

# Biological Effects of Radiation



Kerry Lee

NASA Johnson Space Center, Houston, TX



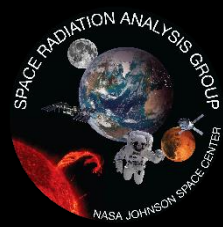
# Outline



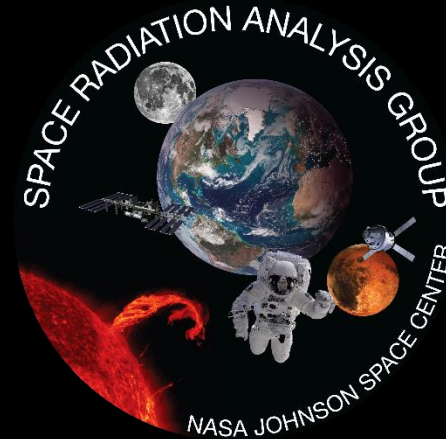
- Space Weather and SRAG
- Vocabulary Zoo
- ISS Instrumentation
- MPCV
  - BIRD Results from EFT-1
  - HERA Overview for EM-X Missions
  - ConOps for SPE Contingency
- Radiation Risk



# Space Radiation Analysis Group



- Located at the Johnson Space Center in Houston, TX
- SRAG, est. 1962
  - Real-time console operations
  - Crew, ambient monitoring
  - Pre-flight planning
  - Design evaluations
- Radiation Health Office
  - Interpretation
  - Record Keeping
  - Risk Estimation
  - Crew Selection





**Integration/Projects**  
 Program Manager - Rhodes, Bradley J.  
 Haven, Cindy  
 Mcmonigal, Kathleen, M.D.  
 Hill, Terry – CH&S Project Manager

**Space & Clinical Operations Division (SD)**  
 Chief – Taddeo, Terrance, M.D.  
 Deputy Chief – Keprta, Sean, MS, CIH  
 Associate Division Chief – Vacant  
 Lead Secretary – Durrschmidt, Doreen, RCJV  
 Admin Officer – Nimmons, Deitra



**Budget**  
 Hare, Stacy  
 Sanchez, Maria

**Human Space Flight Operations Branch (SD2)**  
 Acting Chief – Srb, Nancy  
 Secretary – Durrschmidt, Doreen, RCJV

**Occupational Health Branch (SD3)**  
 Chief – Tarver, Bill, M.D.  
 Deputy Chief – Fogarty, Jennifer, Ph.D.  
 Delgado, Richard – Business Support  
 Secretary – Gaines, Karen, RCJV

**Medical Operations Group (MOG)**  
 Lead – Dervay, Joseph, M.D.  
 Alexander, David, M.D.  
 Bauer, Pete, M.D.  
 Beven, Gary, M.D.  
 Chamberlain, Blake, M.D.  
 Effenhauser, Rainer, M.D.  
 Gilmore, Steve, M.D.  
 Hart, Steve, M.D.  
 Holland, Al, Ph.D.  
 Law, Jennifer, M.D.  
 Moynihan, Shannan, M.D.  
 Scheuring, Richard, D.O.  
 Schmid, Josef, M.D.  
 Stepaniak, Phil, M.D.  
 Ulissey, Lars, M.D. (USAF)

**Mission Support Group (MSG)**  
 Lead -  
 Blatt, Terri  
 Haven, Cindy  
 Kinder, Kristen  
 McCollum, Suzanne  
 Neigut, Joe  
 Sams, Clarence, Ph.D.  
 Sanchez, Rick  
 Schwanbeck, Nicole  
 Spence, Lisa  
 Torney, Susan  
 Weaver, Anita

**Radiation Group (RG)**  
 Lead - Semones, Edward, CHP  
 Bahadori, Amir (SRAG)  
 Fry, Dan (SRAG)  
 Lee, Kerry (SRAG)

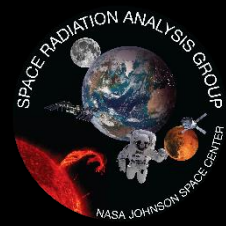
**Occupational and Aerospace Clinics**  
 Johnston, Smith, M.D.  
 Locke, James, M.D.  
 Shackelford, Linda, M.D.

**Human Test Support**  
 Locke, James, M.D.

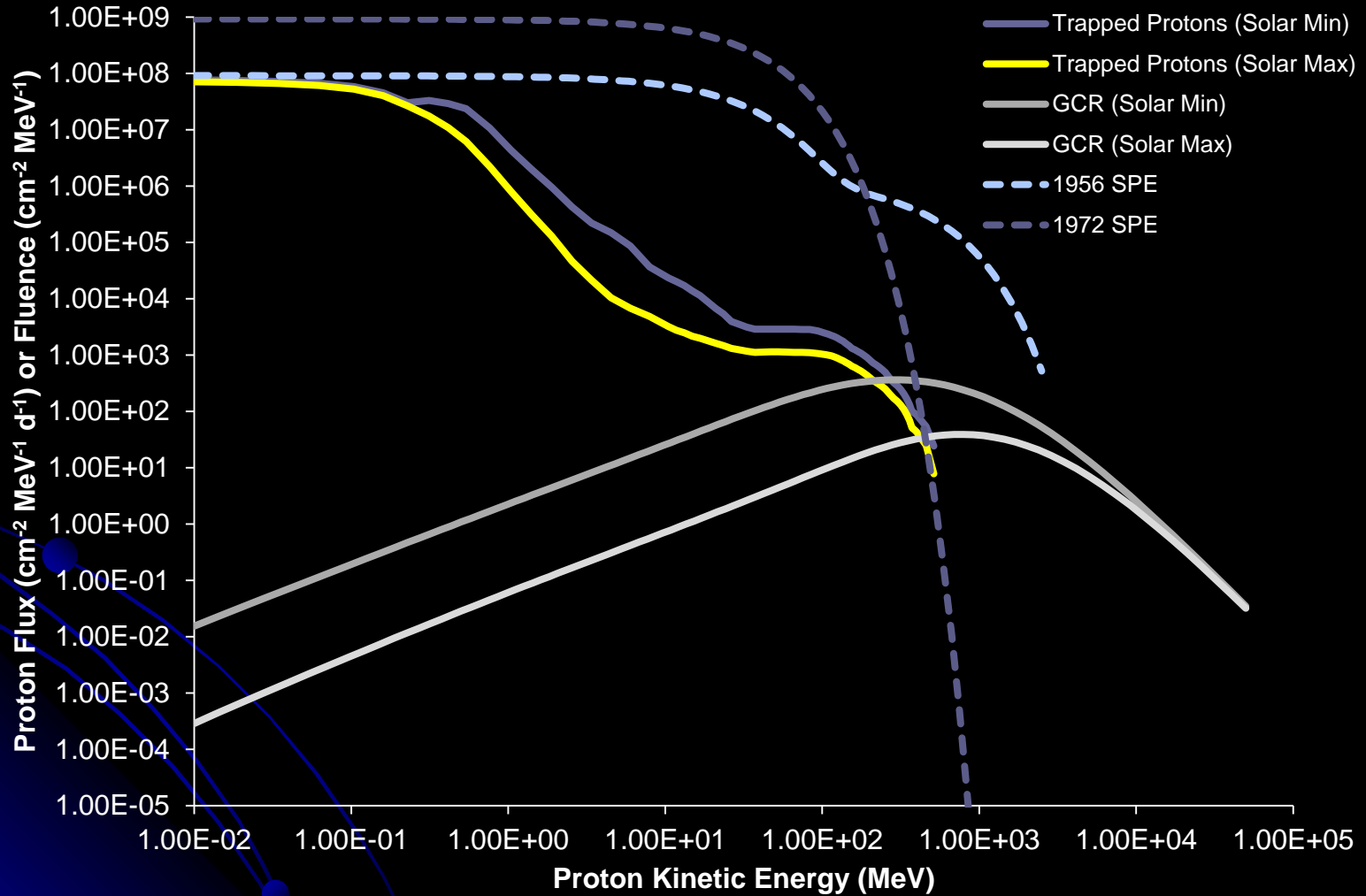
**Human Test Subject Selection**  
 Shackelford, Linda, M.D.

