

UNH-CCMC Collaboration

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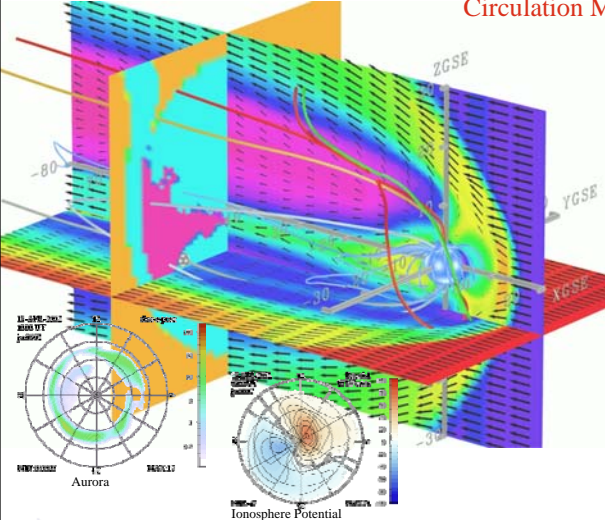
CCMC Workshop, Key Largo, FL, January 25, 2010

Overview

- What is OpenGGCM, where does it come from?
- OpenGGCM at CCMC.
- OpenGGCM use:
 - Tail instability (Siscoe et al.)
 - Ionosphere currents (Vennerstroem et al.)
 - THEMIS support.
 - Metrics.
 - Other science.
- Development under LWS / Strategic Capabilities.
- Development under NSF/PetaApps.
- Other Developments.
- Future Releases (v4.0).

OpenGGCM: Global Magnetosphere Model

The Open Geospace General Circulation Model:



- Coupled global magnetosphere - ionosphere - thermosphere model.
- 3d Magnetohydrodynamic magnetosphere model.
- Coupled with NOAA/SEC 3d dynamic/chemistry ionosphere - thermosphere model (CTIM).
- Coupled with inner magnetosphere / ring current models: Rice U. RCM, NASA/GSFC CRCM.
- Model runs on demand (>300 so far) provided at the Community Coordinated Modeling Center (CCMC at NASA/GSFC).
<http://ccmc.gsfc.nasa.gov>
- Fully parallelized code, real-time capable. Runs on IBM/datastar, IA32/164 based clusters, PS3 clusters, and other hardware.
- Used for basic research, numerical experiments, hypothesis testing, data analysis support, NASA/THEMIS mission support, mission planning, space weather studies, and Numerical Space Weather Forecasting in the future.
- Funding from NASA/LWS, NASA/TR&T, NSF/GEM, NSF/ITR, NSF/PetaApps, AF/MURI programs.

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Heritage

- (early 80's) First global magnetosphere MHD models: LeBouef, Ogino @UCLA, Lyon, Brecht, Fedder @NRL.
- (~1993) First parallelized global MHD magnetosphere model (UCLA-MHD).
- (~2000) UCLA global MHD model + NOAA/CTIM model ==> OpenGGCM.
- (2002) Second model to be implemented at the CCMC. >300(?) runs on demand at CCMC to date.
- (2005) Start RC coupling with RCM (NASA SR&T).
- (2006) NASA/NSF "Strategic Capabilities" funding for CTIPe/RB/RCM/CRCM coupling and V&V.
- Current uses: mission planning (THEMIS, Swarm, Mag-Con), complement data analysis, study fundamental processes, numerical experiments.
- >30 data comparison studies in the refereed literature since 1995: tail physics, magnetopause, ionosphere, ground mags ...
- Real-time capable with modest resources.

