

National Aeronautics and Space Administration

DSCOVR Solar Wind Observations

Adam Szabo, Andriy Koval
Justin Kasper, Michael Stevens



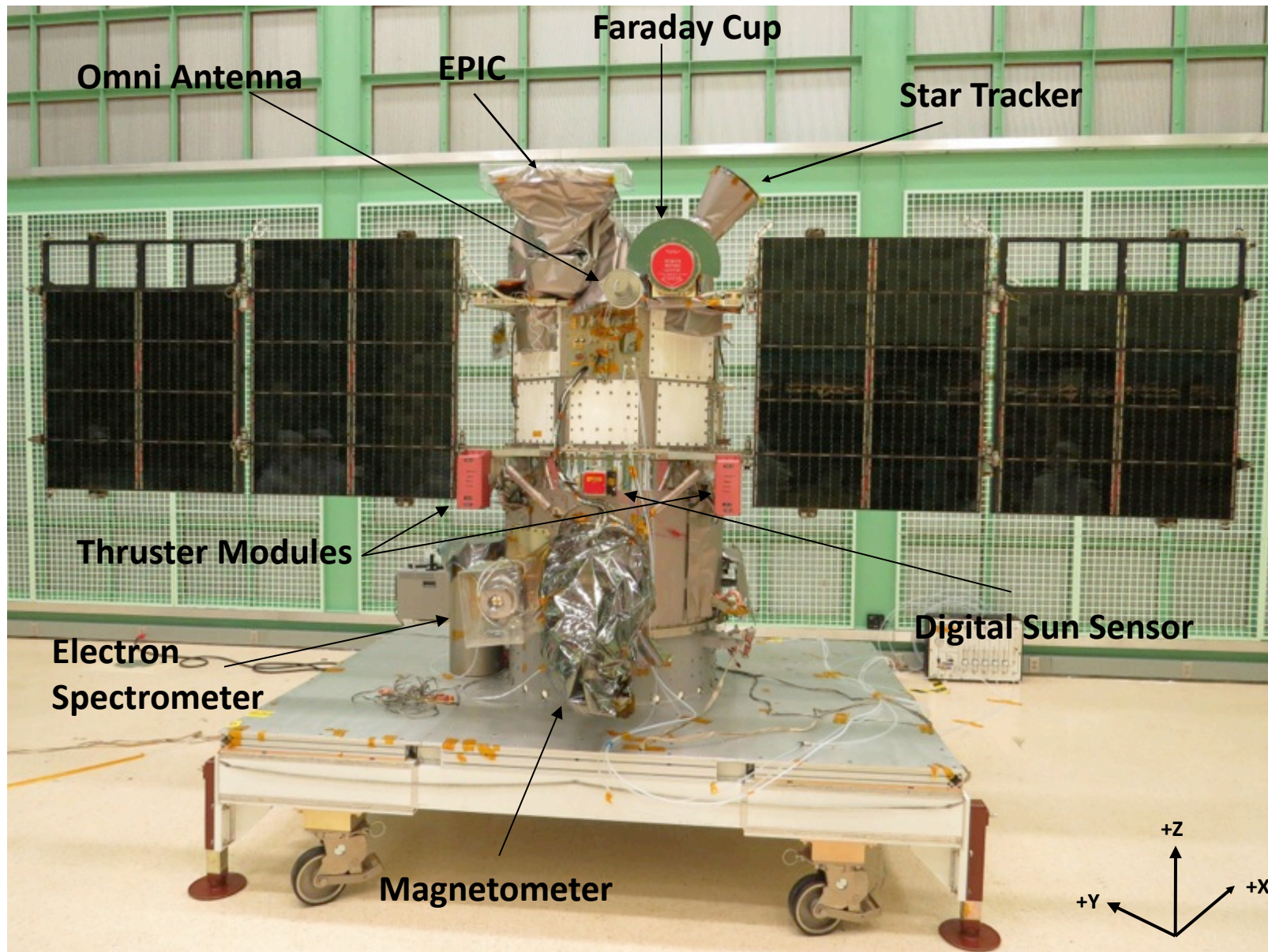
DSCOVR **DEEP SPACE CLIMATE OBSERVATORY**

advanced warning of approaching solar storms

<http://www.nesdis.noaa.gov/DSCOVR>



Locations of the Instruments



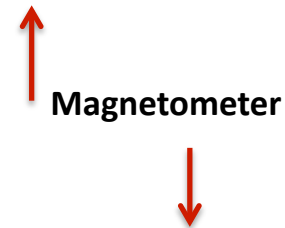
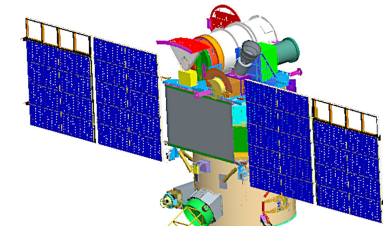


Goddard Fluxgate Magnetometer

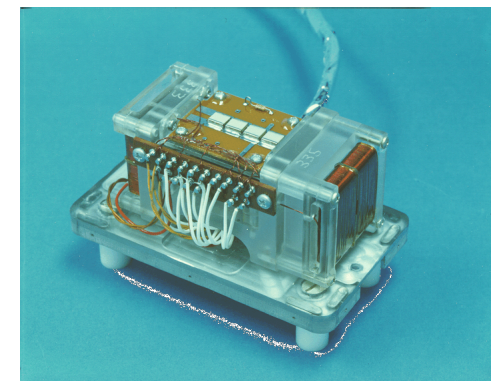
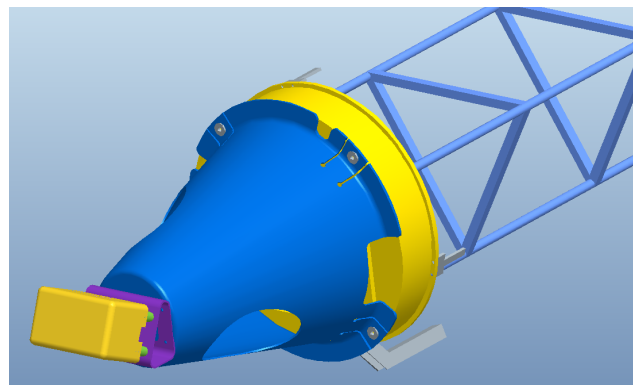


The Fluxgate Magnetometer measures the interplanetary vector magnetic field

It is located at the tip of a 4.0 m boom to minimize the effect of spacecraft fields

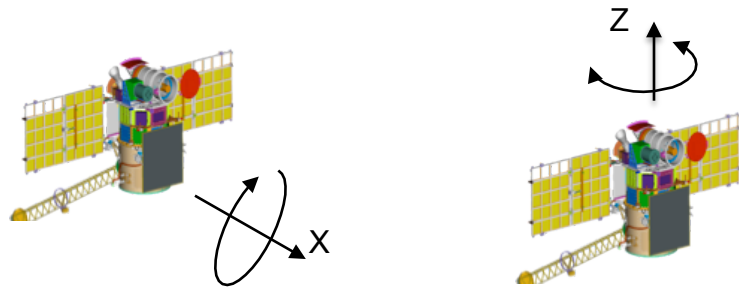


| Requirement | Value | Method | Performance |
|-------------|------------|----------|-----------------|
| Range | 0.1-100 nT | Test | 0.004-65,500 nT |
| Accuracy | +/- 1 nT | Measured | +/- 0.2 nT |
| Cadence | 1 min | Measured | 50 vector/sec |