**Astronaut Occupational Health/Epidemiology**  
 Van Baalen, Mary  
 Kadwa, Binaifer (Bini)

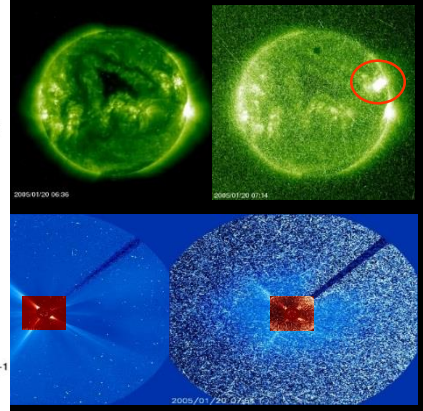
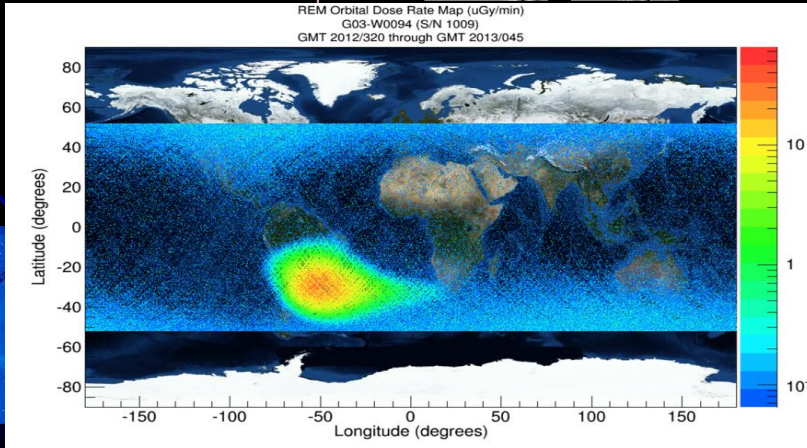
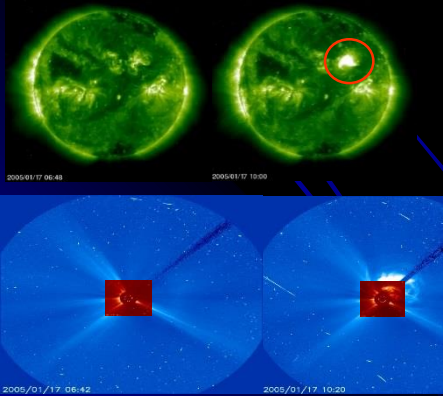
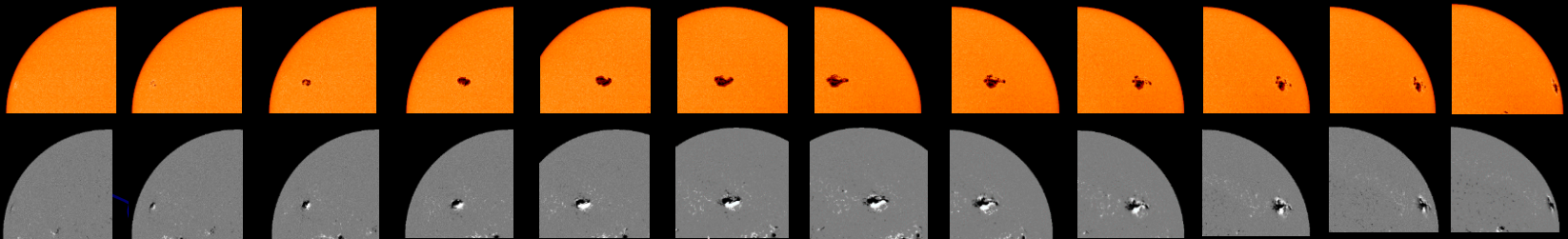
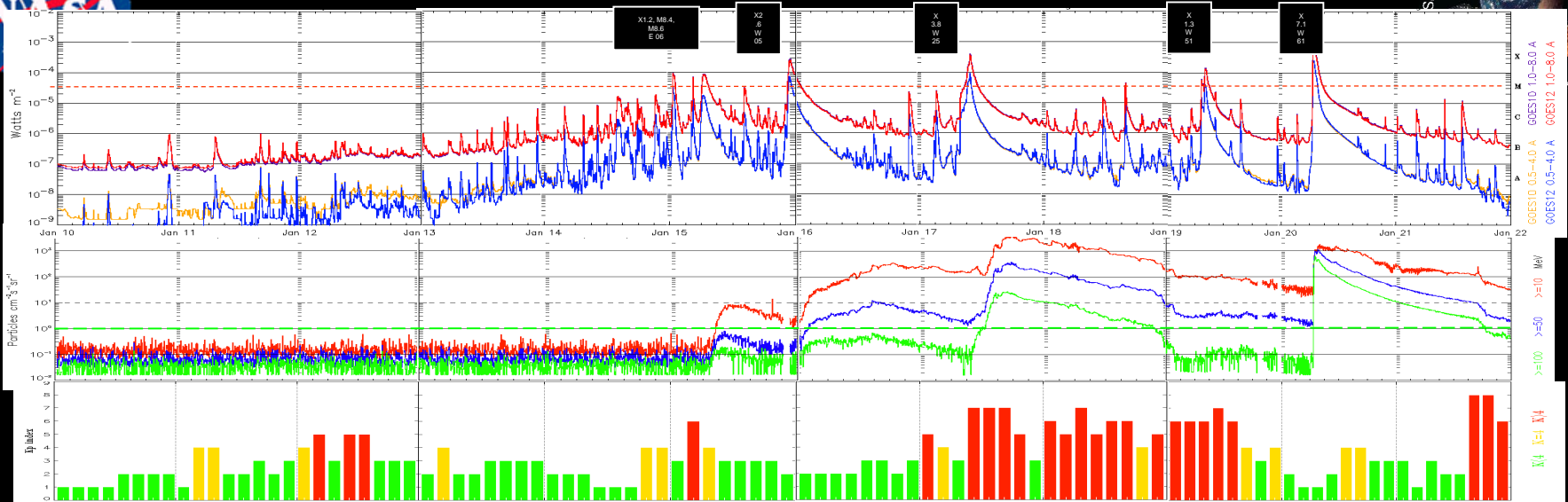
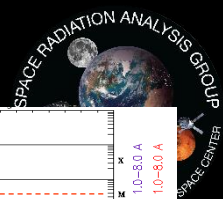
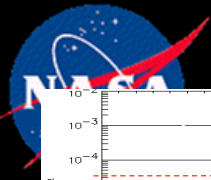
**Institutional Support**  
 Martel, Robert, MPH  
 Plaza, Angel, Ph.D.  
 Stanch, Penney, CIH, CSP



# Radiation Sources

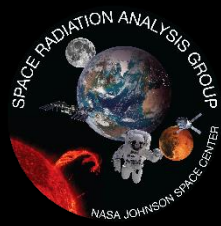


# NOAA Active Region 0720 January 10, 2005 - January 21, 2005

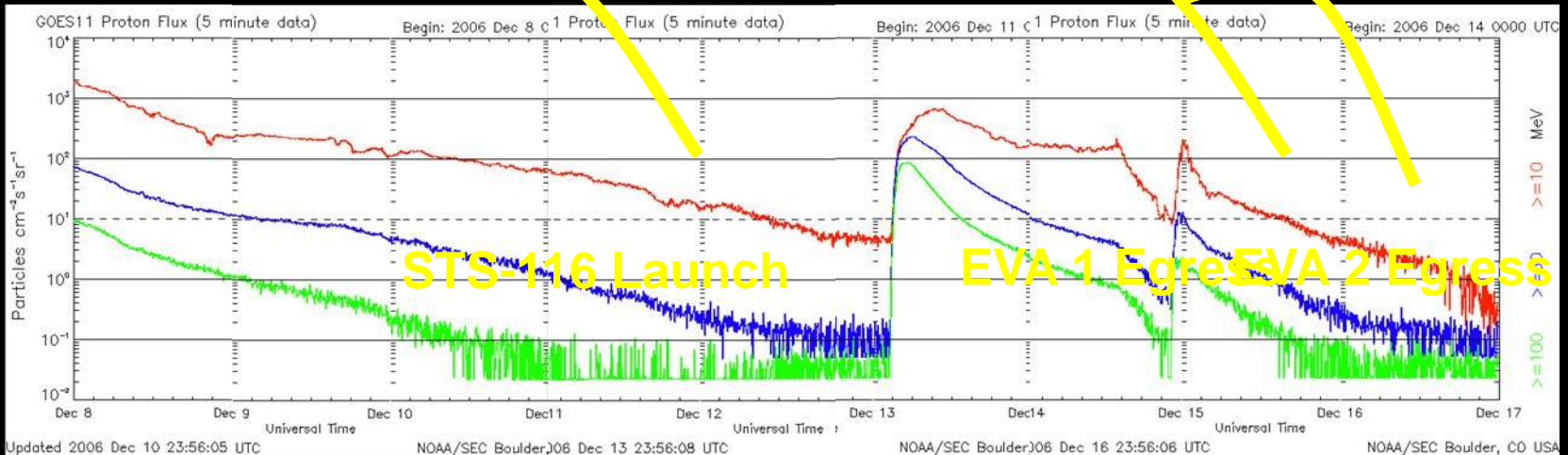
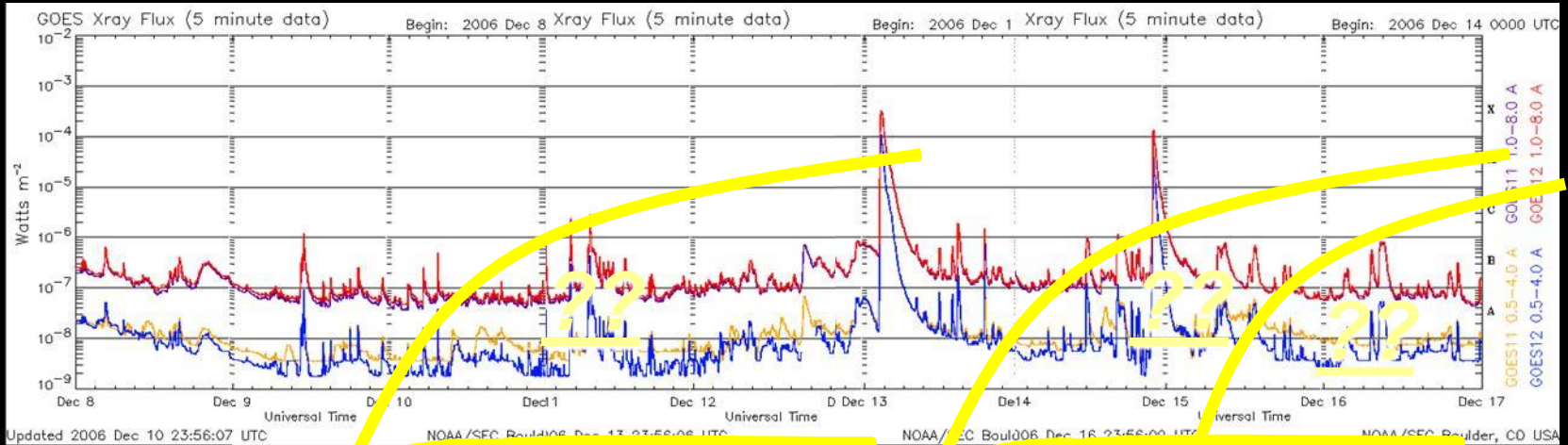




# STS-116: Example Progression



## Probability?





# Space Weather Terminology: Words are Important



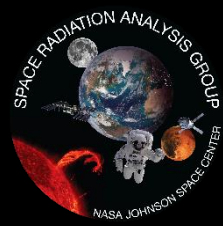
- **Solar Flare** – large amount of electromagnetic energy released from the Sun. In particular, X-ray emission is monitored as there is an observed correlation between X-ray energy release and particle acceleration. The electromagnetic energy release has no operational impact. The X-rays has much lower intensity than dental x-rays
- A **Solar Particle Event** (SPE) is denoted by observation of proton flux of Solar Energetic Particles (SEP) greater than a given energy (e.g.  $>10\text{MeV}$  or  $>100\text{ MeV}$ )
- **Coronal Mass Ejection** (CME) – results when the energy released by a flare is great enough to throw solar mass (mostly protons) with a velocity great enough to escape the Sun's gravity and magnetic fields. CMEs take 1-3 days to arrive at Earth and disturb the geomagnetic field.