Example of OpenGGCM use: Investigation of tail instability by George Siscoe

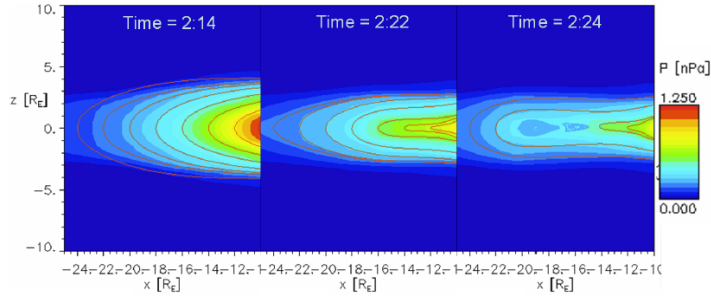


Fig. 1. Three frames from the output of CCMC run UCLA2_UCLA2.060105_3 showing magnetic field lines and contours of plasma pressure in the midnight meridian plane from $10 R_E$ to $25 R_E$ down the tail. The contours and field lines are changing in time in response to an IMF flip from south to north, which reached the magnetopause at time 2:06. Magnetic reconnection in this plane commenced between 2:22 and 2:24. In the last frame, blue blobs are coiled-up plasmoid field lines. Solar wind conditions for the run are density = 5 n/cc, speed = 400 km/s, temperature = 100 000 K, and field strength = 5 nT.

Siscoe et al., *Annales Geophys.*, 27, 3141, 2009

Example of OpenGGCM use: Investigation of tail instability by George Siscoe

Imbalance of $\text{grad}(p)$ and $J \times B$ forces occurs before substorm onset. Points to an ideal MHD instability such as ballooning.

Siscoe et al., *Annales Geophys.*, 27, 3141, 2009

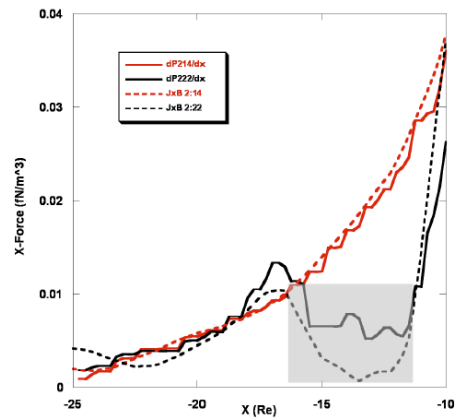
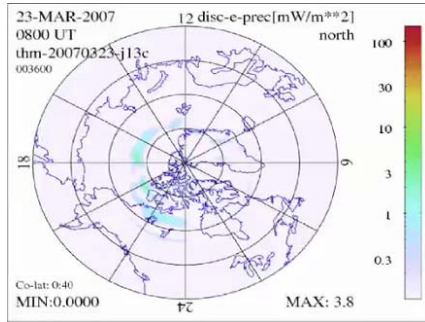
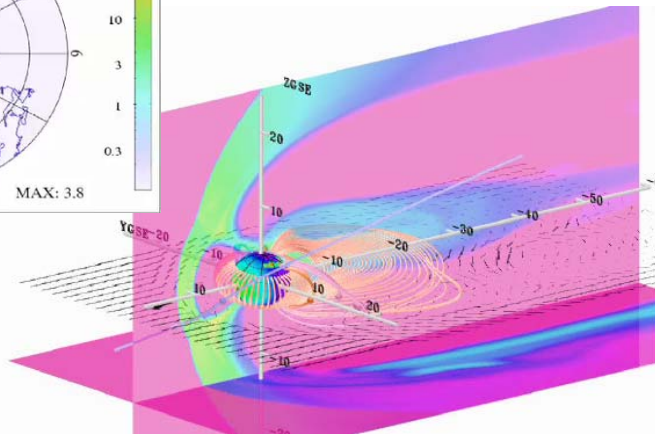


Fig. 2. Corresponding to Fig. 1, this figure shows profiles along the tail axis of the magnetic force ($J \times B$) and the pressure-gradient force (with signed changed for easier comparison) nine minutes prior to reconnection onset (red) and one minute prior to reconnection onset (black). In the gray region the pressure gradient force, which is pushing the plasma away from the Earth, is greater than the magnetic force pulling the plasma toward the Earth.