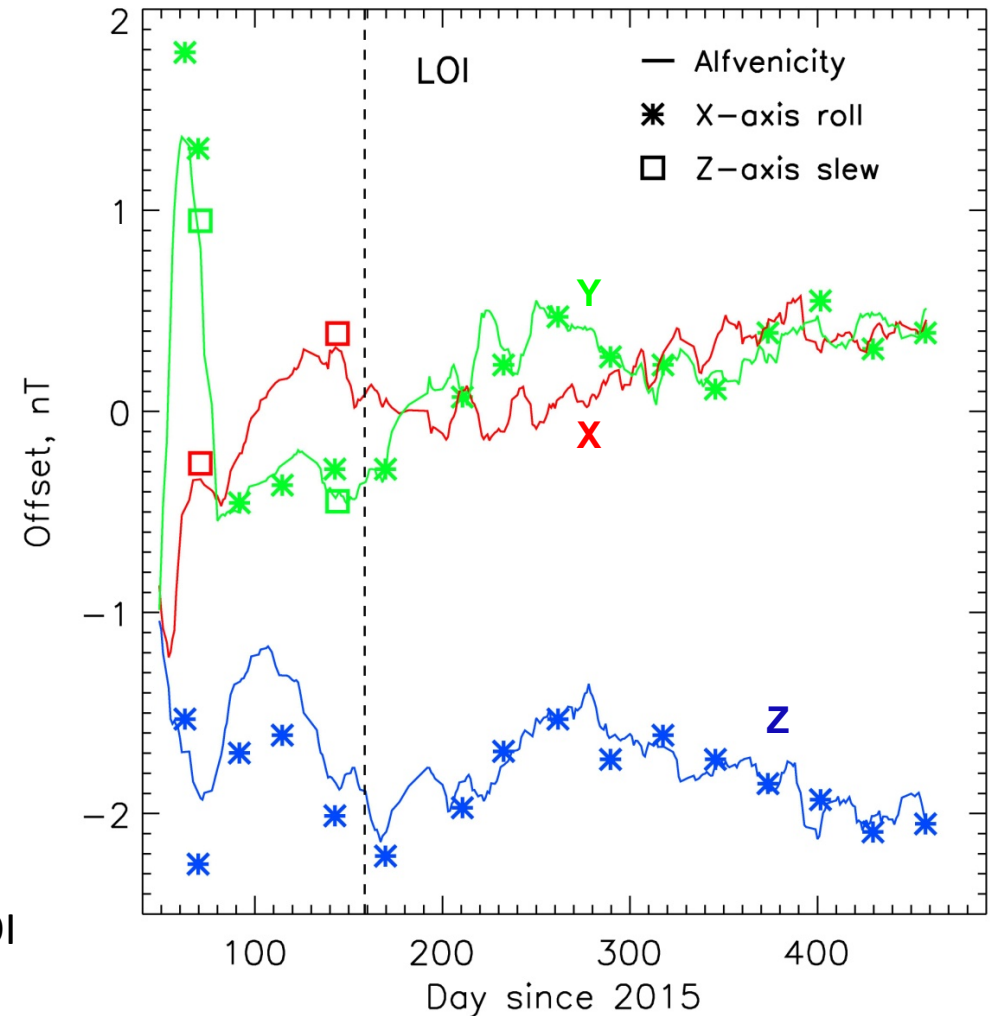


In-Flight Magnetometer Calibrations



- X axis Roll and Z axis Slew data is consistent with ground calibration estimates
- Independent zero offset determination by rolls, slews and using solar wind Alfvénicity give consistent values
- Time variation is consistent with yearly orbital change.
- Resulting magnetic field accuracy since LOI is ~ 0.2 nT, exceeding requirements.

Magnetometer Zero Offsets



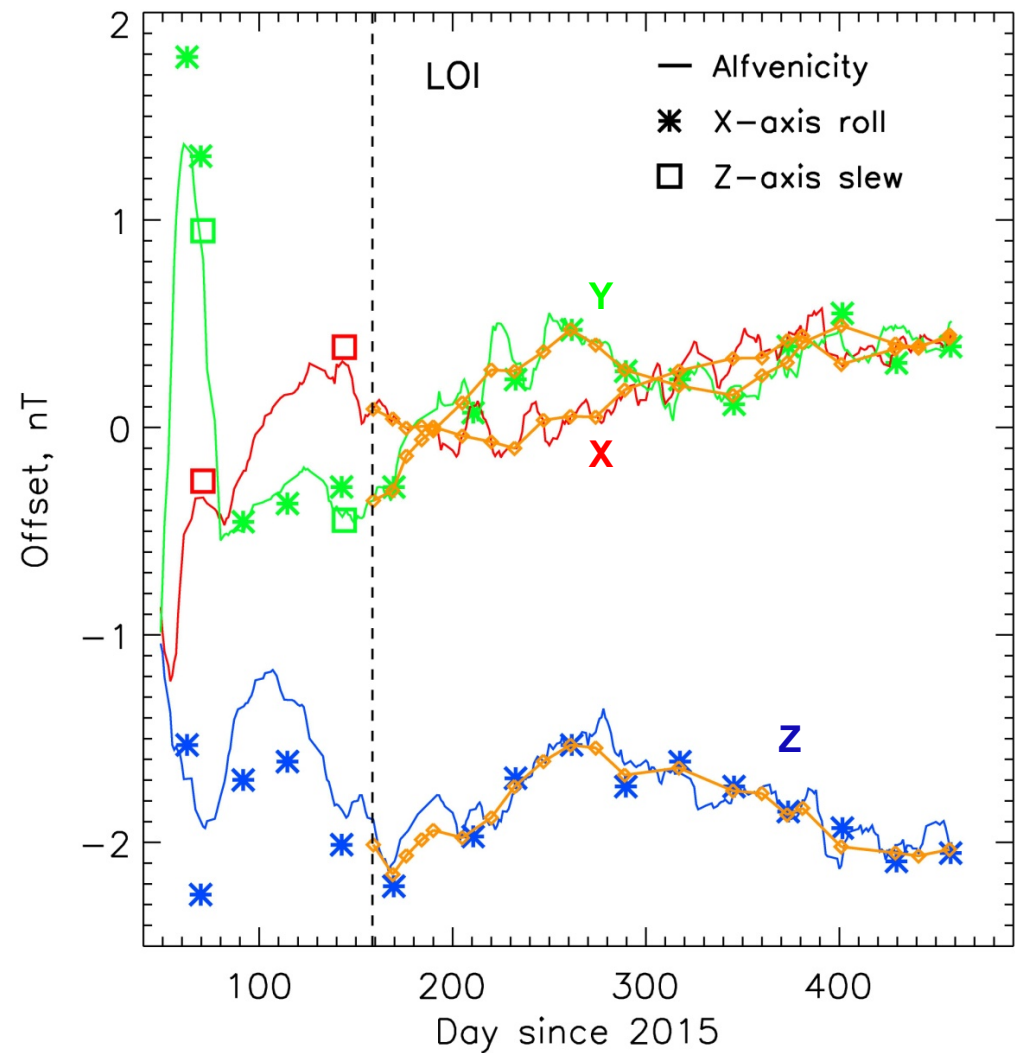


In-Flight Magnetometer Calibrations (2)



Magnetometer Zero Offsets

- Orange curve shows the offset values provided to NOAA SWPC.
- Updates are provided at least monthly or when sudden changes are identified.





Intercalibrations with ACE and Wind



DSCOVR, ACE and Wind Orbits

Require spacecraft separation <25 Re.

ACE – DSCOVR:

June 8 – 19, 2015

Sept 1 – 18, 2015

Nov 27 – Dec 13, 2015

Feb 24 – Mar 10, 2016

Wind – DSCOVR:

May 12, 2015

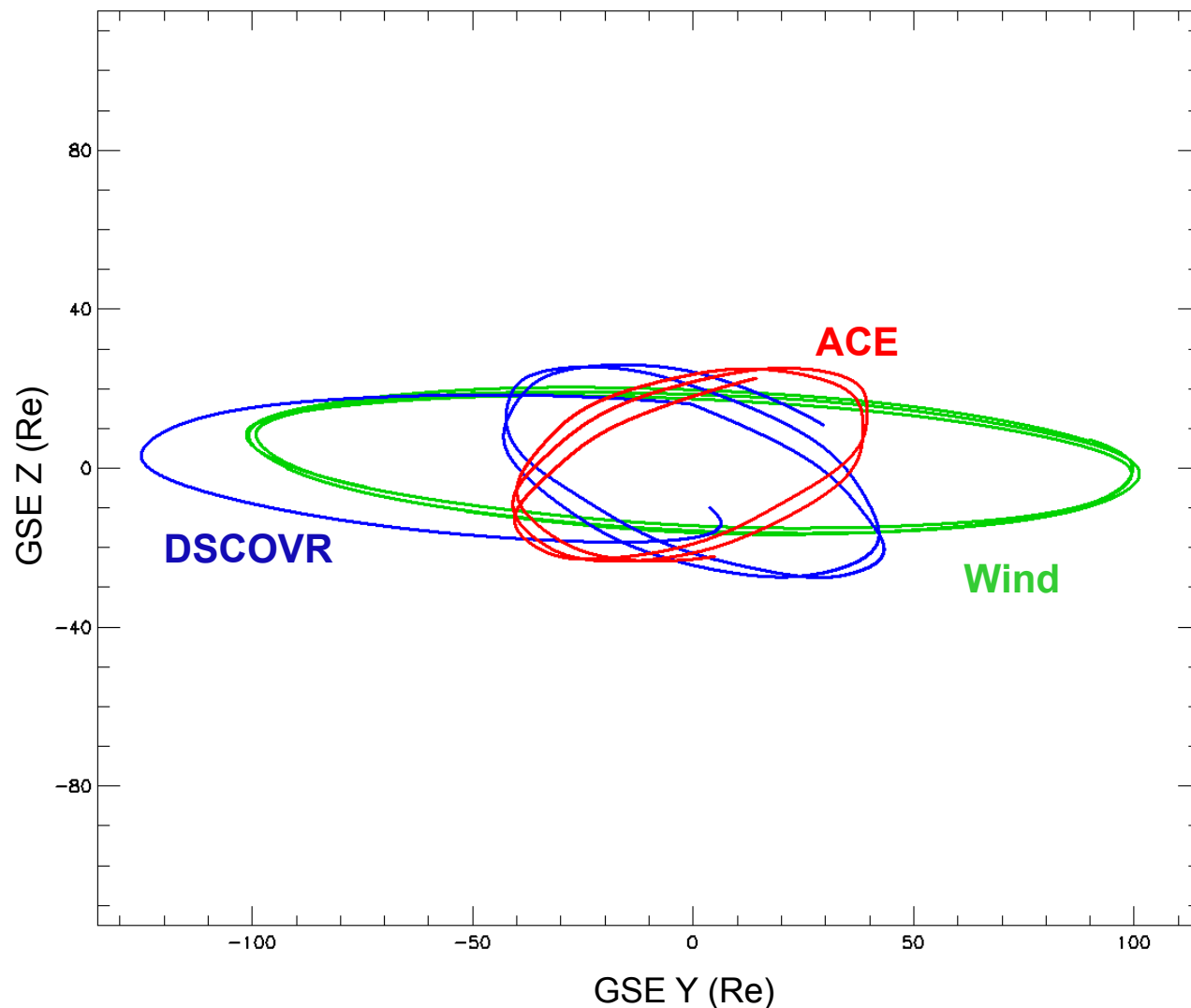
July 4 – 20, 2015

Oct 6 – 23, 2015

Dec 29, 2015 –

- Jan 3, 2016

Mar 23 – Apr 6, 2016

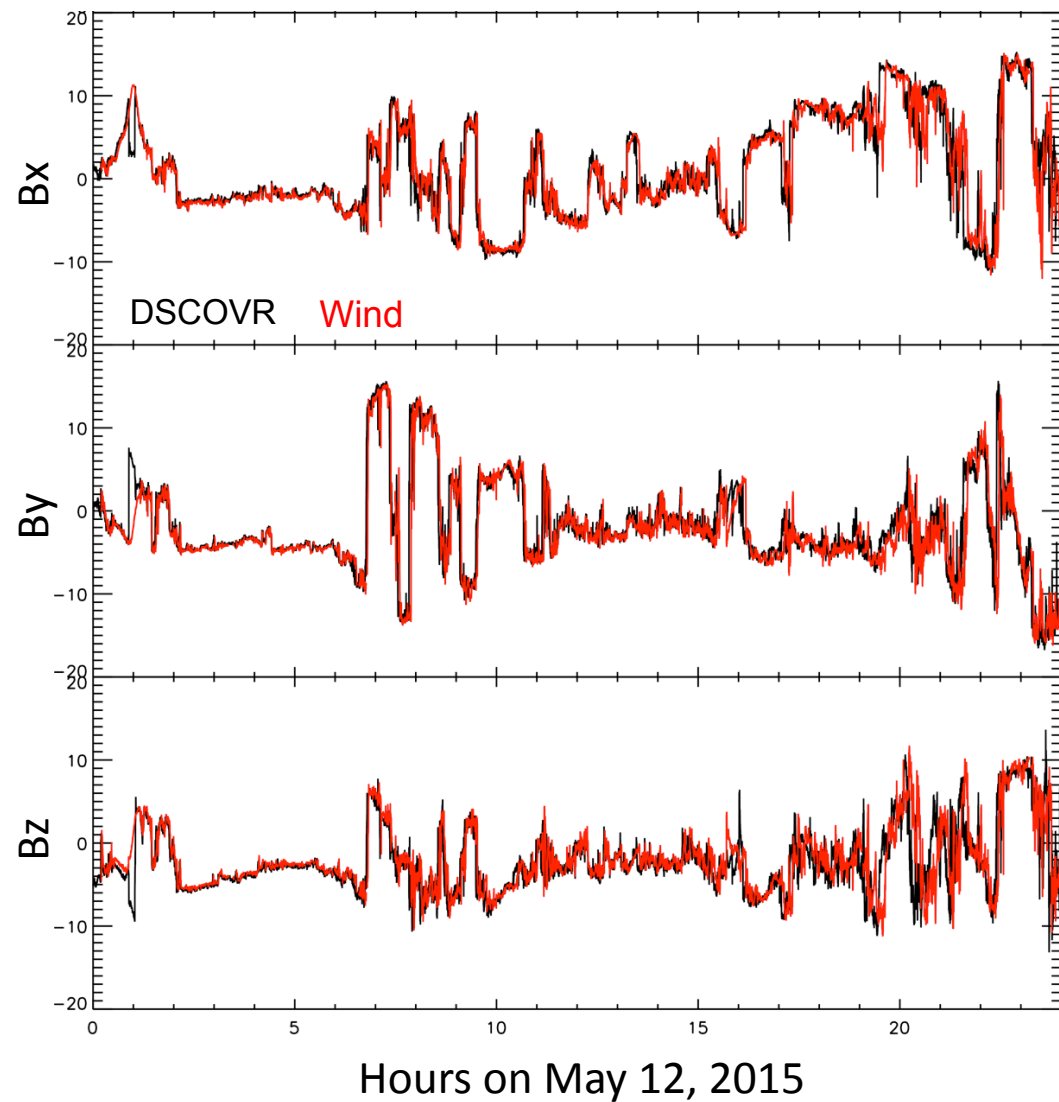




Wind-DSCOVR Comparison



Comparison with Wind spacecraft measurements show good agreement. DSCOVR data is in black. The time shifted Wind data (to allow for solar wind propagation) is plotted in red. Small deviations are consistent with spacecraft separation.



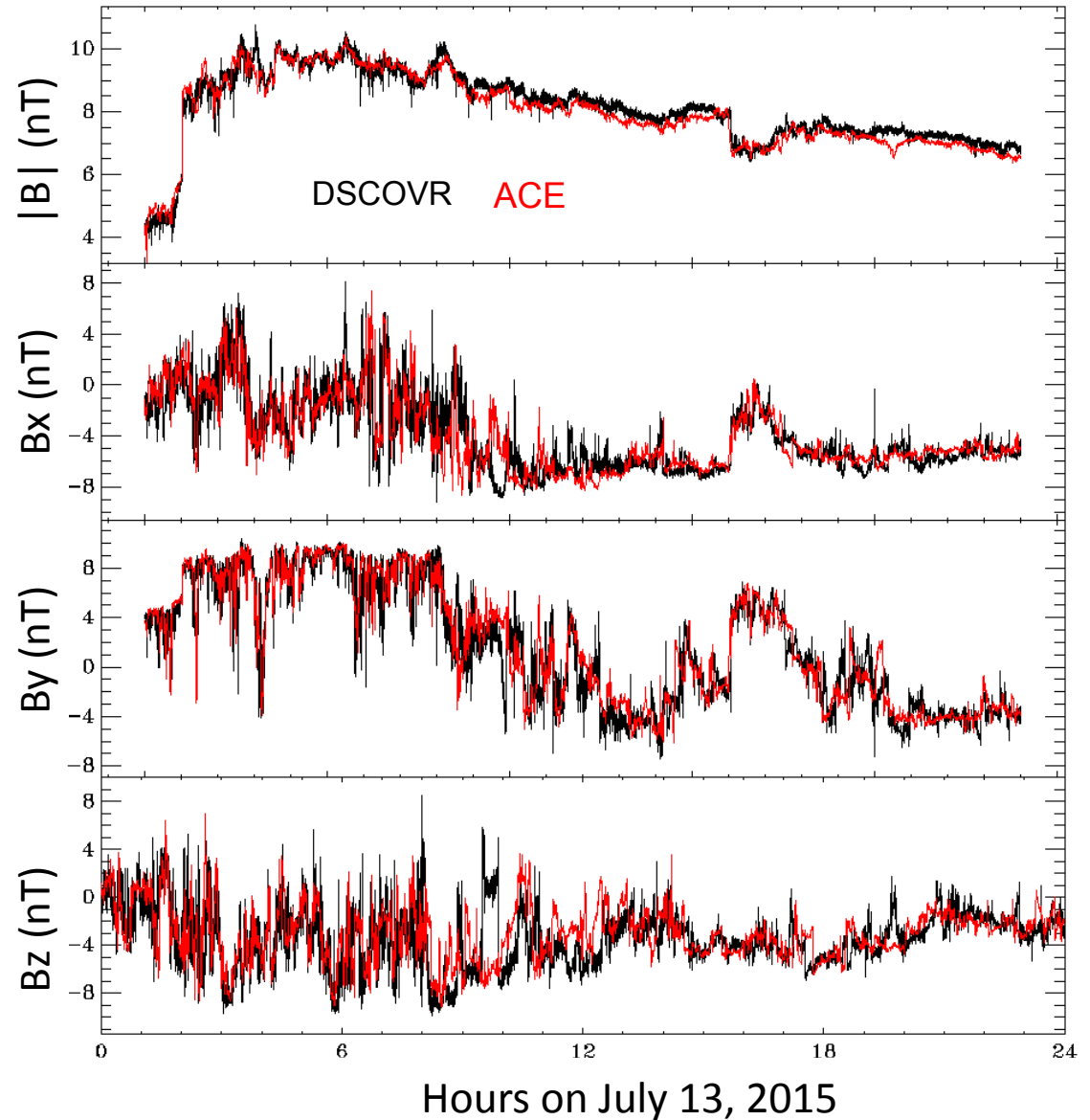


ACE-DSCOVR Comparison



Comparison with ACE spacecraft measurements also show good agreement. DSCOVR data is in black. The ACE data in red is not time shifted. Small deviations are consistent with spacecraft separation.