Visual Terrestrial Analogy





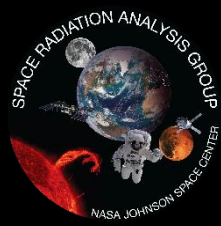
# Biological Radiation Terminology



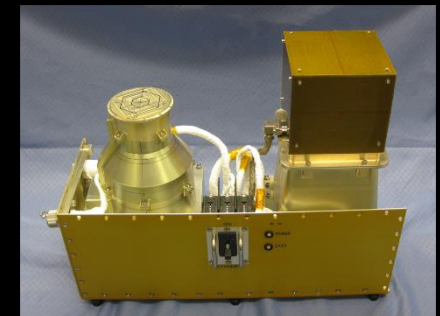
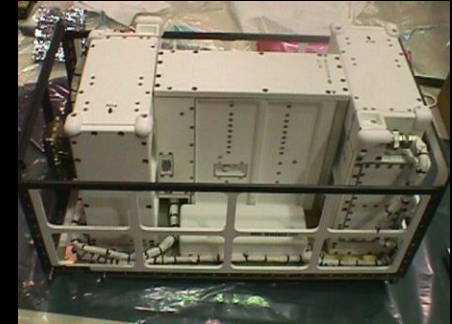
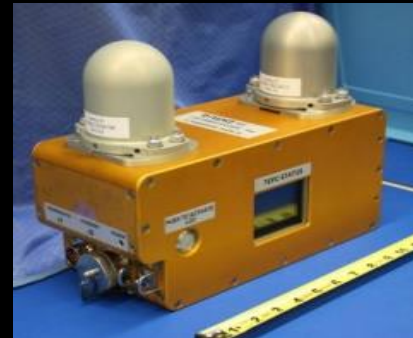
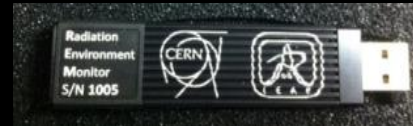
- Dose
  - Measureable quantity of Energy per mass ( $\text{J/kg} = \text{Gy}$ )
- Dose Equivalent
  - Dose multiplied by a biological effectiveness quality factor (Q), that is a function of LET (Linear Energy Transport) (units of Sv)
- Equivalent Dose
  - Dose multiplied by radiation type specific weighting factor. This weighting factor is a function of particle type and energy
- Effective Dose
  - A way to determine whole body radiation effects by weighting equivalent dose within an organ by the organs tissue weighting factor and summing over the exposure to all organs.
- Radiation Cancer Risk
  - A calculated chance of some given radiation exposure to result in cancer



# ISS Instrumentation

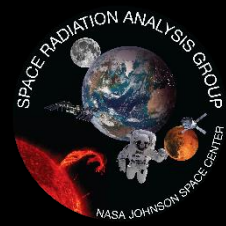


- CPDS – Charged Particle Directional Spectrometer
- REM – Radiation Environment Monitor
  - Active dosimeter with USB interface
- TEPC – Tissue Equivalent Proportional Counter
  - Located in ISS Service Module
- IV-TEPC – new TEPC detector
  - Moves about ISS every 4-6 weeks
- ISS-RAD – Radiation Assessment Detector



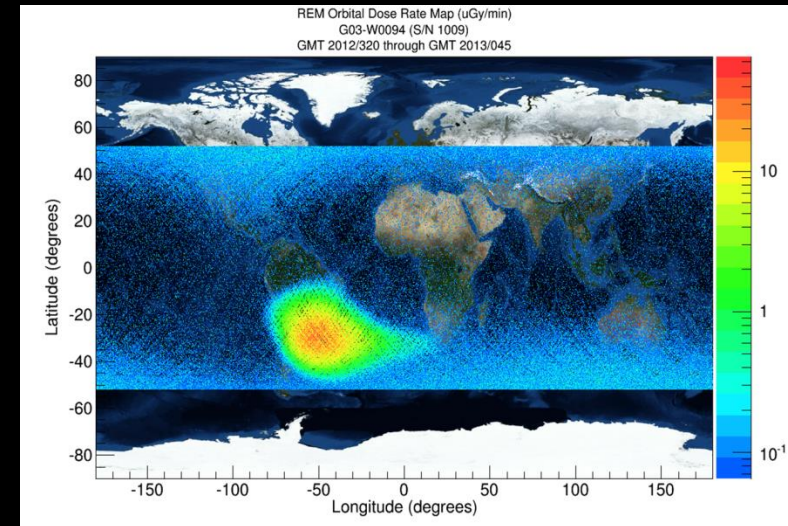
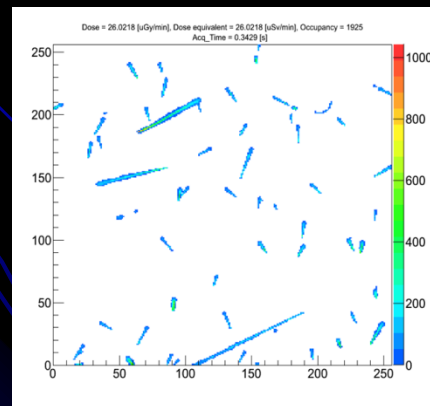
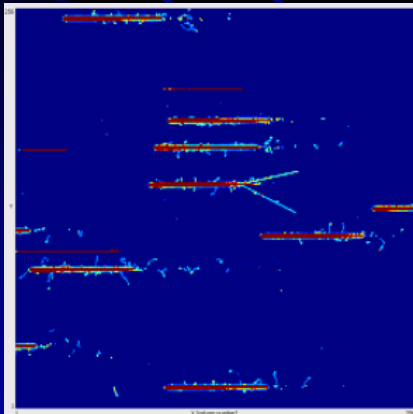
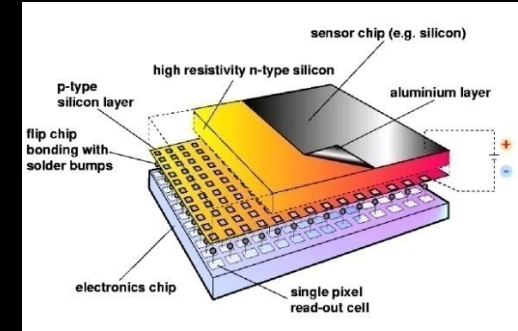


# Radiation Environment Monitor (REM)



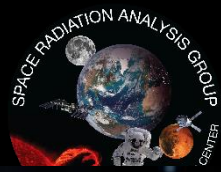
## The Timepix Detector

- Developed as a High Energy Physics application of medical imaging technology
- Hybrid Pixel Detector with independent counting and readout circuitry in each pixel footprint
- 256 x 256 pixel grid with total area of 2 cm<sup>2</sup>





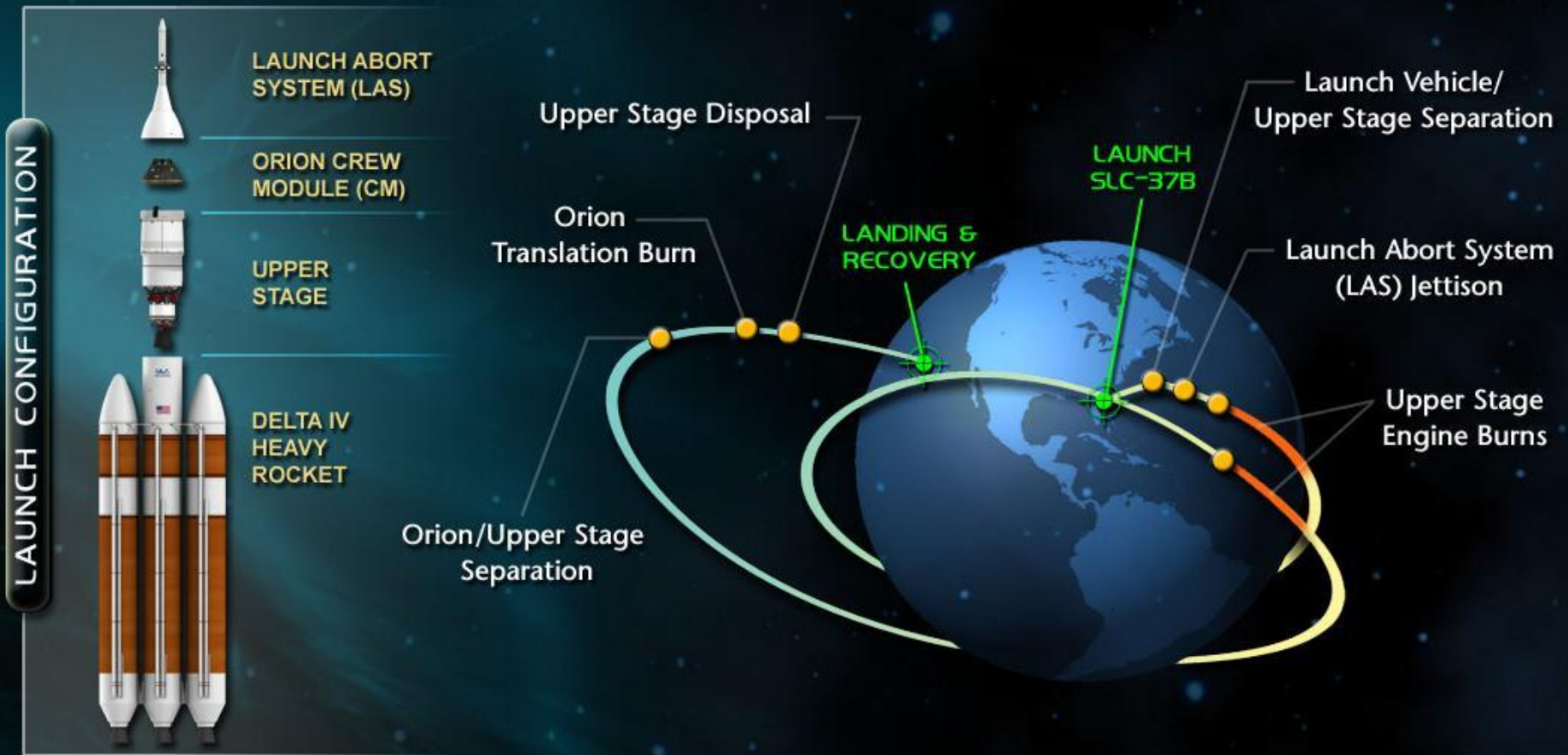
# Exploration Flight Test 1 – Dec 5, 2014