Follow-up with study of observed THEMIS substorm of March 23, 2007



Aurora and
Westward
Traveling Surge



Force balance breakdown before tail reconnection onset

Example of OpenGGCM use: Field-aligned currents by Susanne Vennerstrom

How do FACs react
to changes in the
IMF direction?

Vennerstrom et al.,
JGR, 110, A06205,
2005

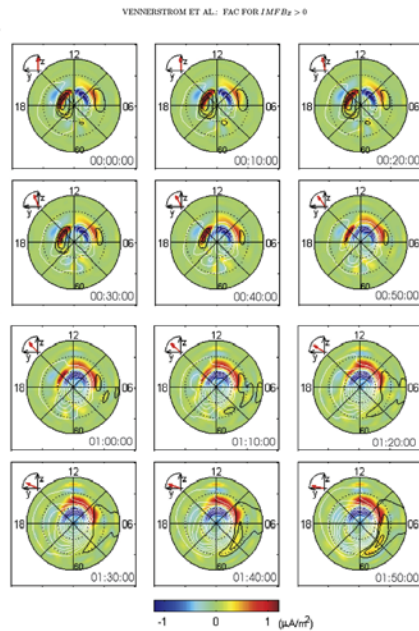


Figure 2. Development of the simulated FAC-patterns (color-coded) and electric potential (contour-lines) in the northern polar region when the IMF (shown in the upper left corner of each panel) rotates from northward to eastward. Upward FAC's are blue and downward FAC's are red. White contour lines signify negative potential and black contour lines positive potential. Contourline spacing is 10 kV.

THEMIS Support

- THEMIS has an OpenGCM modeling component.
- THEMIS researchers bring problems forward and we run the model.
- Example: Plasma entry into magnetosphere under northward IMF:

Oieroset et al., GRL, L17511, 2009:

Li et al., JGR, 114, A00C15, 2009:

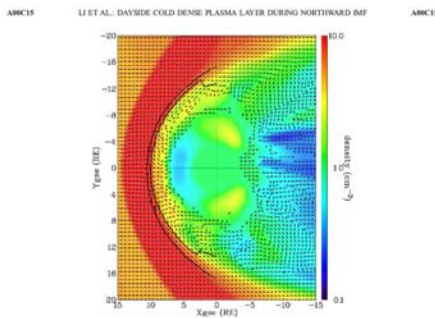


Figure 9. The plasma density and flow velocity field projected on GSE equatorial plane at time 16:30 UT in the simulation. The vectors are normalized to unit vectors to represent the flow directions projected on the equatorial plane. Vectors with velocity magnitude on the plane less than 5 km/s are not shown. The outer black line shows the boundary between the IMF and the open field, while the thick black line further earthward shows the boundary between the dayside closed field and the open field.

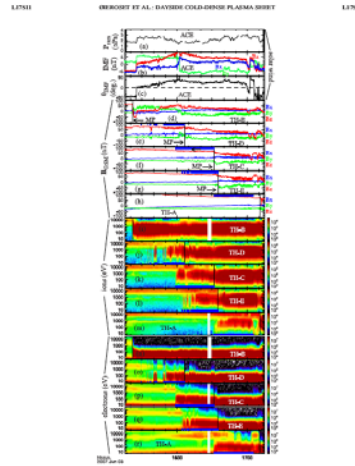


Figure 4. (a)-(c) ACE solar wind parameters, magnetic field, and magnetic field beta angle $\beta_{\text{mag}} = \tan^{-1} \sqrt{B_x^2 + B_y^2} / B_z$. (d)-(f) magnetic field components, (g)-(i) ion magnetograms, and (j)-(l) electron magnetograms from THEMIS. D, C, E, A respectively. Horizontal dark lines mark for CDF's intervals. The magnetograms marked by a vertical black line correspond to the magnetized field boundary defined from the data in this study distribution.