Interplanetary shock jump conditions at the beginning of the day agree as measured by the two spacecraft.

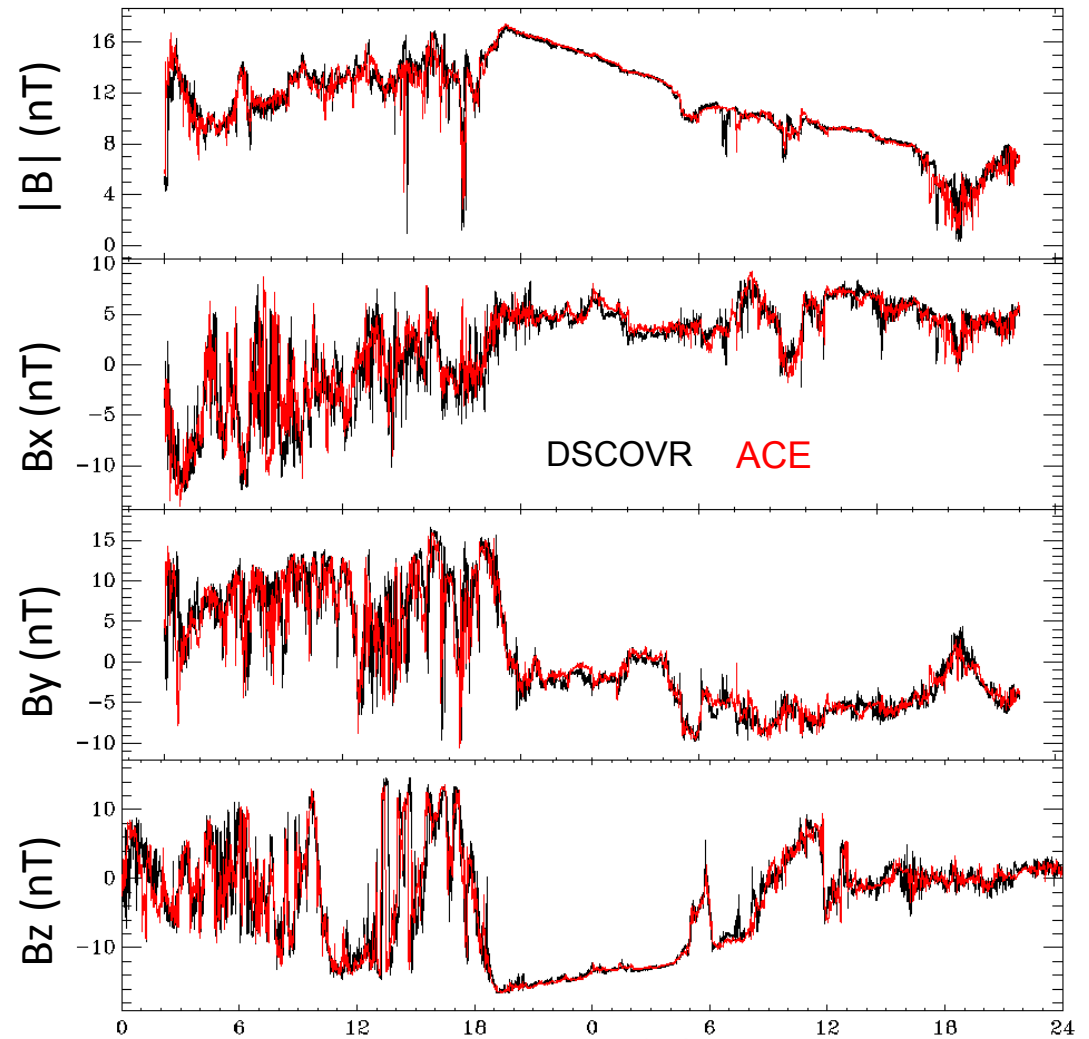




ACE-DSCOVR Comparison (2)



The 2015 Dec 31 – 2016 Jan 1 ICME was measured by both spacecraft with identical values.



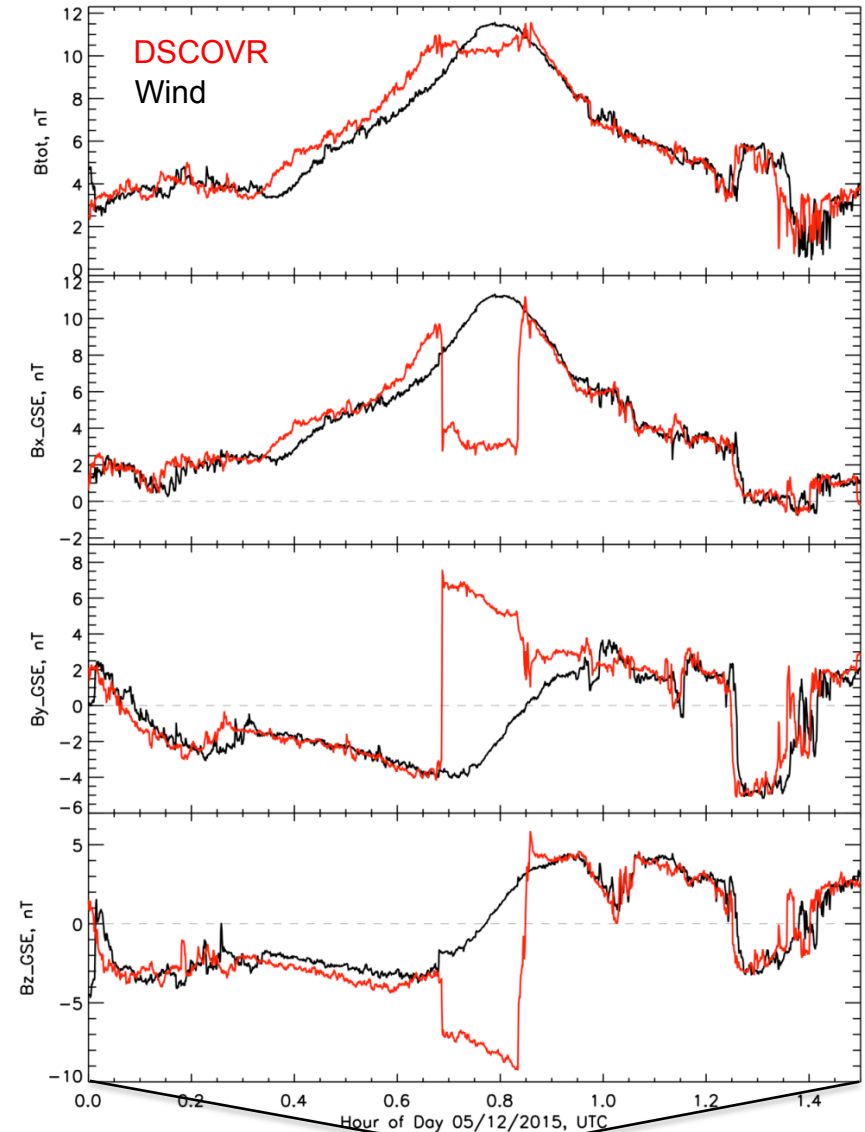
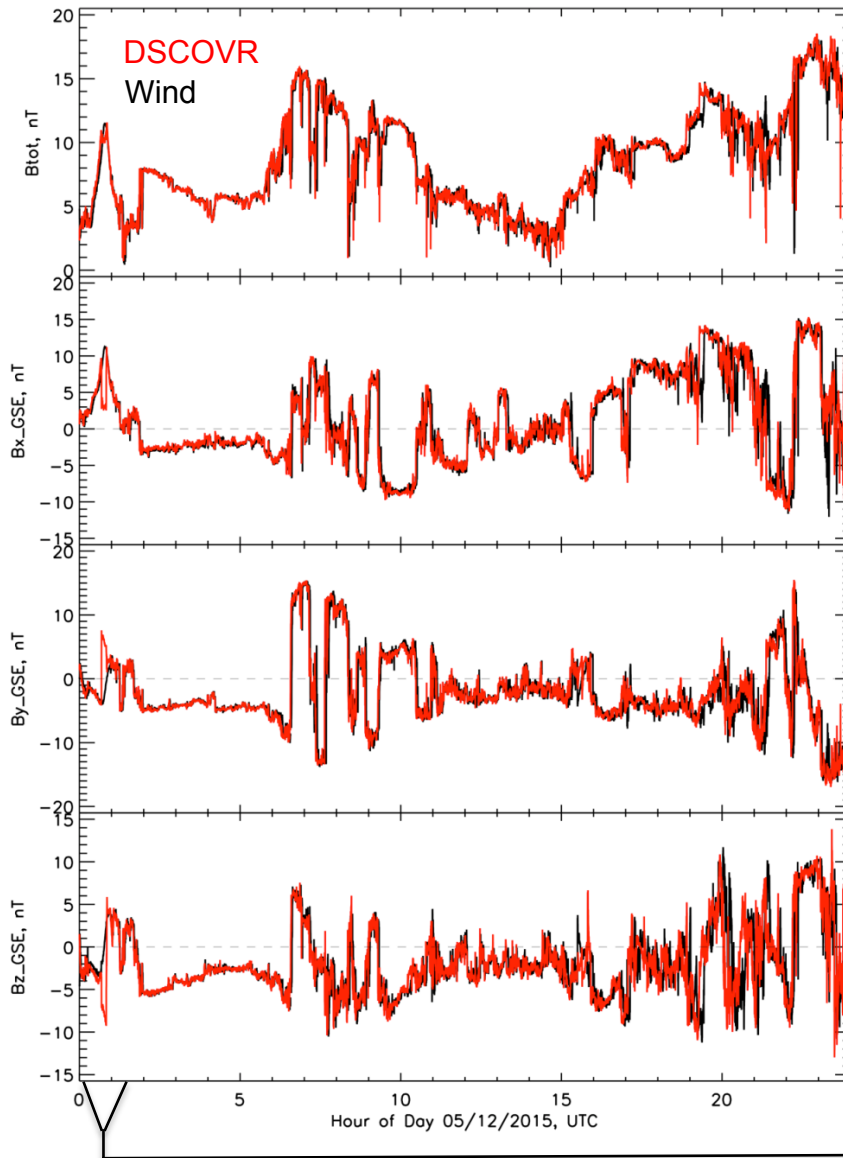
Hours on 2015/12/31 and 2016/1/1