## EXPLORATION FLIGHT TEST ONE

OVERVIEW

TWO ORBITS ♦ 20,000 MPH ENTRY ♦ 3,671 MILE APOGEE ♦ 28.6 DEGREE INCLINATION



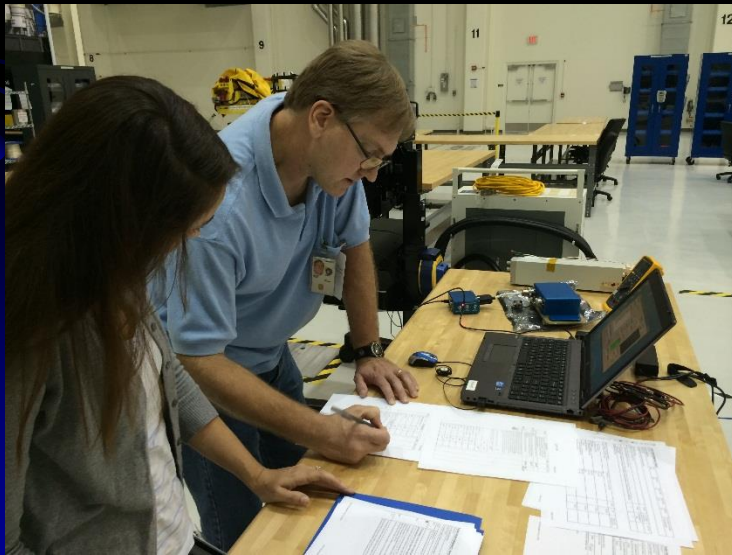
Credit: NASA



# BIRD Overview



- Oct. 2014: Flight hardware shipped to KSC
- Nov. 2014: Installed into EFT-1 vehicle
- 5 Dec. 2014: EFT-1 Launch
- 9 Dec. 2014: Recovered from vehicle
- Feb. 2015: Data report delivered to HQ
- June 2015: NASA Technical Publication (NASA/TP 2015-218575)



