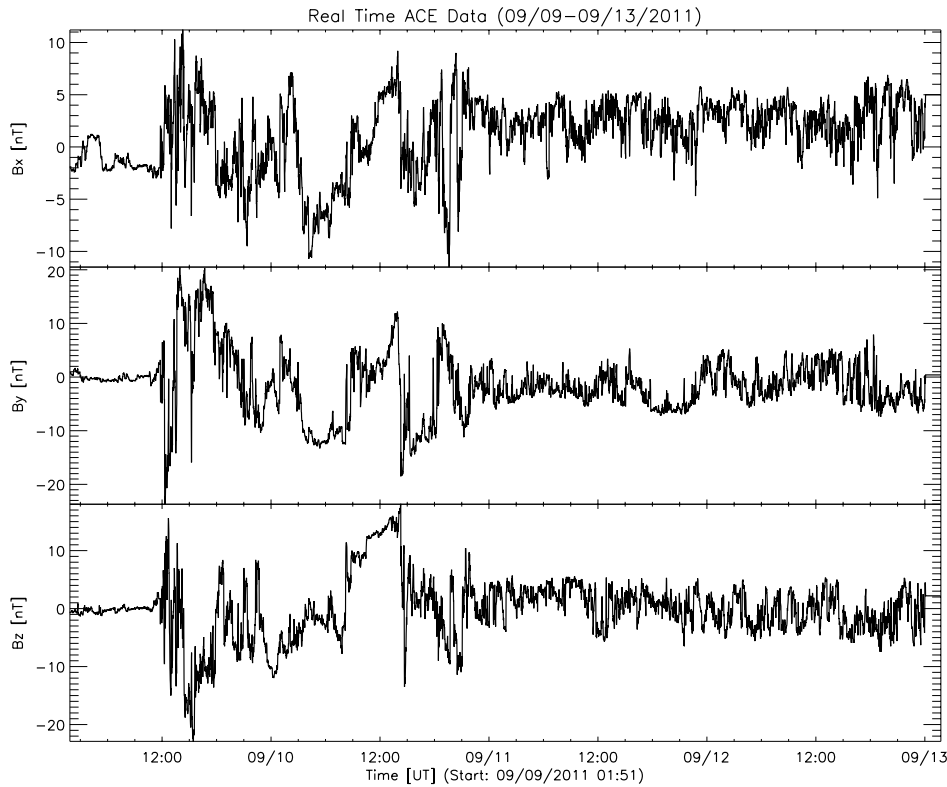
Dst min. -113 nT on 6 Aug

5-12 August 2011 Dst min -113 nT, AL min -2000 nT

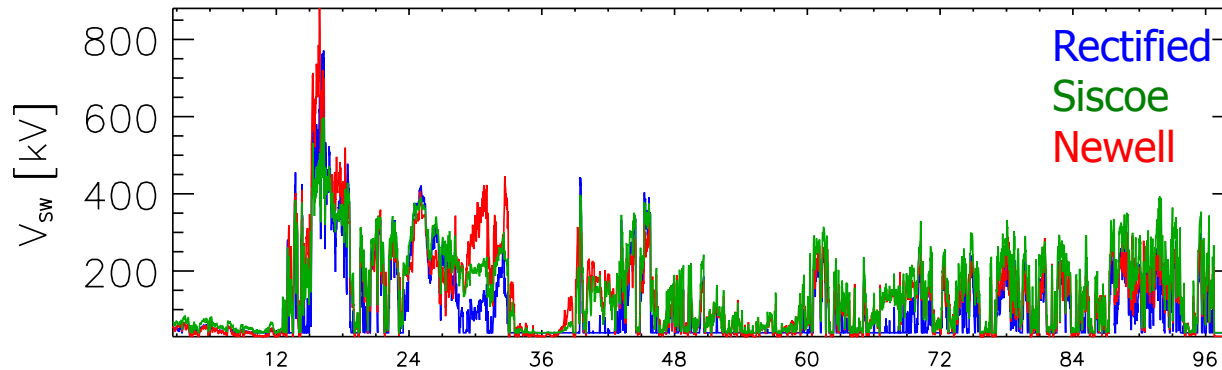
WINDMI Results (08/05-08/09/2011)