Metrics

- We participate in CCMC's metrics and evaluation efforts.
- Antti will say more....

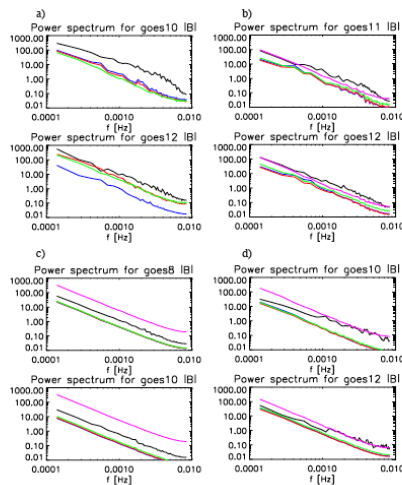
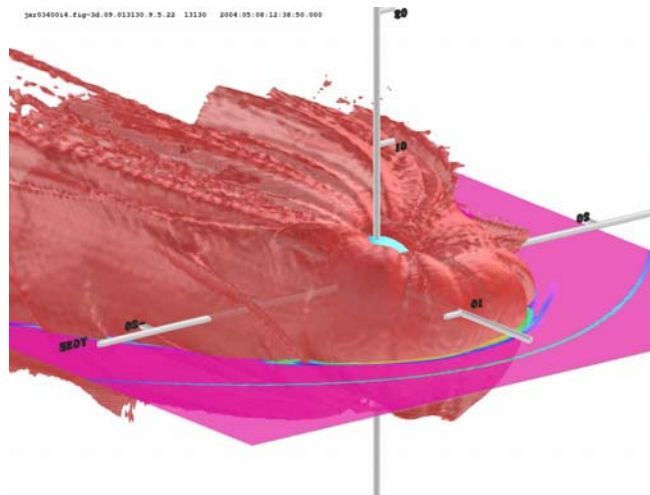


Figure 6. Spectra of the modeled and the observed geostationary magnetic field fluctuations in Fig. 5 for storm events (panels a-d corresponding to events 1-4) given in Table 1. Different colors indicate the power associated with different models. Black curves indicate the observed spectra. The color codes corresponding to different models are given in Table 3.

Lpulkkinen et al., JGR, in press, 2010:

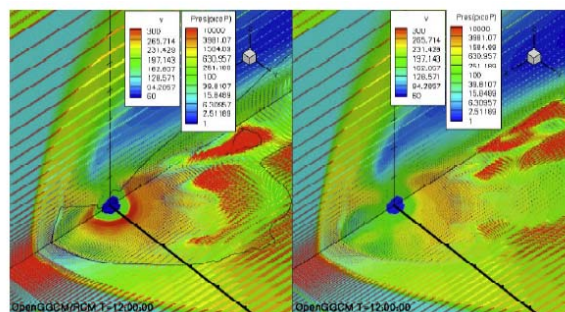
Other science: FTEs

- Flux Transfer Events during east-west IMF, May 8, 2004 Cluster event.
- FTEs for in subsolar region, only move towards dawn, observed at dawn MP.
- OpenGGCM predicts dawn-dusk asymmetry depending on IMF By sign and tilt.