DSCOVR Science: Small Structures

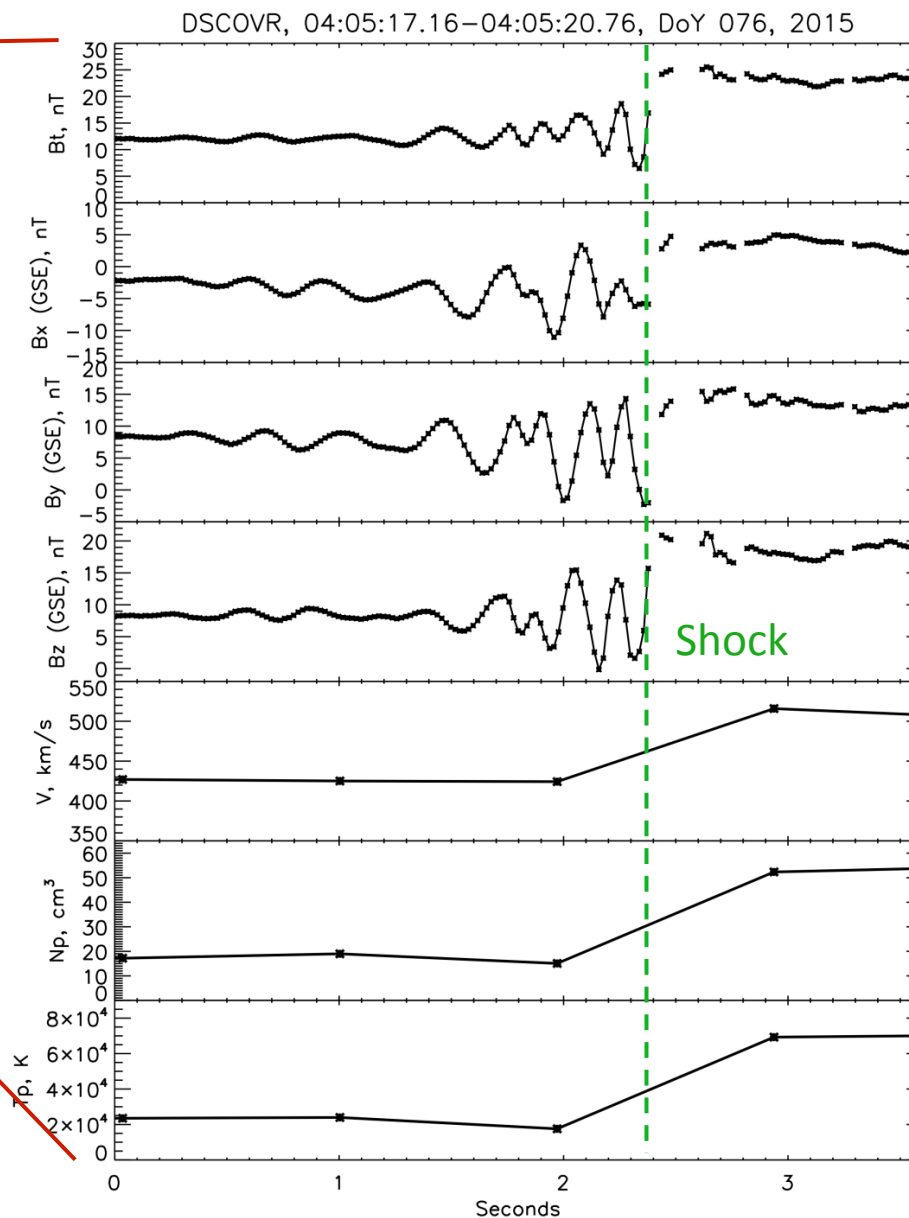
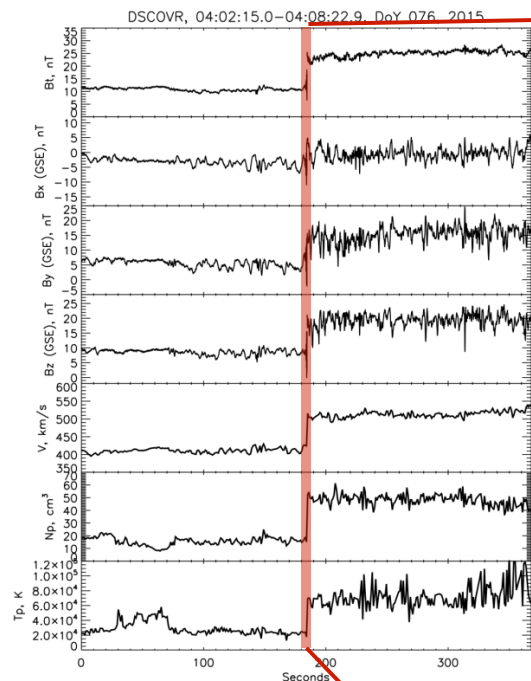