9-14 September 2011 Dst min. -64 nT, AL min. -1300 nT

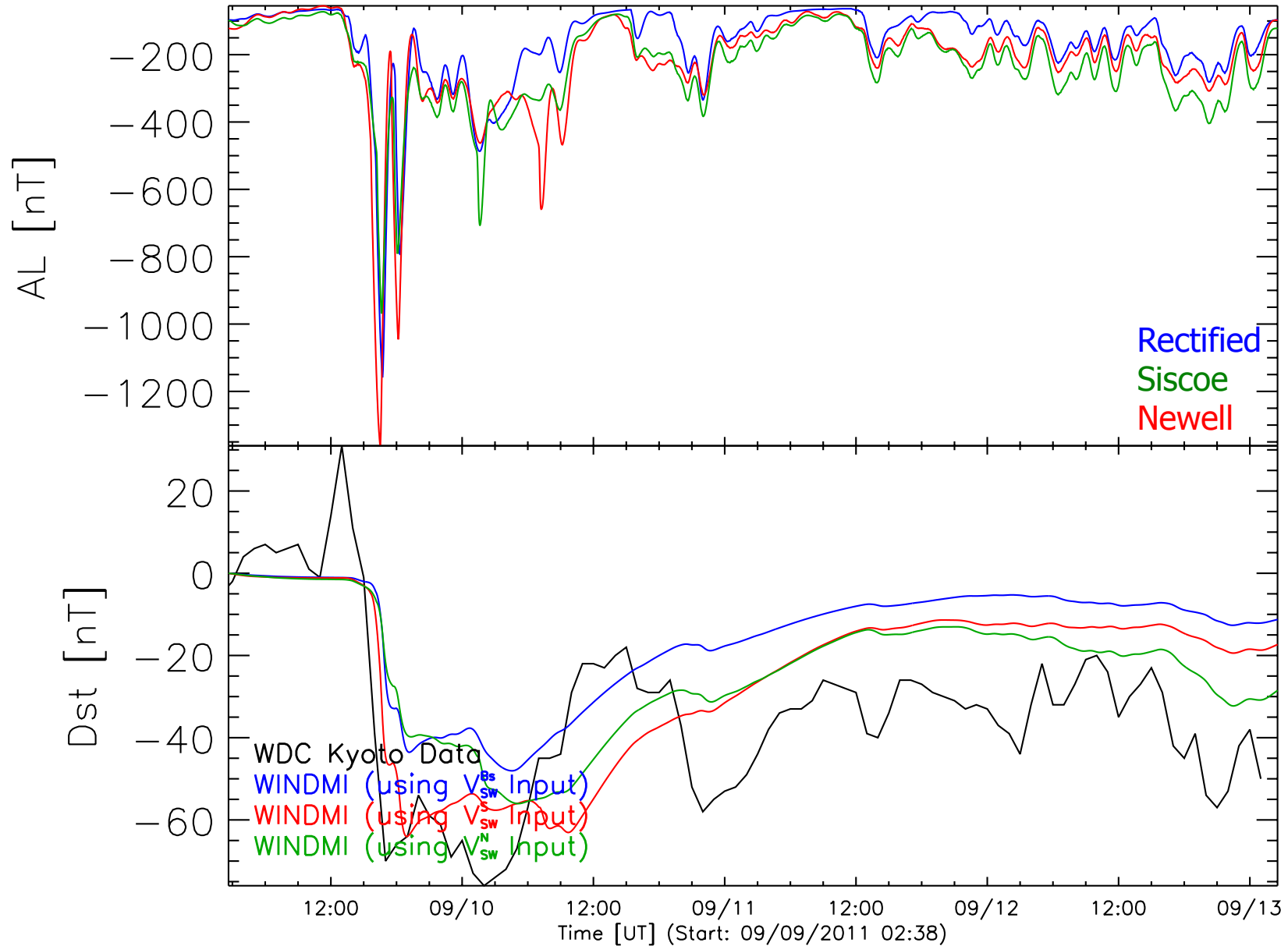


Input Driving Voltage V_{sw} (09/09-09/13/2011)