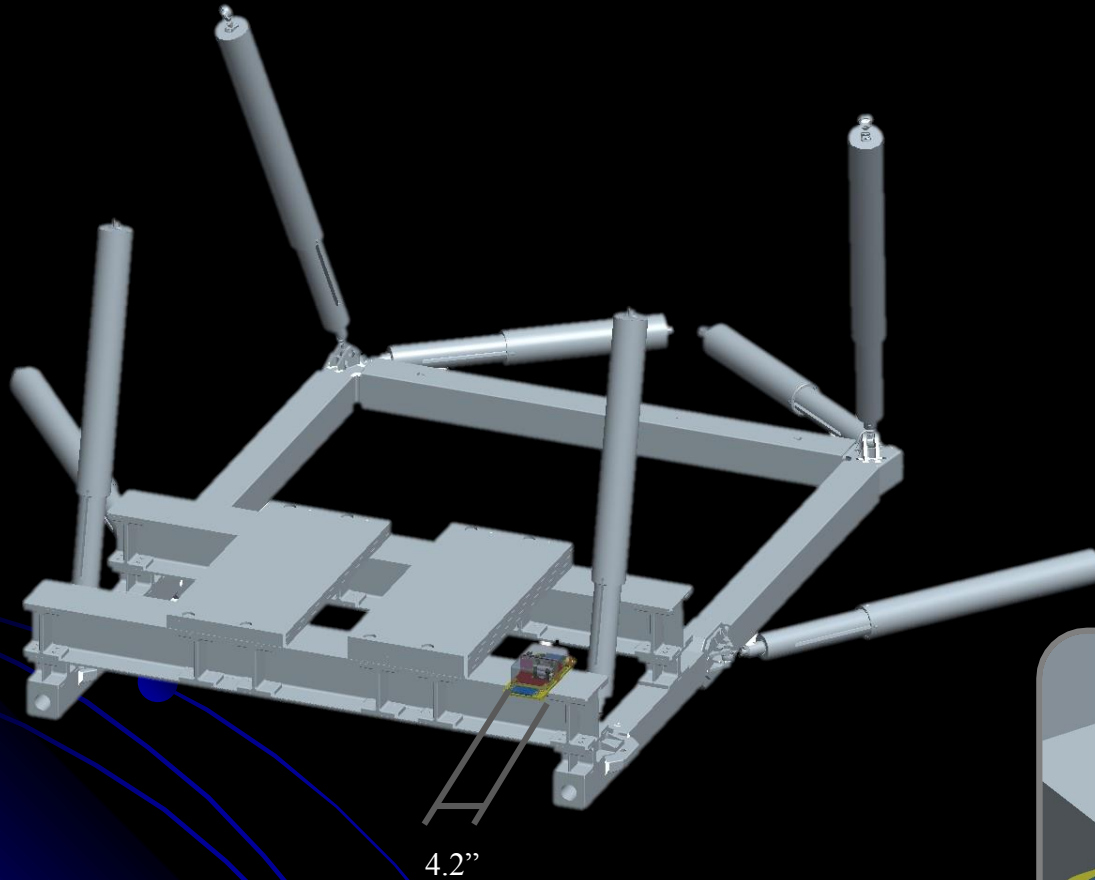
*BIRD Preflight checkout at KSC*



*Hardware post flight*

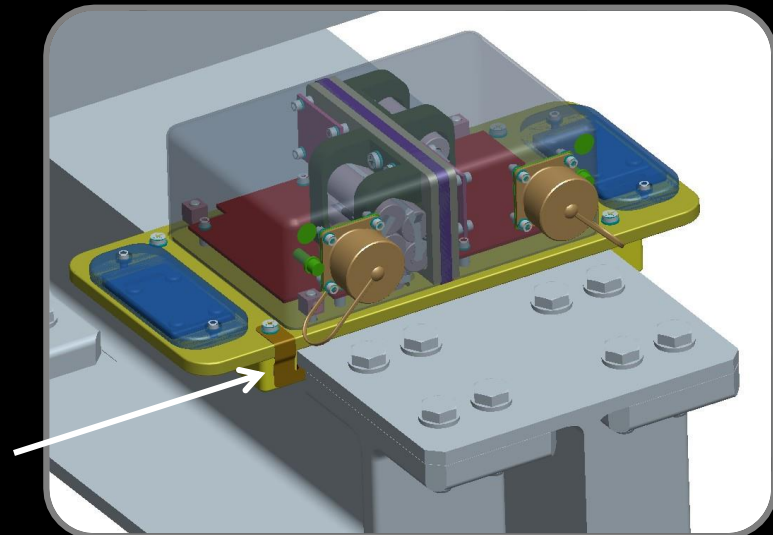


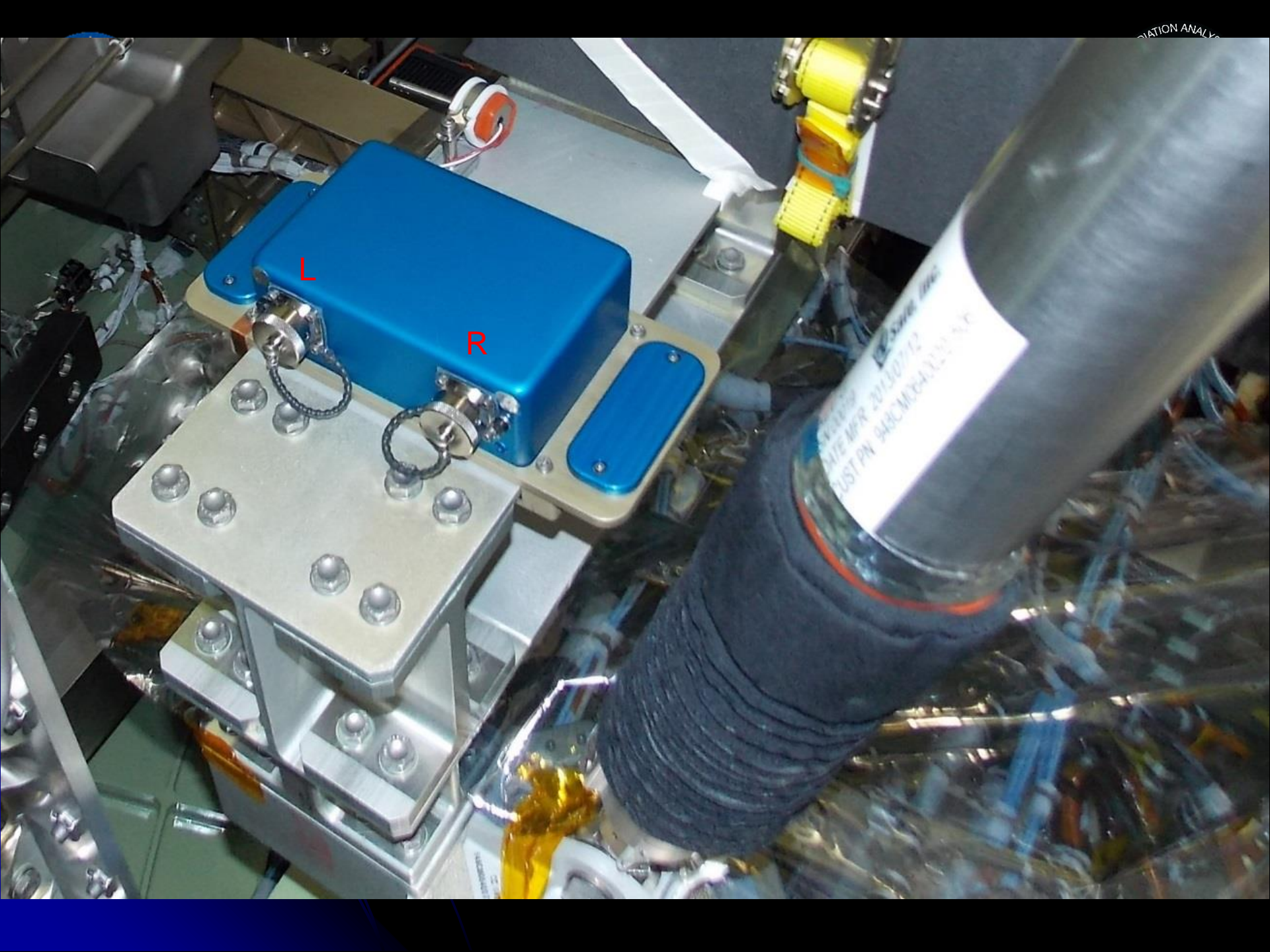
# Vehicle Attachment



4.2"

Copper bonding strip positioned to contact an alodined surface





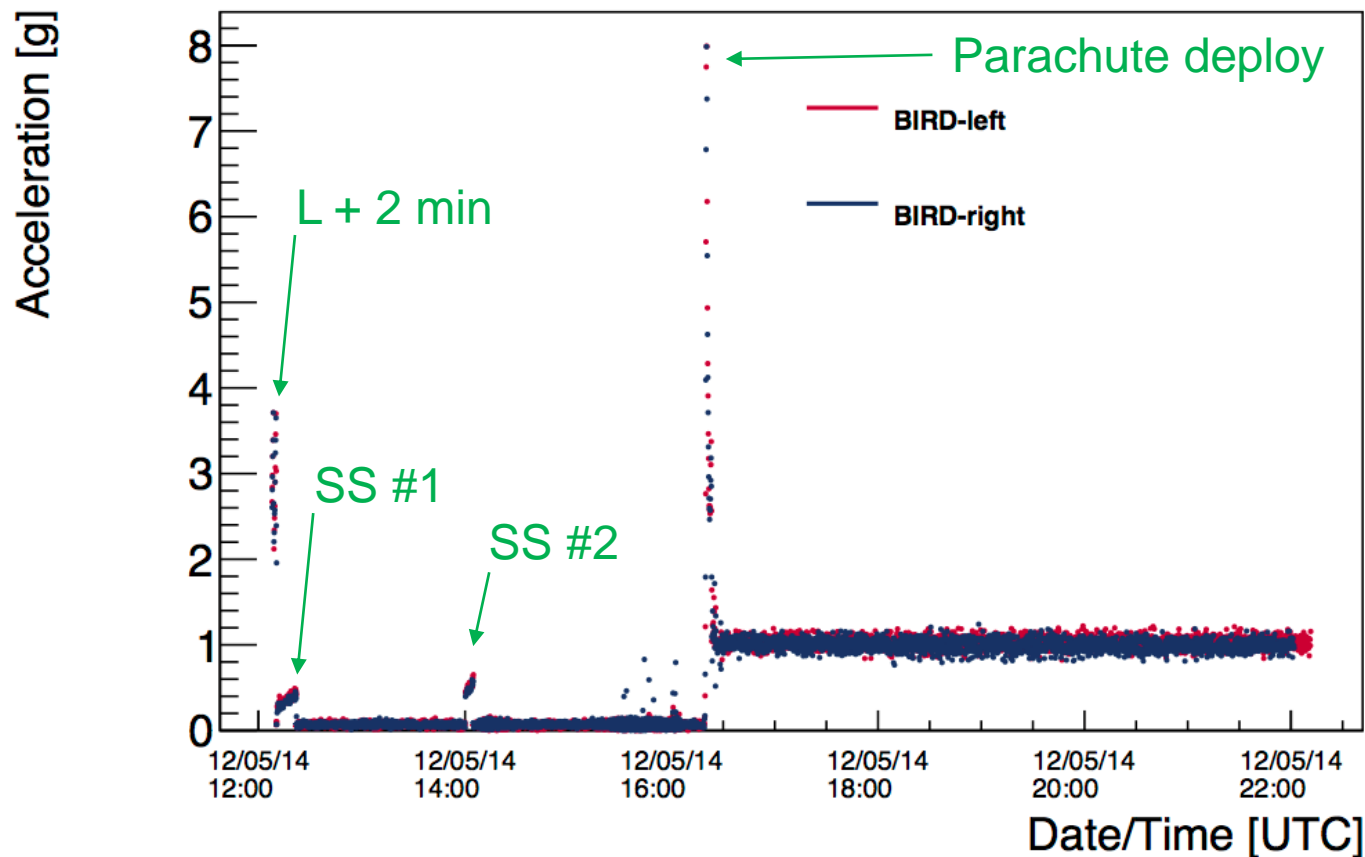
L

R

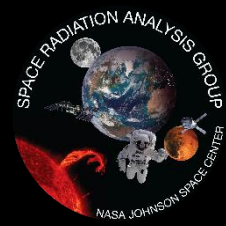
20120712  
DATE APR 2012  
JUST IN SAC/MPA/2012