- DSCOVR and Wind separated by < 10 Re perpendicular to SW





High Time Resolution Data



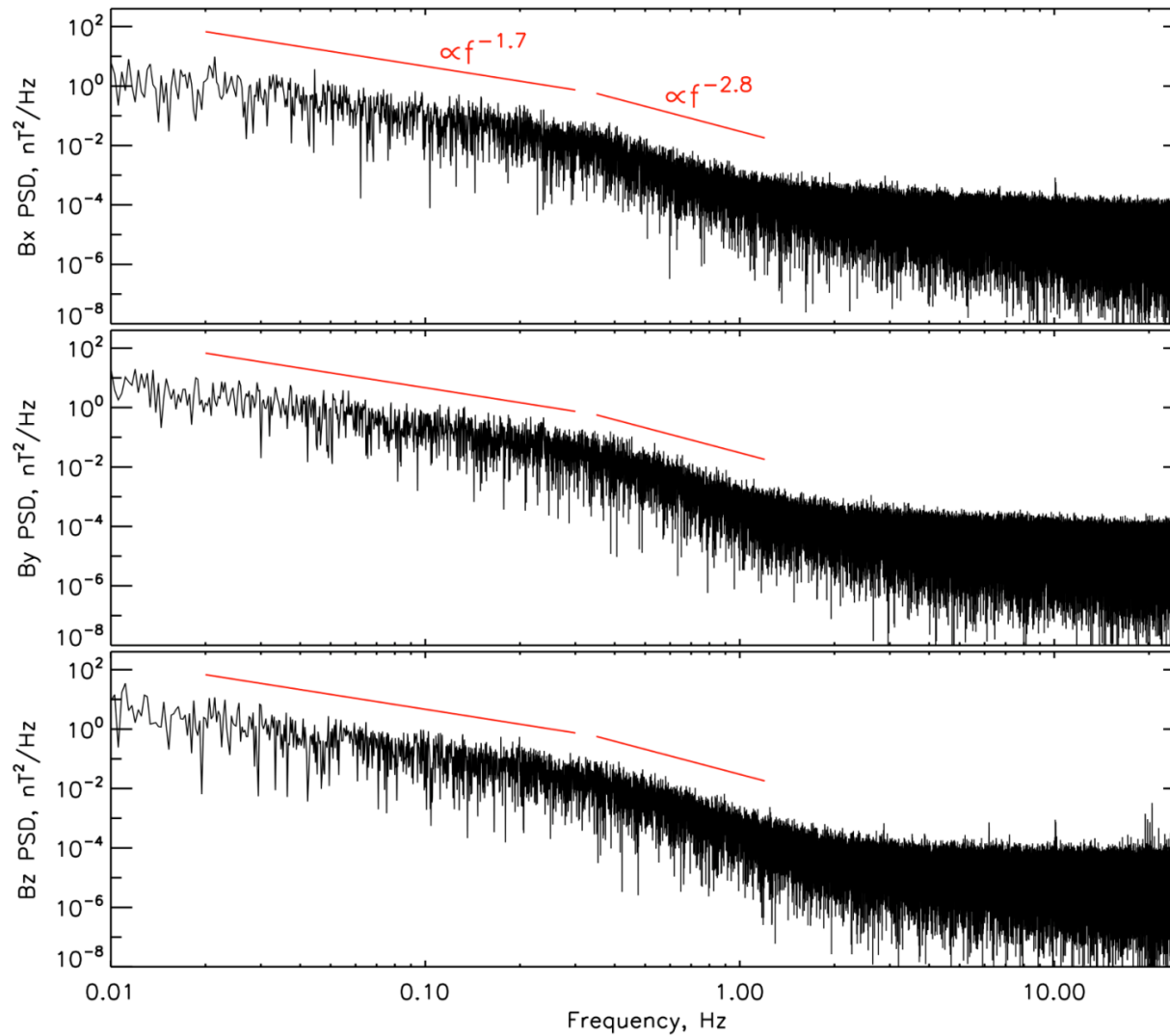
Large amplitude waves upstream of shock



Magnetic Power Spectrum



- Inertial and dissipation ranges of magnetic turbulence



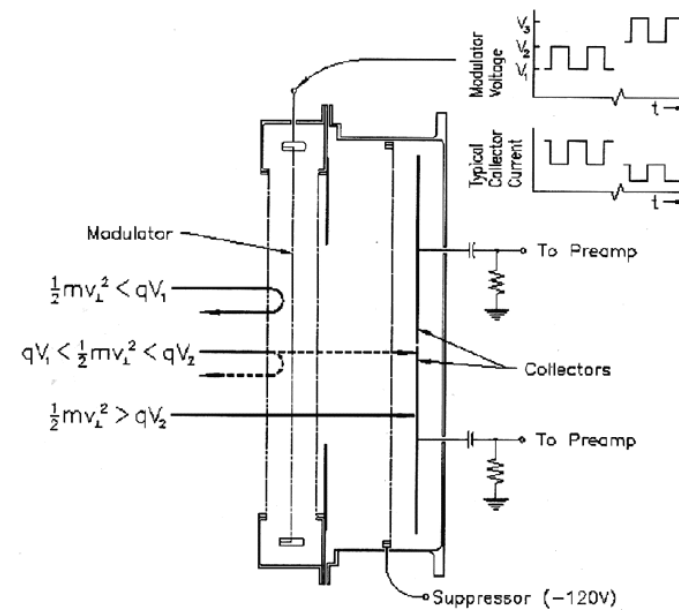
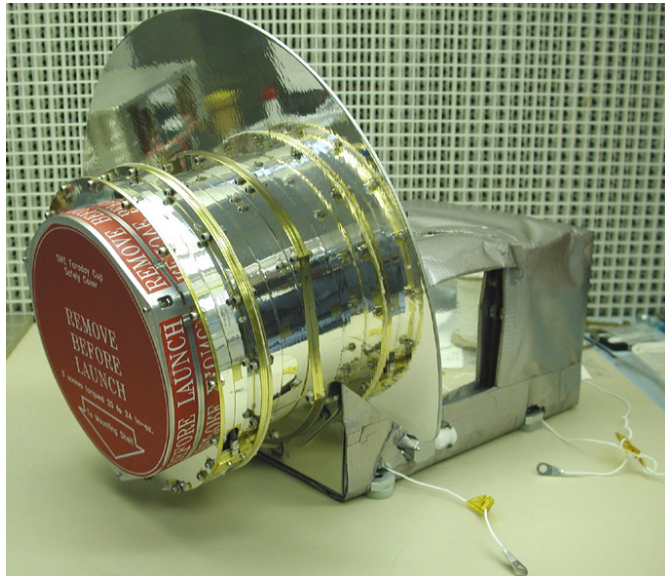


DSCOVR Faraday Cup - SAO/MIT/GSFC



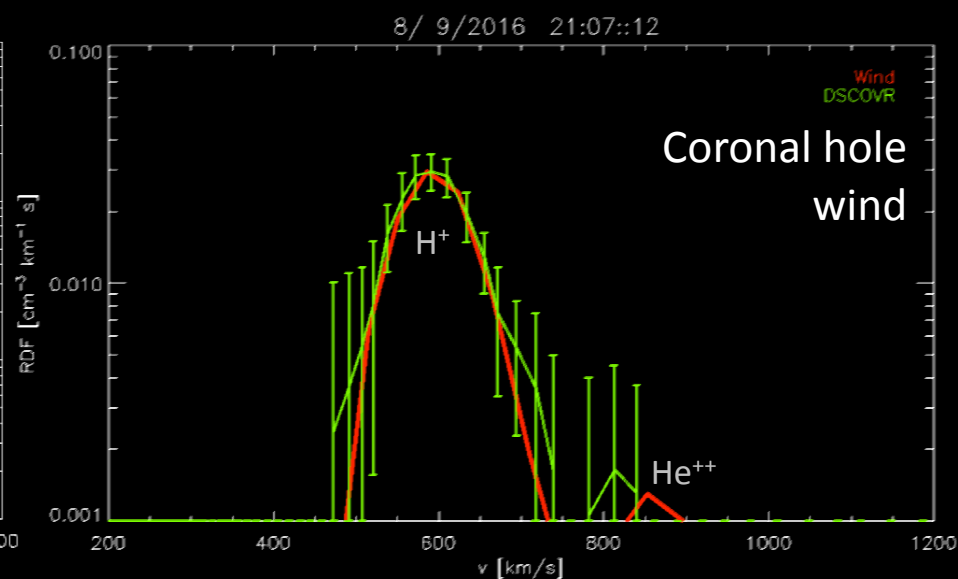
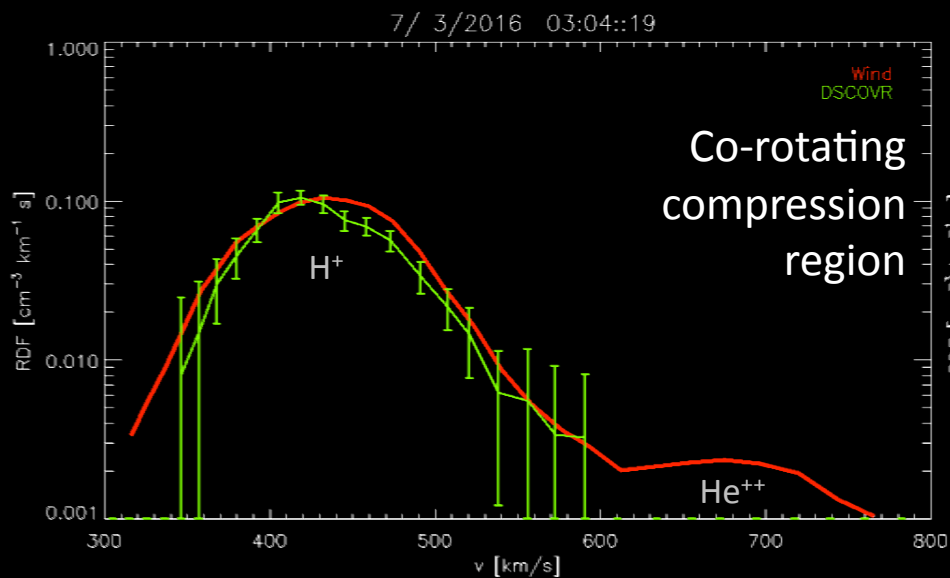
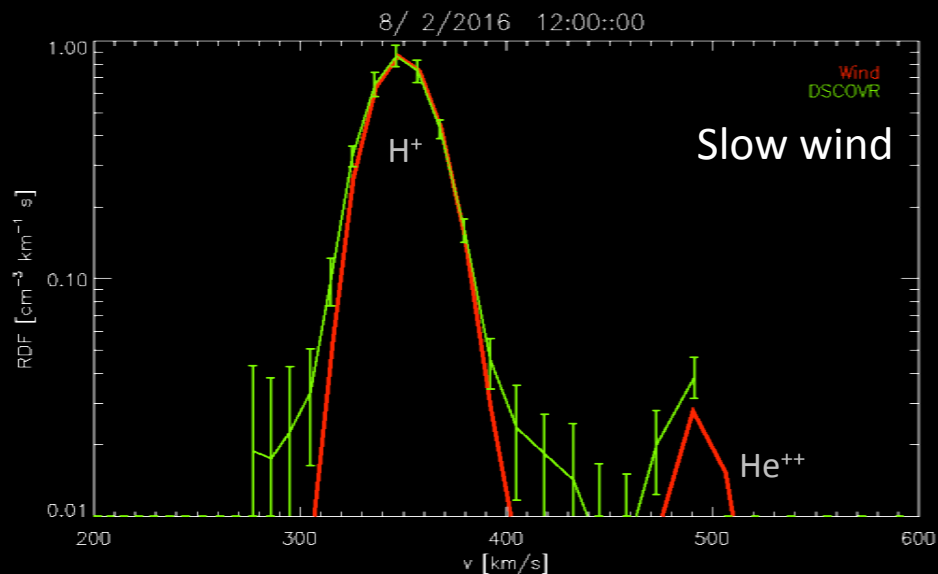
The Faraday Cup is a retarding potential particle detector that provides high time resolution solar wind proton bulk properties (wind speed, density and temperature)