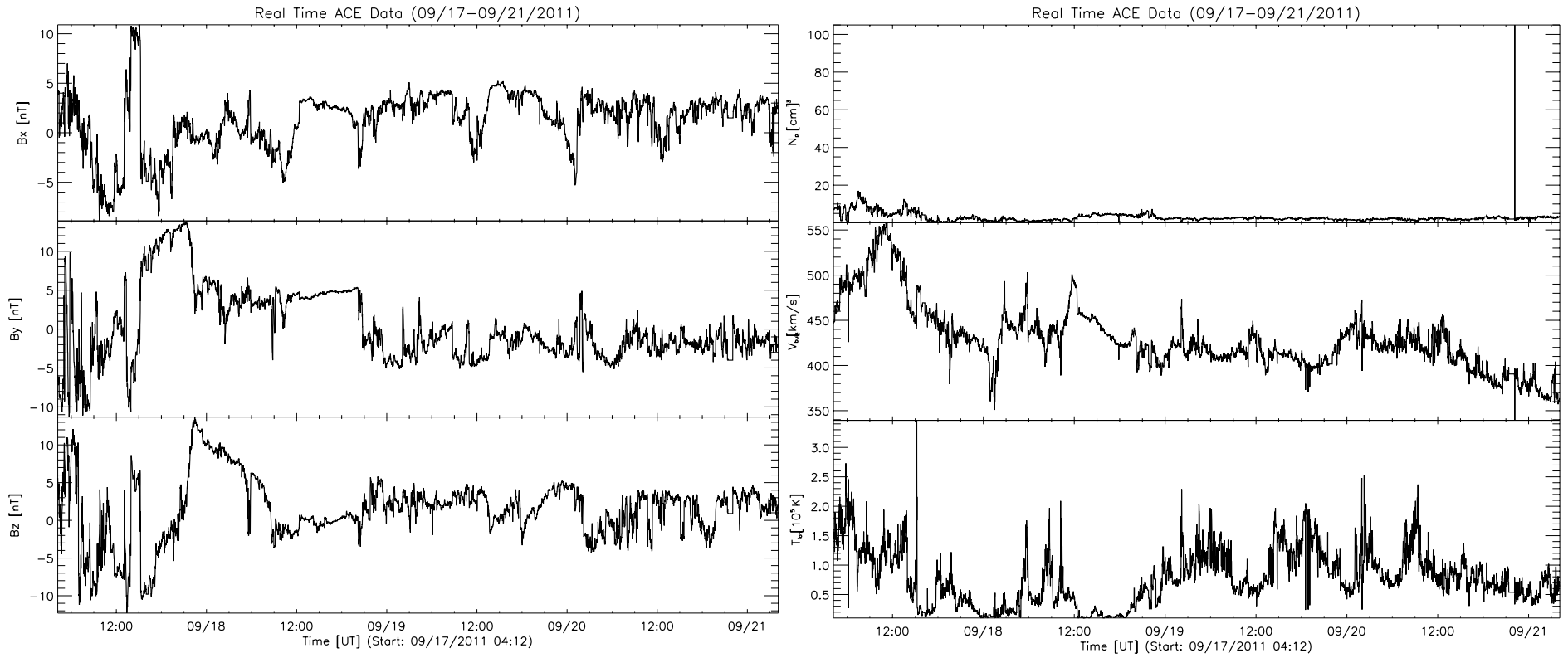


9-14 September 2011 Dst min. -64 nT, AL min. -1300 nT

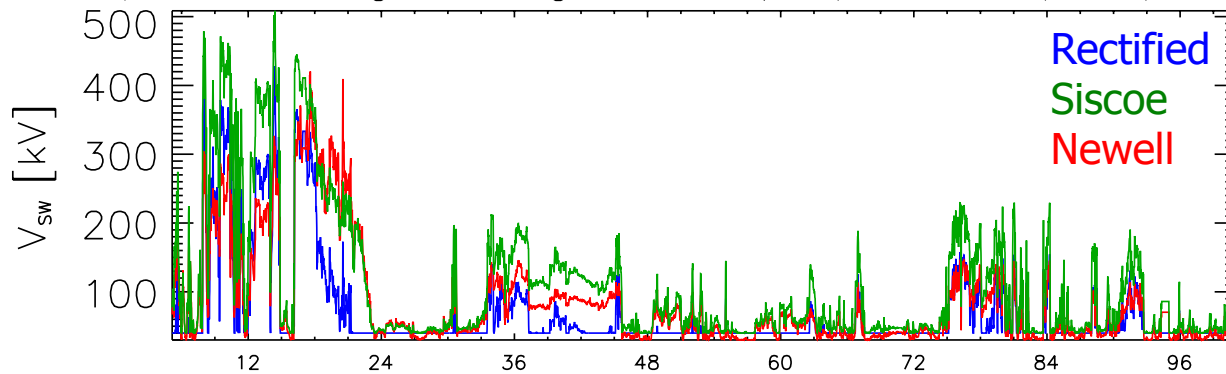
WINDMI Results (09/09–09/13/2011)