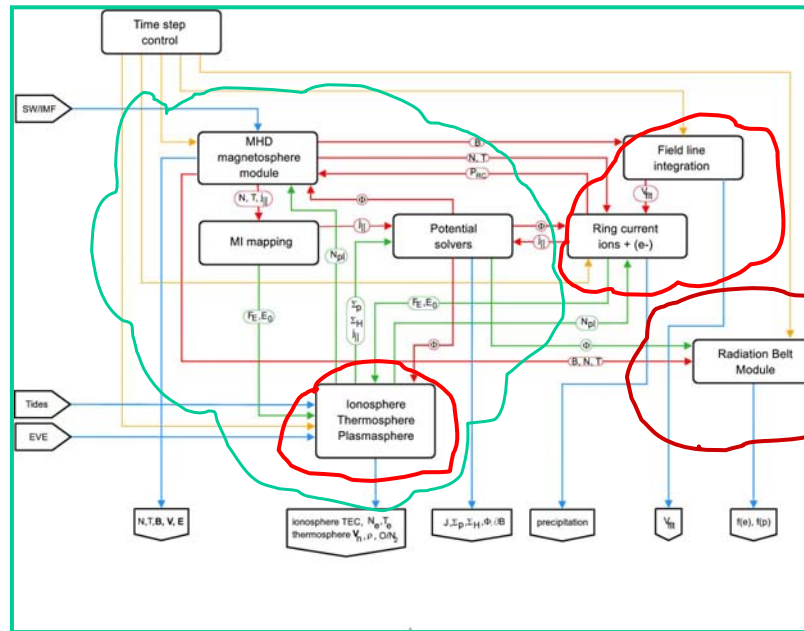


OpenGGCM development under LWS/SC

- Collaboration with:
 - NASA/GSFC (M.-C. Fok, A. Gloer): CRCM/RBM coupling.
 - Rice U. (F. Toffoletto, B. Hu, A. Chan): RCM coupling, RB diffusion.
 - U. Colorado (T. Fuller-Rowell, N. Maruyama): CTIPe/GT-GIP coupling.
 - NCAR/HAO (A. Richmond, A. Maute): Asymmetric potential solver.
- Runs through 09/2011, just passed mid-term review.
- Accomplishments:
 - RCM coupling basically finished.
 - CRCM coupling basically finished.
 - Potential solver tested.
 -



Model Data Flow



Development under PetaApps: Tuning for CBE

Computing: The Old Days...

15 years ago one would have needed something like a Cray Y/MP:

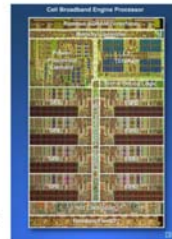
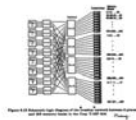
- 8 processors
- ~ 1 GB memory
- ~ 8 x 0.2 Gflops
- ~ \$10M
- ~ 200 kW power/cooling
- ~ \$1M/y operations.

First global models ran on such machines.



Computing: Enter the Cell Chip

1992 Y/MP block diagram:



2006 Cell Broadband Engine (CBE) chip (IBM/Sony/Toshiba):

Computing: Cray on a Chip



Cell chip versus Y/MP processor board (2 ft wide)



Computing: Comparing the Vitals

	Y/MP	Cell chip
Processors:	8	1+8(6)
Memory:	0.25-1.0 GB	0.25 GB
Peak speed:	~ 2 Gflops	~ 256 Gflops
Memory bw:	2 GB/s	25.6 GB/s
Clock:	0.333 GHz	3.2 GHz
Power:	100 kW	200 W
OS:	UNICOS	Linux
Price:	\$10M	599.00 @ Walmart
Compilers:	Crap back then	Crap now

Development under PetaApps: 40 PS3 Cluster



- 40 PS3 from Best Buy + GB Ethernet switch + PC head node + cables + monitor – games ~\$24k.
- New firmware, Linux, MPI libs etc.
- Uses 5 kW of power, though.
- Motivates middle-schoolers, newspaper writers.

Page A2 • NEW HAMPSHIRE FROM LEISURE, Monday, June 23, 2008

UNH's supercomputer to predict 'space weather'

Advanced math: By combining 40 PS3 gaming consoles, the institute created a computer that, in theory, can perform 8 trillion calculations per second.