Robust instrument – Can operate through high energy particle storms that commonly accompany critical space weather events



FC Data Quality

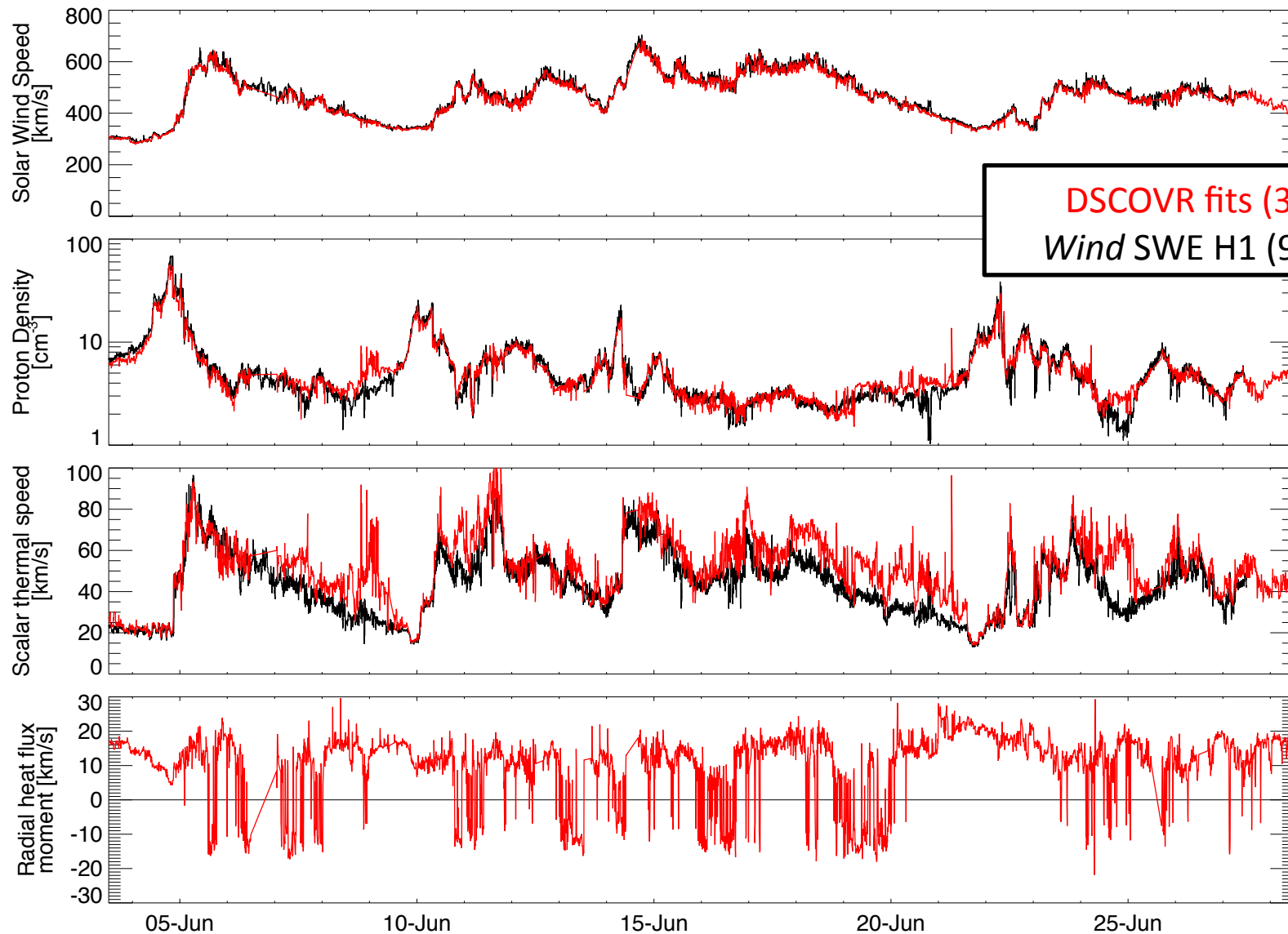
“flyback” noise eliminated and low-energy range extended



Wind SWE data (not fits)