17-19 September 2011 Dst min. -58 nT, AL min. -1200 nT

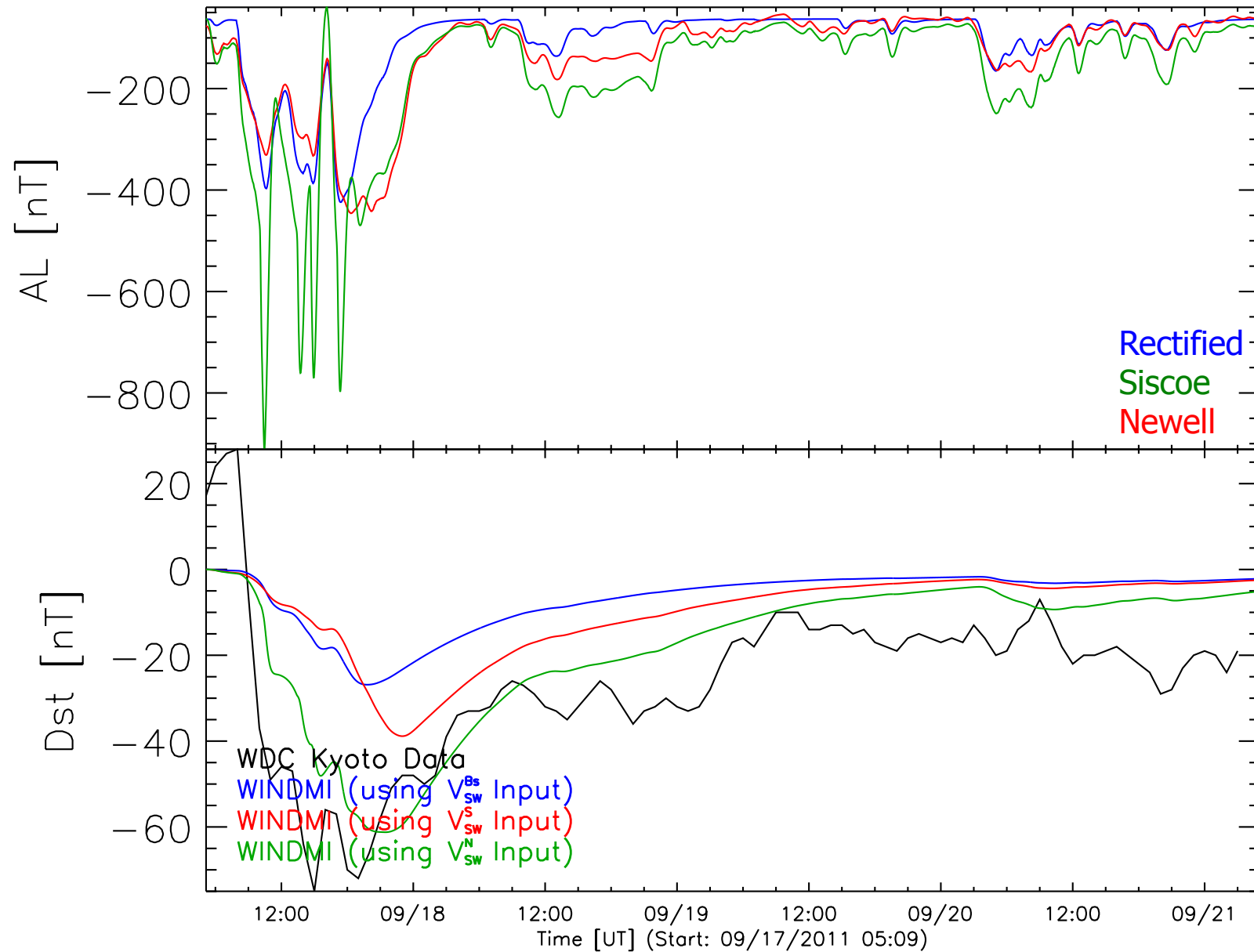


Input Driving Voltage V_{sw} (09/17-09/21/2011)