# Acceleration

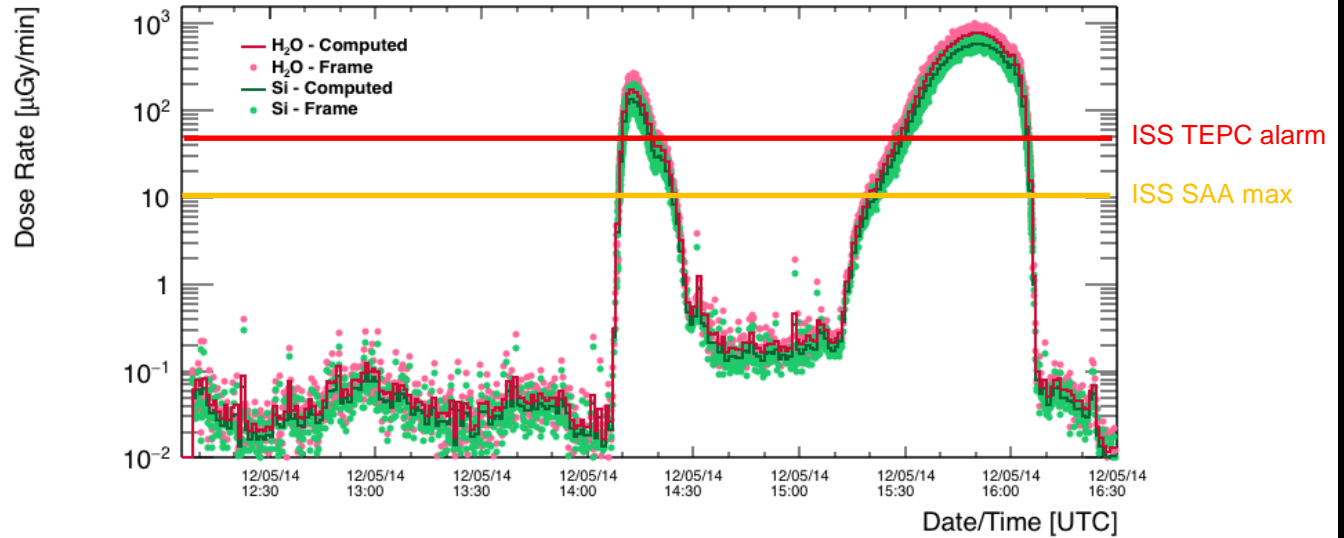




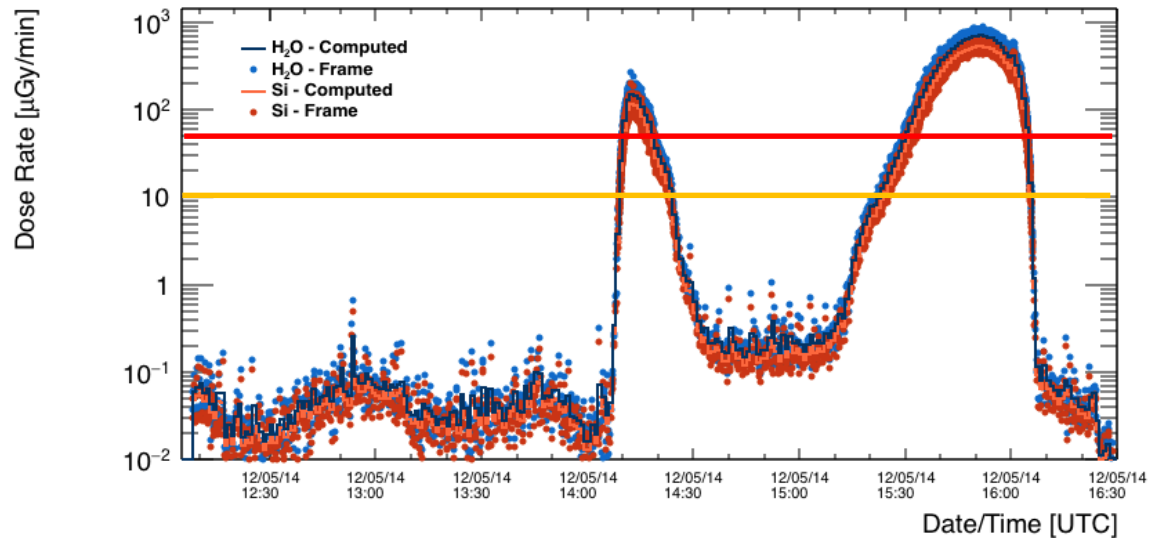


# Absorbed Dose Rates

Left Detector

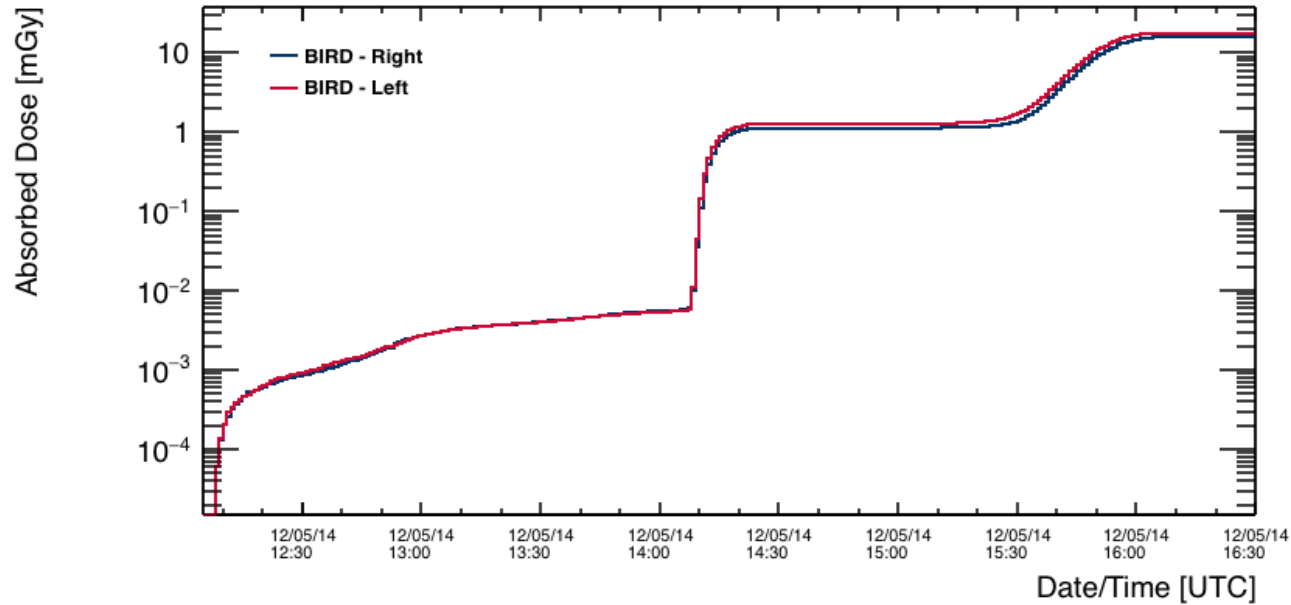
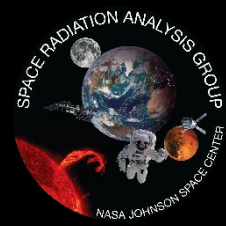


Right Detector

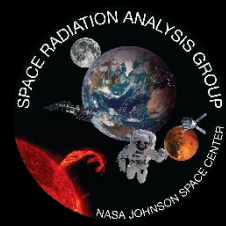




# Cumulative Absorbed Dose

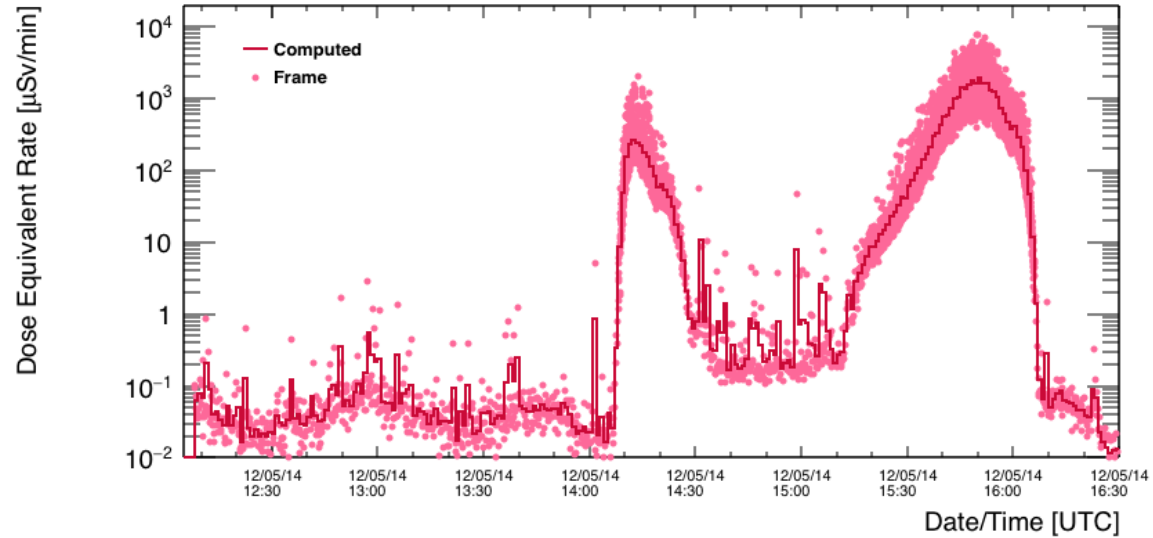


	BIRD [mGy]	RAM [mGy]	ISS-TEPC [mGy]
Left	17.9	$15.1 \pm 0.3$	0.015
Right	15.7	$13.5 \pm 0.2$	