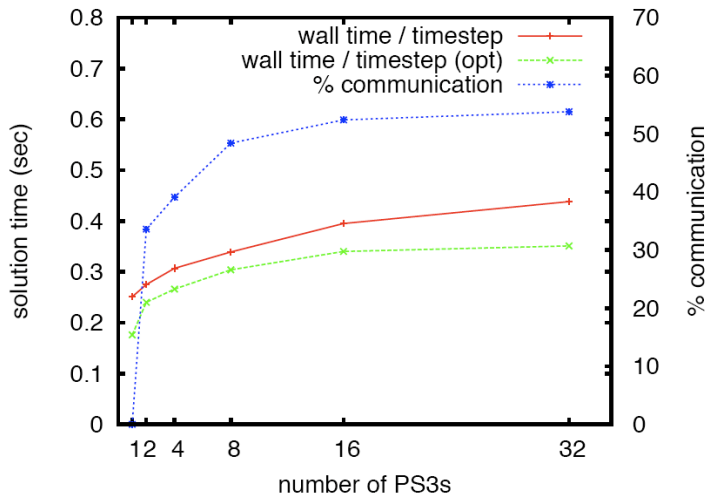
BY CLYTON NANGO
From our computer
news column

UNH — Video game
nerds realize themselves at
the University of New Hamp-
shire have hatched an idea: to
use their consoles to form a
supercomputer that will use to
help predict "space weather."

UNH's supercomputer is
made up of 40 PS3 consoles
connected to a central PC
head node. The system is
designed to perform 8 trillion
calculations per second. The
UNH supercomputer is being
used to help predict "space
weather," which is a term
used to describe the solar
wind and other phenomena
that can affect Earth's
magnetic field and atmosphere.

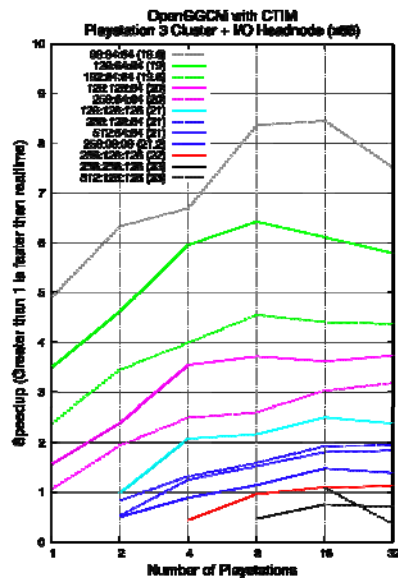
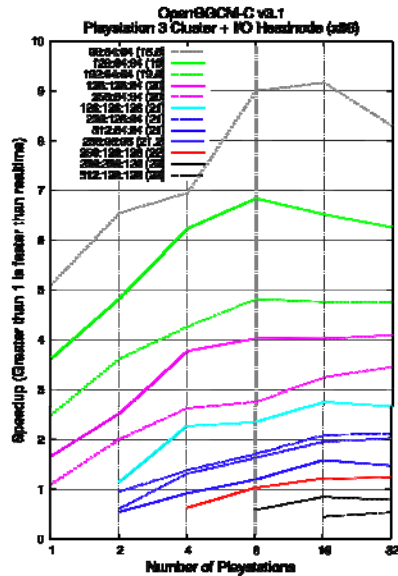
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Scaling with number of PS3's



Weak scaling on 40 PS3 cluster: Not perfect. Too much time spent on inter-node communication (GB Ethernet switch). New QS22 IBM Cell blade cluster should be much better. Cluster (42 nodes / 84 CBE, NSF CISE funding + IBM donation) delivered November 2009.

Scaling for PS3 cluster: Event of 31 August, 2001



Latest results: ~ factor 1.8 better!

Other developments

- Code now completely under subversion control.
- Bug reporting system: <http://fishercat.sr.unh.edu/trac/openggcm>
- Compiling with autoconf/make.
- Ultra-high resolution runs for ballooning (1200x600x600, Ping Zhu, U. Wisconsin).
- Include Hall physics (LWS grant, Bhattacharjee).
- Particle tracing / reconstruction of $f(v)$ (for now cusp ion structures, grad student project).
- Calculating proton aurora (grad student project).
- Moving dipole (grad student project).
- Version 4.0: not so soon. Still many loose ends that need to be tied up.