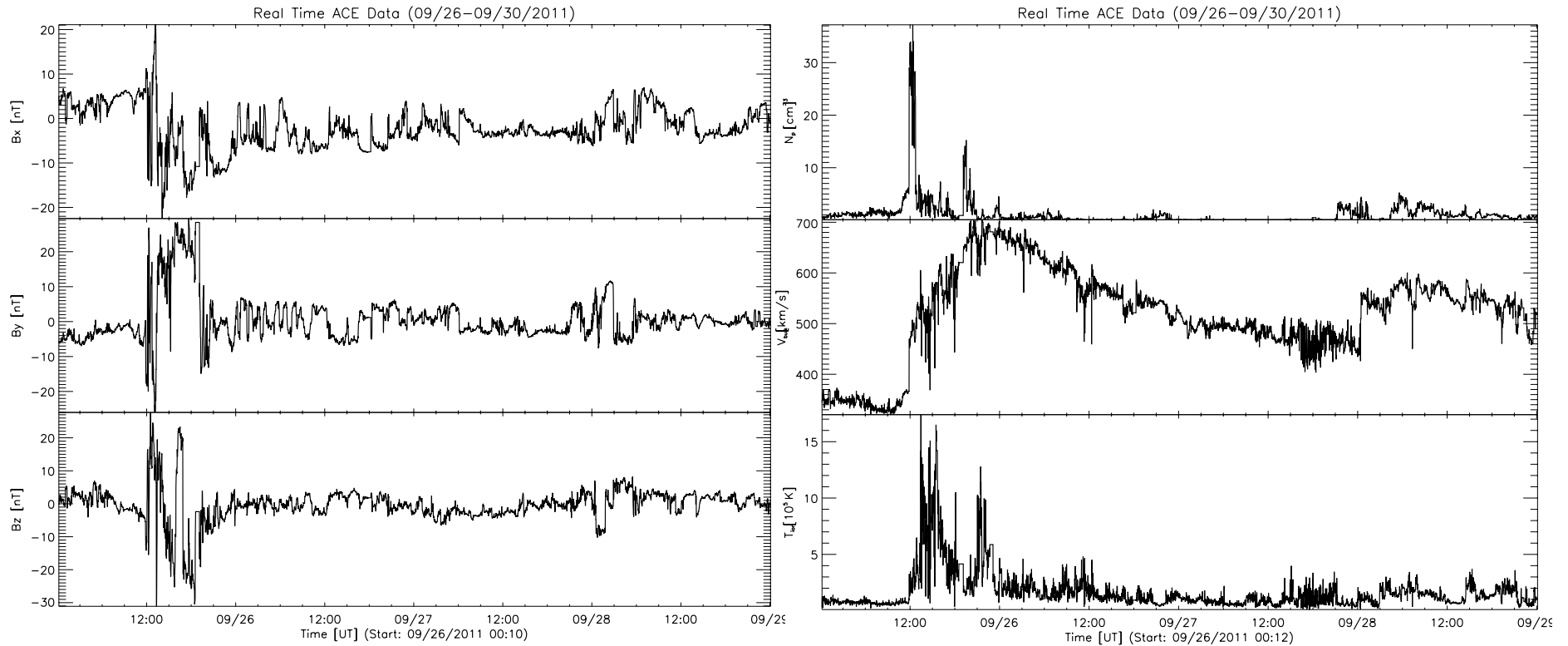


17-19 September 2011 Dst min. -58 nT, AL min. -1200 nT

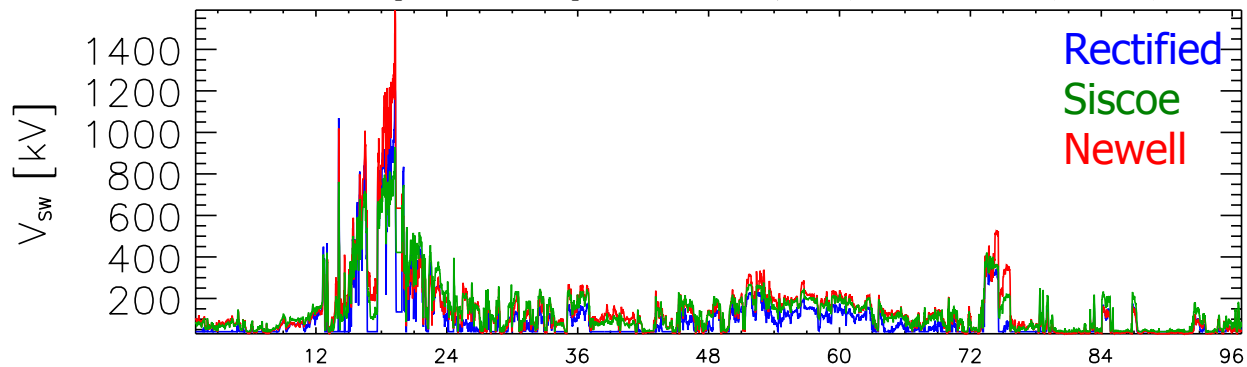
WINDMI Results (09/17-09/21/2011)