DSCOVR – Wind Comparison

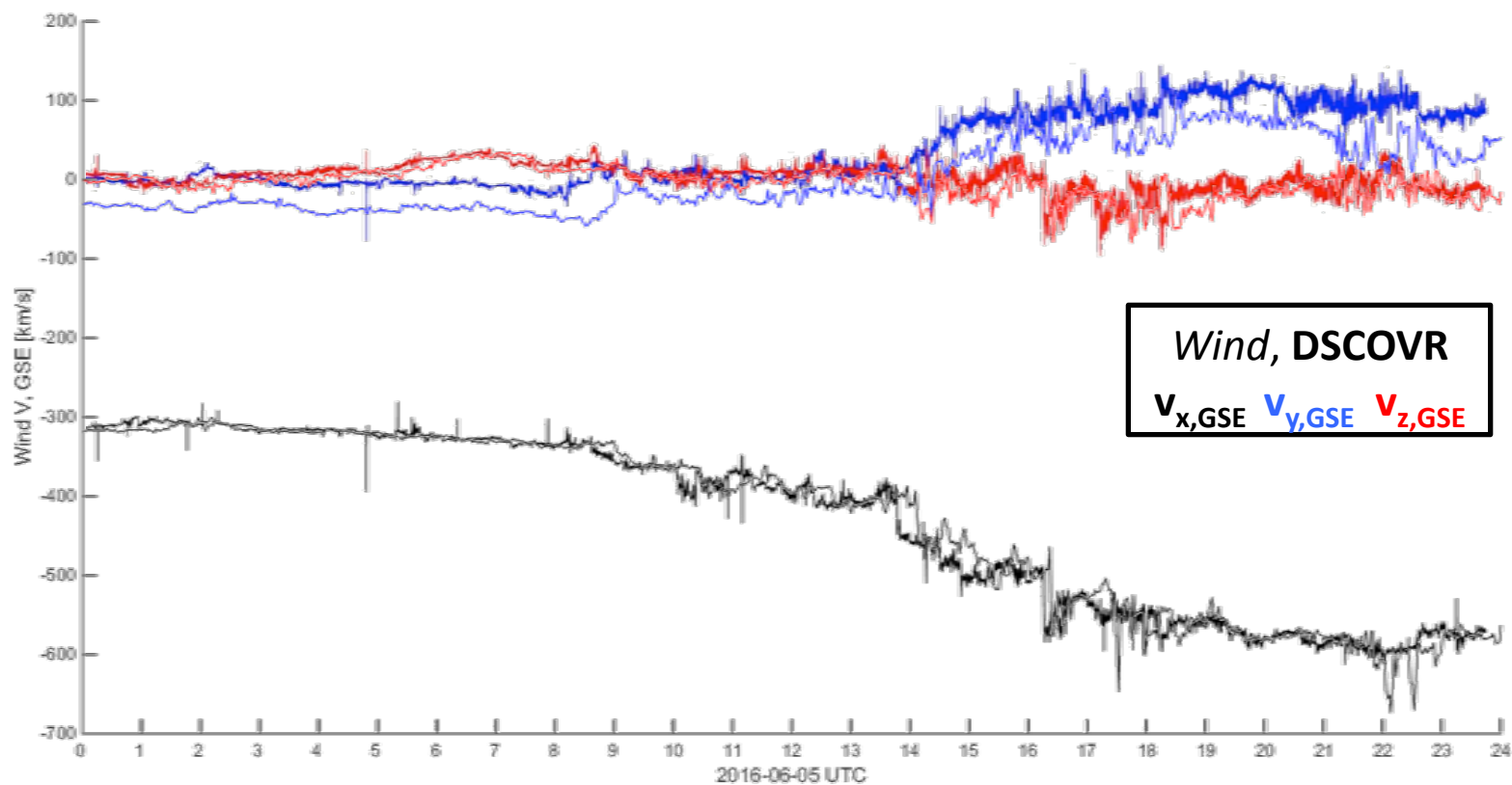


DSCOVR fits (3s)
Wind SWE H1 (92s)

2016

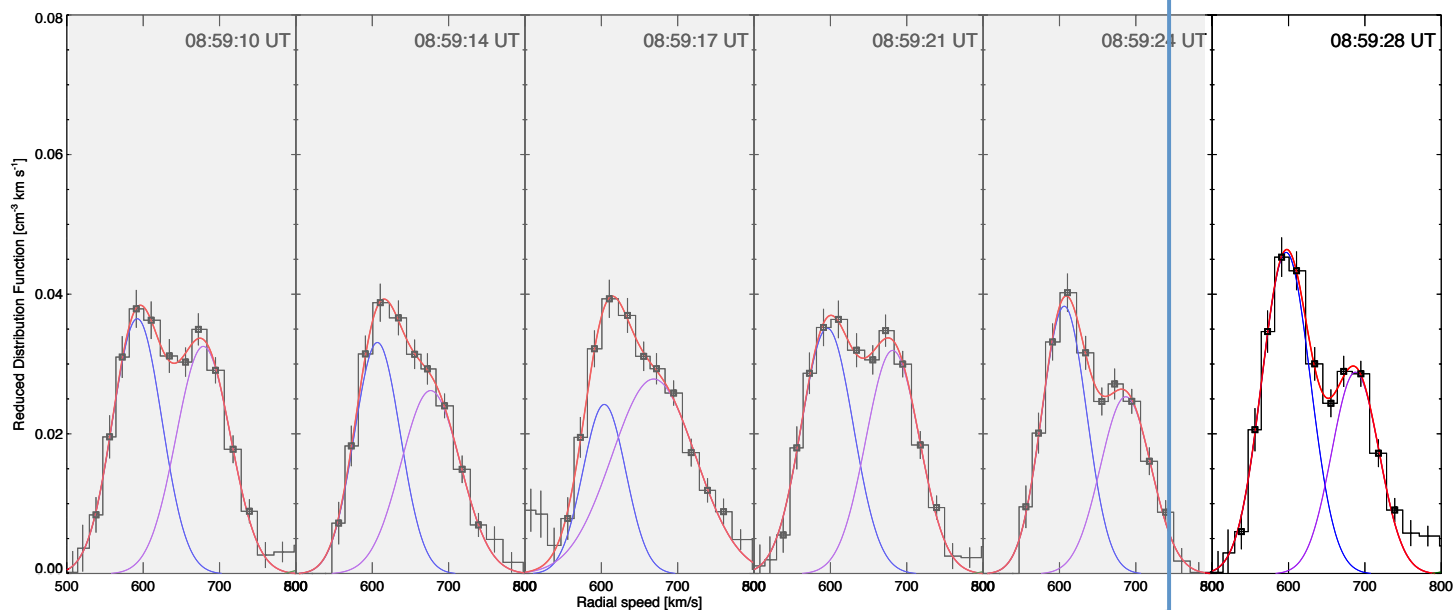


Non-radial velocity components





Counter Streaming Proton Beams



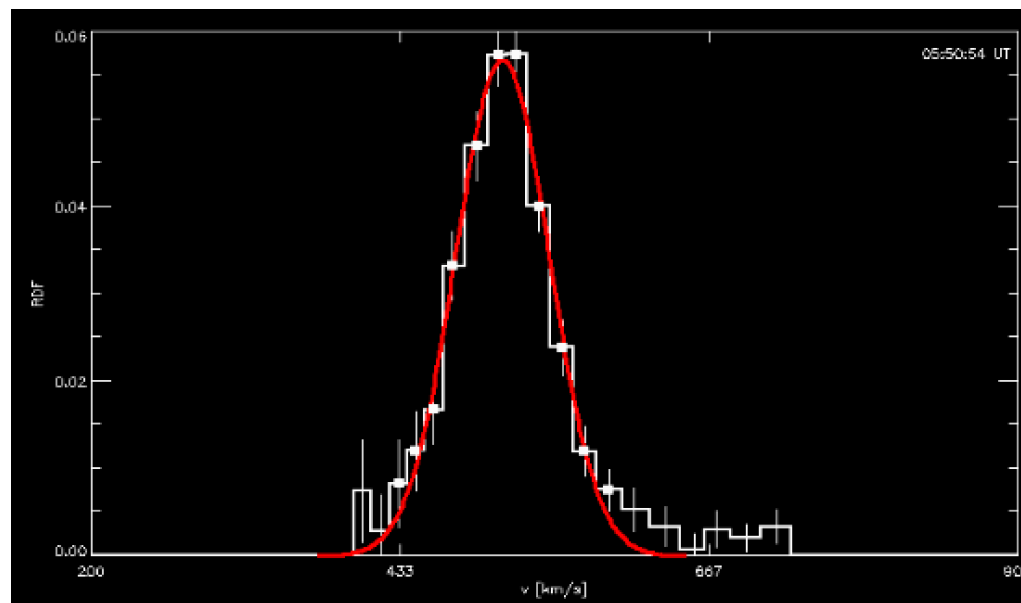
| Type of peak analysis | Speed, v (km/s) | | | Density, n (cm^{-3}) | | | Effective thermal speed, w (km/s) (i.e. Temperature) | | | |
|---|-------------------|-------|------------------|-----------------------------------|-------|------------------|---|-------|------------------|------------------|
| | u_1 | u_2 | u_{eff} | n_1 | n_2 | n_{tot} | w_1 | w_2 | Δv_{1-2} | w_{eff} |
| Two-peak fit Best known values | 622 | 634 | 631 | 3.7 | 2.2 | 5.9 | 45 | 44 | 91 | 63 |
| MOMENTS <i>DSCOVR realtime</i> | 634 | | | 5.7 | | | 52 | | | |
| Single peak fit <i>like Wind NRT keys</i> | 622* | | | 6.4* | | | 93* | | | |
| ACE realtime | 550-590* | | | 3.0-5.3* | | | 31-40* (*flagged for poor quality) | | | |



Typical Solar Wind Conditions



A more typical, equilibrium-like measurement from July 12, 2016



| Type of peak analysis | Speed, v (km/s) | | | Density, n (cm^{-3}) | | | Effective thermal speed, w (km/s) (i.e. Temperature) | | | |
|---|-------------------|-------|------------------|-----------------------------------|-------|------------------|---|-------|------------------|------------------|
| | u_1 | u_2 | u_{eff} | n_1 | n_2 | n_{tot} | w_1 | w_2 | Δv_{1-2} | w_{eff} |
| Two-peak fit | | | | | | | | | | |
| Best known values | 511 | - | 511 | 5.1 | - | 5.1 | 50 | - | - | 50 |
| MOMENTS <i>DSCOVR realtime</i> | 511 | | | 5.2 | | | 49 | | | |
| Single peak fit <i>like Wind NRT keys</i> | 510 | | | 5.1 | | | 50 | | | |
| ACE realtime | 460-512 | | | 2.8-4.8 | | | 36-48 | | | |



Conclusions



- DSCOVR became the NOAA operational L1 solar wind monitor on July 27, 2016 at 16:00 UTC
- DSCOVR is providing interplanetary magnetic field and solar wind proton key parameters to NOAA SWPC.
- Archived data is available through NOAA's National Centers for Environmental Information (NCEI) [previously NGDC]
- Reprocessed, science data will be available through CDAWeb starting in December, 2016.