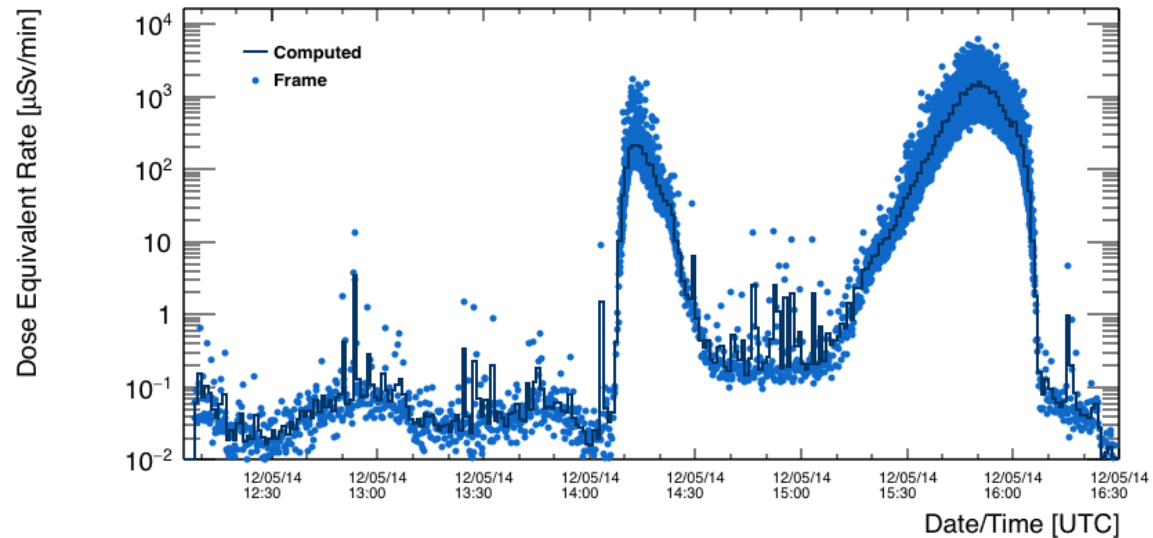


# Dose Equivalent Rates

Left Detector

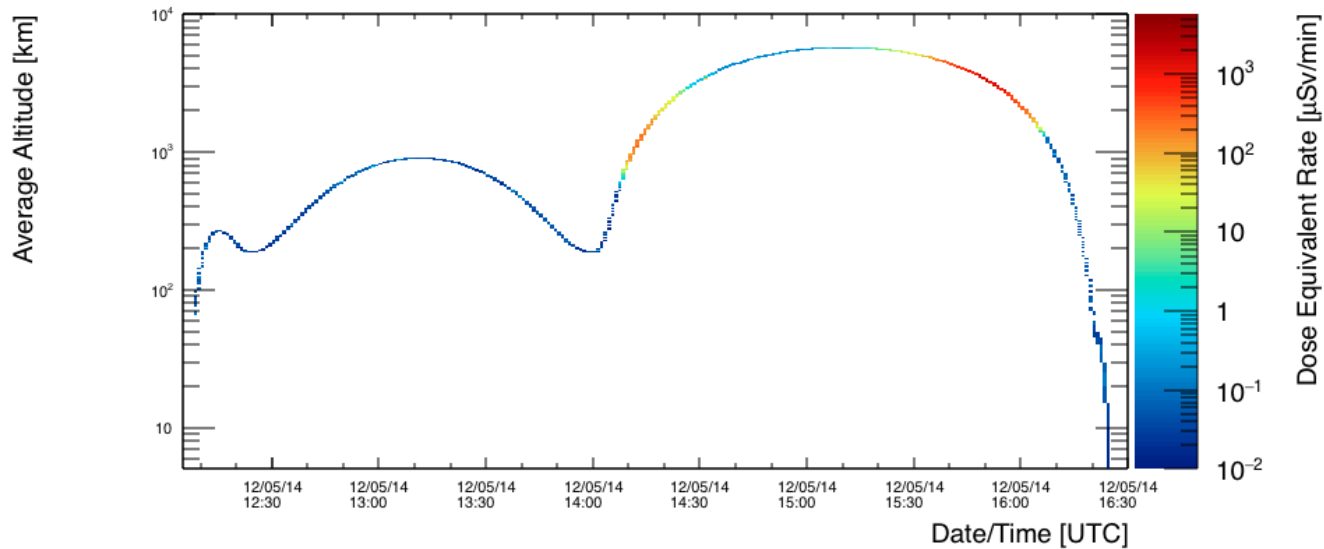
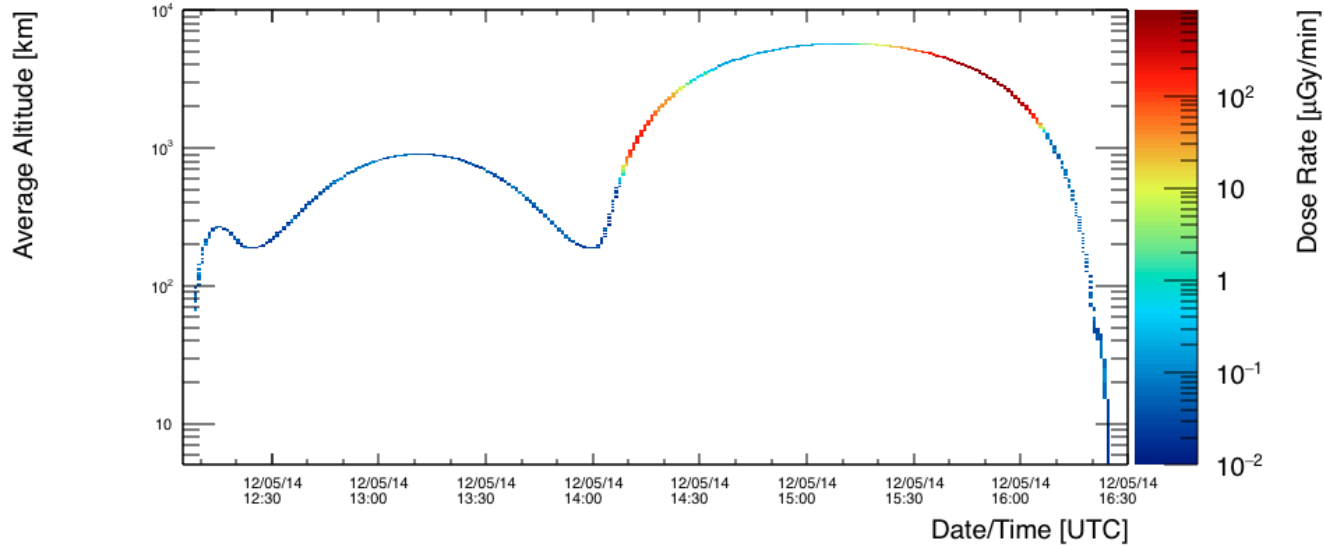


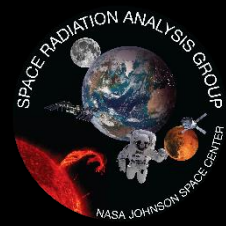
Right Detector



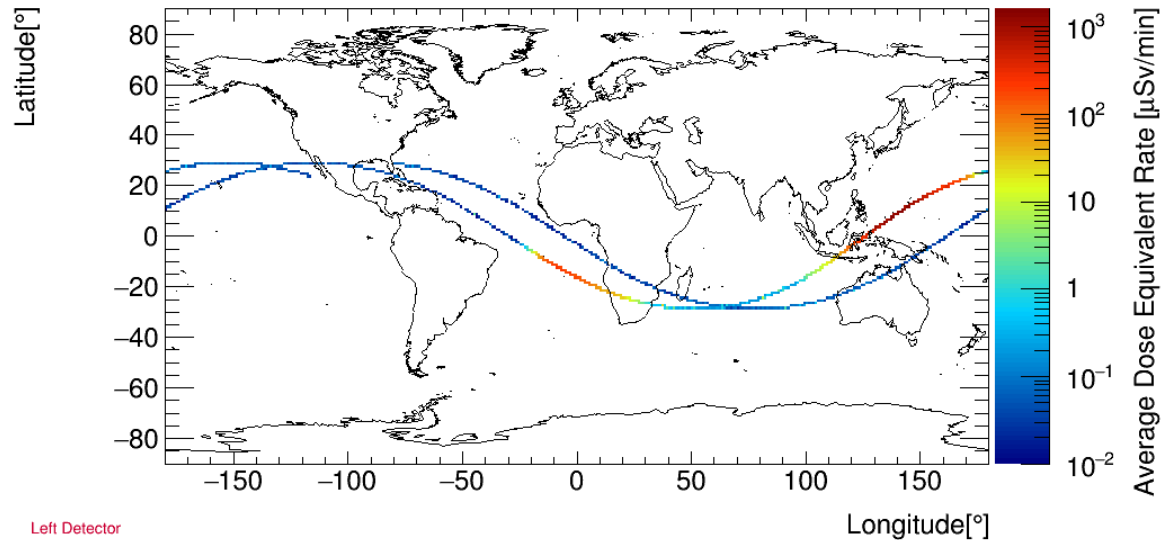
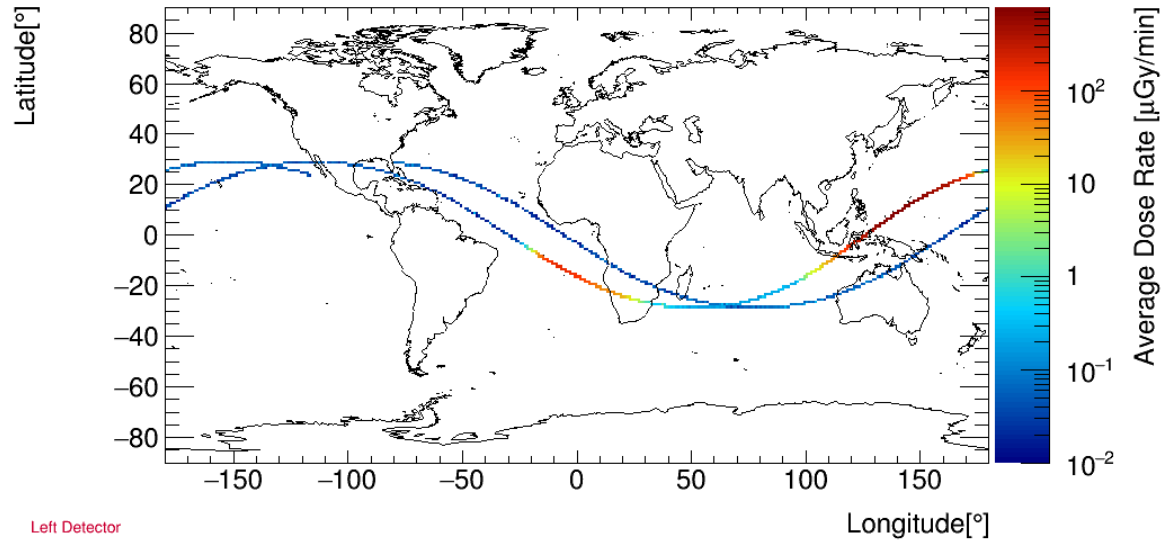


# Rates vs. Altitude and Time





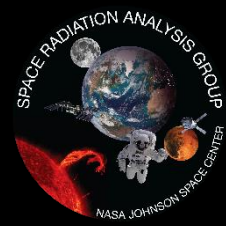
# Rates vs. Latitude/Longitude





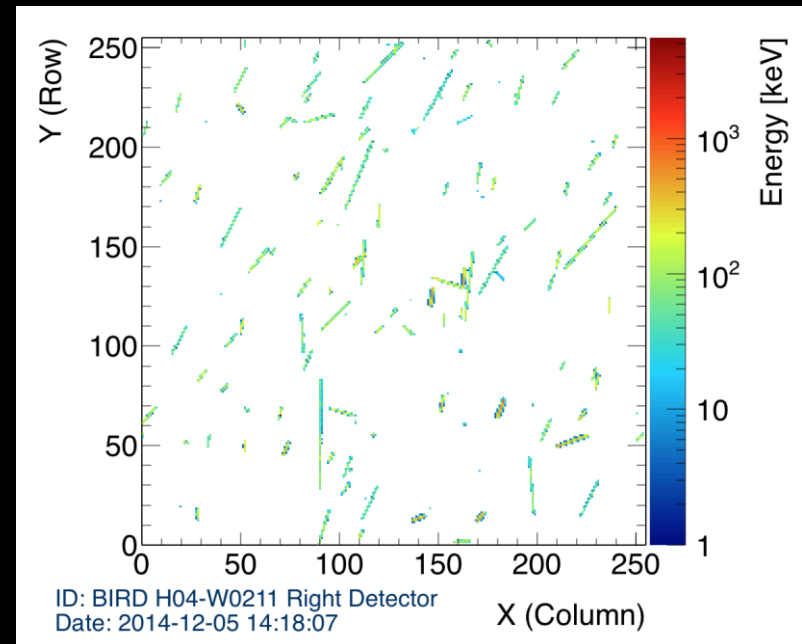
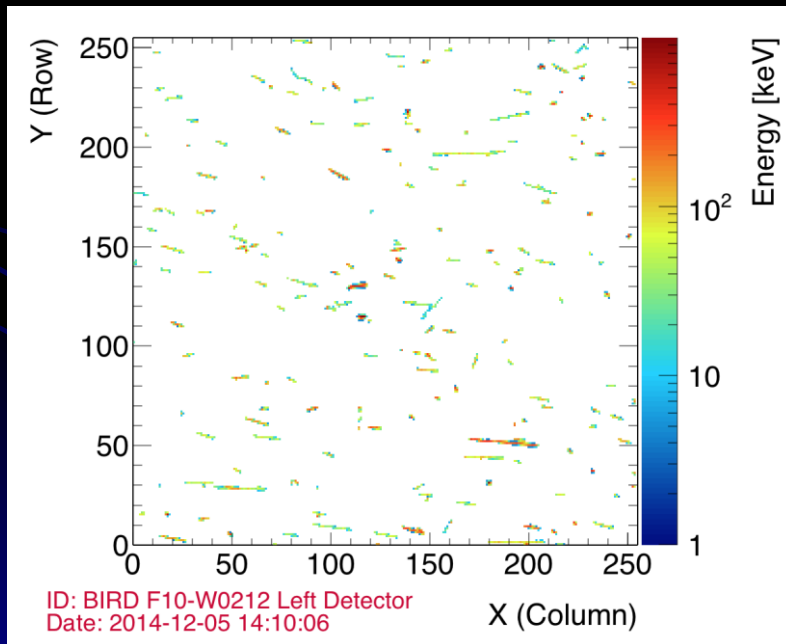
# Google Earth Video





