Current Users

- 1. Unifying theme was an appreciation of CCMC staff for their efforts to assist (All).
- 2. Magnetospheric models powerful tools providing results that immediately translate into discoveries and research papers (Siscoe).
- IT services are impressive. Can compare with observations to determine what features models have to work on more. (Mannucci)

Current Users

• 4. Suggestions

- a. Continue your efforts to couple models, e.g. magnetosphere/ionosphere to thermosphere (Siscoe, Mannucci)
- b. Need for enhanced user-provider contacts to enhance flexibility and user control in solar/heliospheric models to the levels available for magnetospheric models (Luhmann, Devore-A, Mannucci).
- c. Continue STEREO support (Luhmann, Devore-A)

Current Users

- 4. Suggestions (con)
 - d. Provide read me files to accompany output downloads (Devore-A)
 - e. Integrate model coronal densities for comparison with coronograph images (Devore-A)

- Dikpati- solar dynamo models. Flux transport dynamo: generation of toroidal field by shearing poloidal field, regeneration of poloidal field by lifting and twisting a toroidal flux tube by helical turbulence. Flux transport by meridional circulation.
- Predicts surface radial and toroidal fields for comparison with observations of photospheric fields.
- Predict peak amplitudes of solar cycles 12 through 24 with flux transport dynamo model. Based on memory of magnetic field.
- Trying to predict Maunder minimum.

2. George Fisher (given by Janet Luhmann)

- ANMHD solve 3D MHD equations to study subsurface evolution, convective dynamo, magnetoconvection, active region formation
- FLCT uses local correlation tracking to determine spatial displacements of image structure.
- AMPS. 3D MHD equations to study solar magnetic field from convection zone to corona.
- Developing SANMHD. 3D MHD in spherical geometry.

• 3. Solomon

- Models. TIE-GCM, TIME-GCM, WACCM big models.
- AMIE, GLOW, CMIT, GSWM, CAM3, WRF
- TIE-GCM. Thermosphere ionosphere. Solar flux, solar energy deposition, auroral forcing, tidal oscillations, code engineering.
- TIME-GCM Includes mesosphere. Flux couple to lower atmosphere.
 - WACCM. TIME GCM + MOZART OZONE, CAM3 community atmosphere. Presently available through portal at NCAR, but not with a user-supplier mentality

- 4. **DeVore**. ARMS for the CCMC (not a call for defensive measures)
 - Time dependent MHD equations in 3D in dynamic block adaptive grids. Heliospace display package.
 - Grid: cartesian, cylindrical spherical. Boundaries: periodic, symmetric. Grid adaptive.
 - Dynamic remeshing.
- Transitioning. Ready to provide code to CCMC, but this is a complex research tool, run on demand or pre-operational model. Expert assistance needd. Would have to constrain user inputs. Provide sample runs. Offer steady state inner coronal models to others.

• 5. **Zank**

- Heliopause, Termination Shock, Bow Shock, neutral Hydrogen wall
 - Must include neutrals, charge-exchange interactions, a mechanism for anomalous heat transport.
- Self-consistent models;
- Dynamic plasma and neutrals: four-fluid model
- Monte-Carlo Boltzmann, particle Boltzmann (kinetic)
- Kinetic and multi-fluid approaches are converging.
- Solar wind anisotropy, 2D vs 3D heliosphere, tilted dipole, wavy current sheet. Time dependence solar cycle, solar wind disturbances. Galactic cosmic rays.
- 2D codes ready for CCMC. Boltzman and Kinetic treatments.