26-30 September 2011 Dst min. -103 nT, AL min. -1600 nT

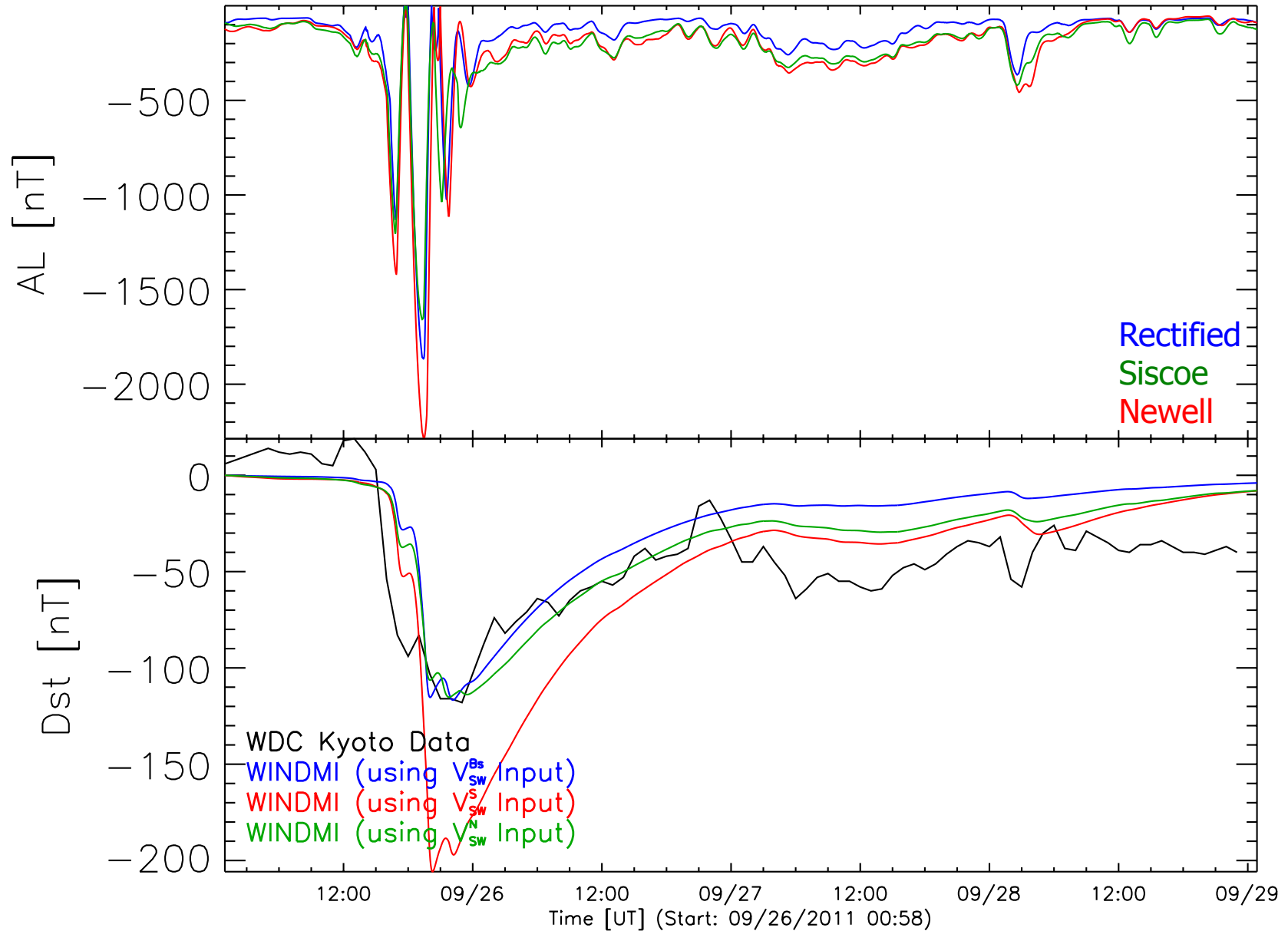


Input Driving Voltage V_{sw} (09/26-09/30/2011)



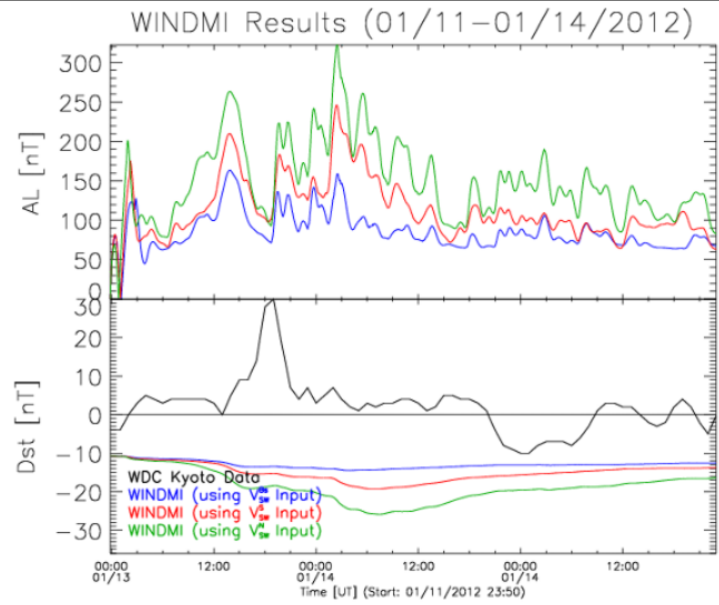
26-30 September 2011 Dst min. -103 nT, AL min. -1600 nT

WINDMI Results (09/26-09/30/2011)



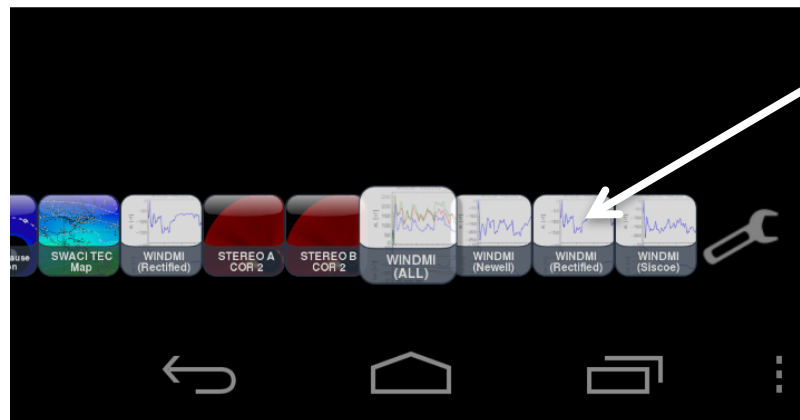
Screen
Shot
from
Smart-
Phone
11-14 Jan
2012

WINDMI - Newell, Siscoe, & Rectified
Solar Wind Coupling Functions/Drivers
2012-01-15 00:56:00.0



iphone store
Apps NASA
Space Weather

Look at
Cygnets for
WINDMI



Conclusions: WINDMI

- WINDMI is a fast, reliable tool for real-time forecasting of Space Weather: Storms and Substorms
- Download runs at CCMC.gsfc.nasa.gov under tab “Instant Runs” and as APP [iTunes NASA Space Weather].
- Algorithms for the sum of the contributions multiple current loops +ring current yield 3 phases of $Dst(t)$ for large storms. Validation pubs in JGR.
- Auroal magnetic index $AL(t)$ reproduces historical substorm data with low Average Rel Variance. (isolated substorms and substorms-within-storms)
- Support of L. Rasttaeller and CCMC team make it work!