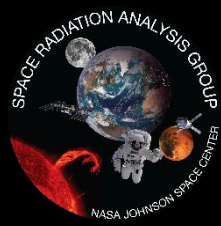
# Anisotropy

- Trapped proton environment below about 2000 km is known to be anisotropic





# EFT-1 Summary

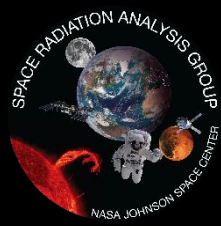


- EFT-1 presented a unique opportunity
  - First measurements in Orion MPCV
  - Information about EM-2
- Detector operation
  - Met all expectations
  - No apparent data corruption
- Data
  - Two peaks caused by spectral changes
  - Max absorbed dose rate about 1 mGy/min
  - Absorbed dose 1000x ISS TEPC



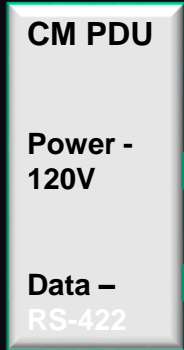


# HERA EM-1 & EM-2 Integration Configurations



Proposed HERA Hardware Block Interface Diagram for Power / Data

## EM-1 Configuration (no C&W, minimal commanding)



Orion/HERA ICD



Stowage



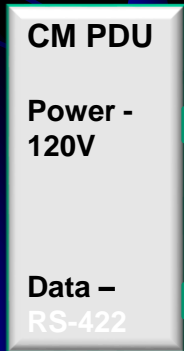
Port ECLSS Wall

Single string of HERA on EM1

C&W Data



FWD Bulkhead



LM Integrated Cabling between HERA boxes

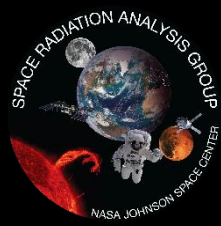
### Integration Responsibilities

LM Vehicle Provided

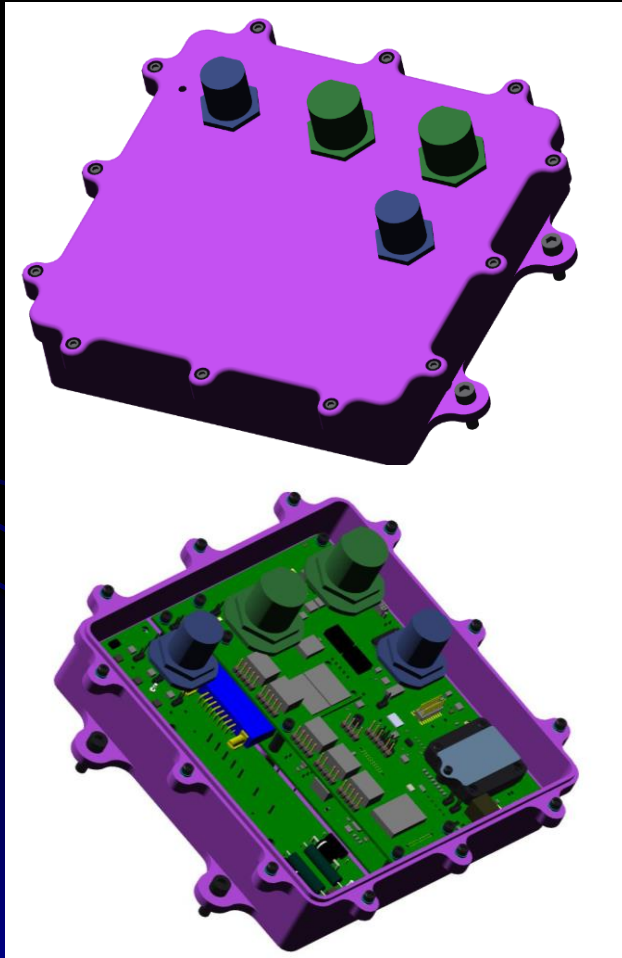
CSM GFE Provided



# HERA Power and Sensor Units



- HERA Power Unit (HPU)
- HERA Sensor Unit (HSU)





# SPE Contingency Plan Scenarios: (Effective Dose due to King '72 SPE)



**Scenario 1:**  
D&E stowage on  
top

Crew1: 114 mSv  
Crew2: 117 mSv  
Crew3: 119 mSv  
Crew4: 113 mSv

**Scenario 3:**  
D&E stowage  
in  
8 boxes on top

Crew1: 109 mSv  
Crew2: 122 mSv  
Crew3: 111 mSv  
Crew4: 106 mSv

**Scenario 4:**  
D&E stowage  
in 16 boxes on  
top

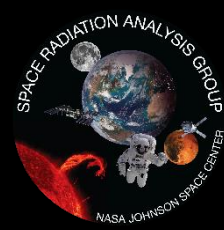
Crew1: 105 mSv  
Crew2: 117 mSv  
Crew3: 106 mSv  
Crew4: 98 mSv

**Scenario 5:**  
18 boxes on top  
and 20 canisters  
in WMS

Crew1: 95 mSv  
Crew2: 110 mSv  
Crew3: 106 mSv  
Crew4: 98 mSv

**Scenario 2:**  
Ideal stowage  
configuration

Crew1: 85 mSv  
Crew2: 102 mSv  
Crew3: 100 mSv  
Crew4: 98 mSv



## A little about Risk

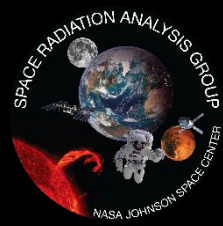
- Calculated quantity – cannot be measured
- Our best knowledge of radiation: tissue interaction, cells damage/repair, mutation, cause of earlier death
- Risks include cancer, CNS, cardiovascular disease
- Based off of limited human statistics (A-Bomb survivors and Nuclear accidents)
- Based off animal exposure studies with relatively high doses
- Extrapolation of mouse models to humans
- Extrapolation of short high doses to low doses over long periods of time

## ERROR BARS ARE BIG

- Deep space missions over ~100days exceed the current NASA limit
- NASA risk limit: 3% REID at upper 95% CI



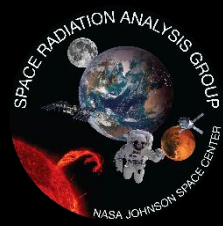
# Basic Radiation Safety and Engineering Problems



- Reduce Time of Exposure
  - Deep space missions lengths are fixed by destination and propulsion system
- Increase Distance from Source
  - Space radiation is ubiquitous
- Increase Shielding
  - Space missions are severely mass limited due to high launch costs
  - Shielding is ineffective against GCR



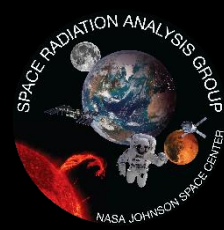
# Role of Dosimetry



- Tools for evaluating risk
  - Environmental models
    - Trapped particles (protons and electrons)
    - GCR (modeled as H through Ni)
    - SPE (probabilistic modeling)
  - Particle transport (HZETRN)
  - Vehicle models and human phantoms
  - Risk models
- Dosimetry provides “anchors” for the tools above
  - RAM/CPD data used to determine relative contributions of GCR and trapped protons for ISS
  - MPCV missions: HERA + RAM/CPD



# Risk Models

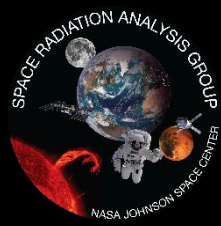


- Evolution of radiation risk models at NASA
  - Pre-2012: effective dose approach
    - LSS mortality-based risk coefficients
    - Gender and age dependency
    - Leukemia and solid cancer risks
  - NSCR 2012\*: organ risk approach
    - Risks for organs tracked separately
    - LSS incidence-based risk coefficients
      - UNSCEAR
      - BEIR VII
      - Preston et al
    - Never-smoker status
  - Future\*: normal weight population, NTE, circulatory effects, CNS effects, etc.

*\*Cucinotta, Kim, and Chappell. Space Radiation Cancer Risk Projections and Uncertainties – 2012. NASA/TP-2013-217375.*



# Quality Factor



- Pre-2012
  - Function of LET in water
  - ICRP 26 then ICRP 60
- NSCR-2012\*: NASA QF
  - Track structure-based model
  - Function of  $Z$ ,  $\beta$
  - Various adjustable parameters informed by radiobiology data
  - Important contributor to uncertainty

*\*Cucinotta, Kim, and Chappell. Space Radiation Cancer Risk Projections and Uncertainties – 2012. NASA/TP-2013-217375.*





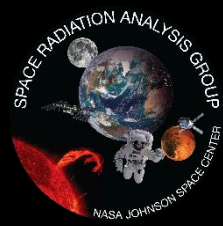
# Uncertainty Assessment



- NASA risk limit is inherently statistical
- Monte Carlo sampling used to determine overall uncertainty distribution
- Subjective probability density functions
  - NCRP Report No. 126
  - Risk models
  - DDREF (Bayesian analysis)
  - Quality factor



# Summary



- SRAG Operations for future exploration missions need additional forecasting capability due to being unprotected in free space
- The ConOps for a radiation contingency event on MPCV is established and the details are being worked. Protection is possible below required limit without flying any parasitic mass
- MPCV radiation instrumentation is being built
- Risk is the quantity that the NASA radiation requirement is built on. Better understanding of biological effects of radiation are needed to reduce the error