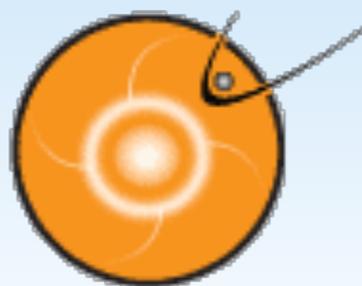


Update on Empirical Models At CCMC

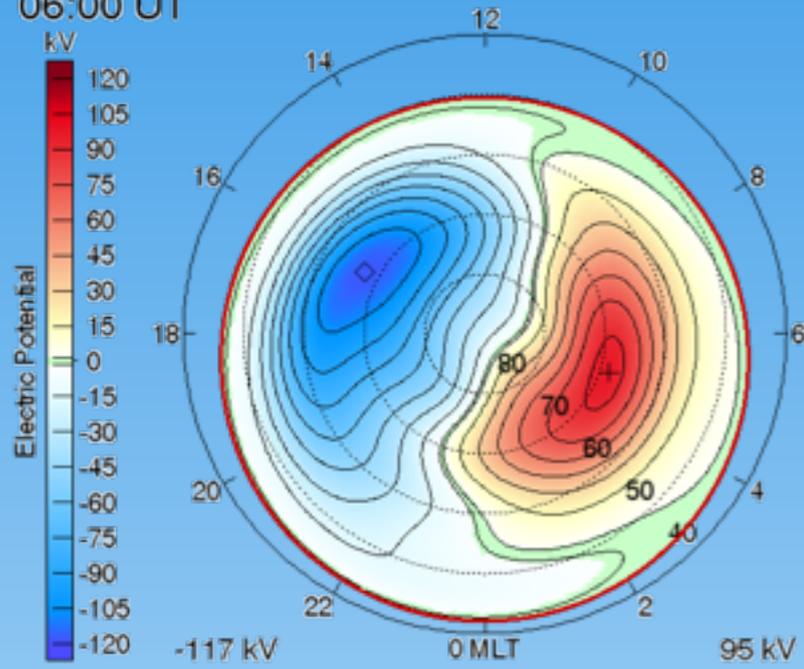
Daniel R. Weimer
Research Professor



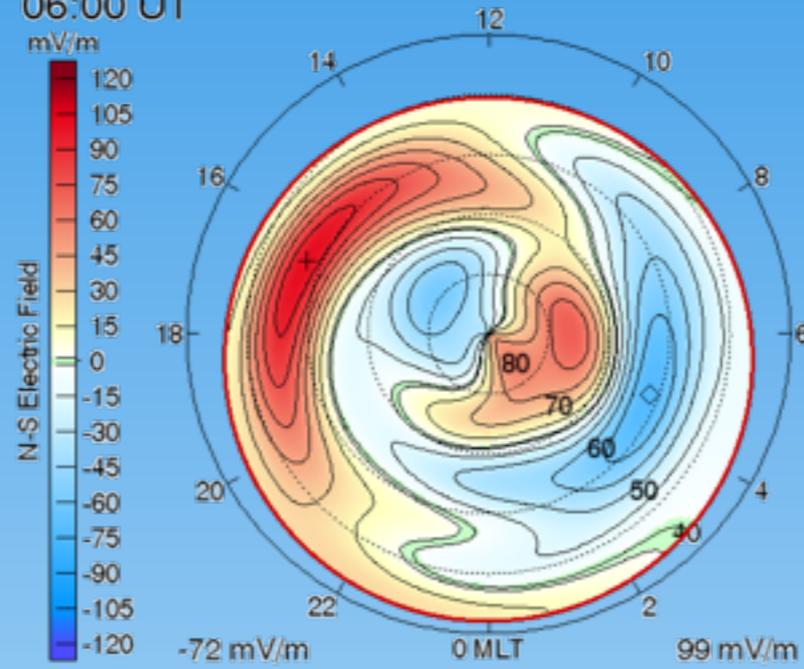
COMMUNITY
COORDINATED
MODELING
CENTER

7th CCMC Community Workshop
March 31– April 4, 2014
Annapolis, MD

8 Nov. 2004
06:00 UT

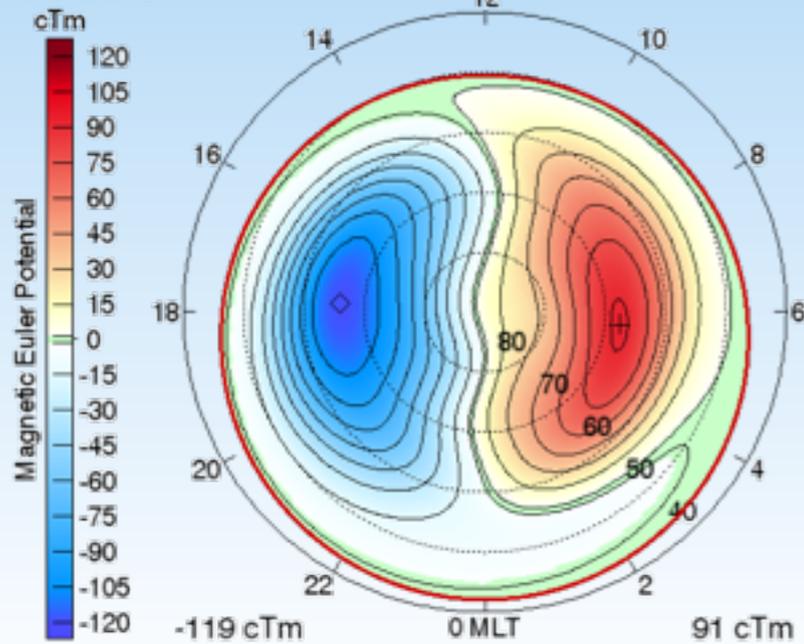


8 Nov. 2004
06:00 UT

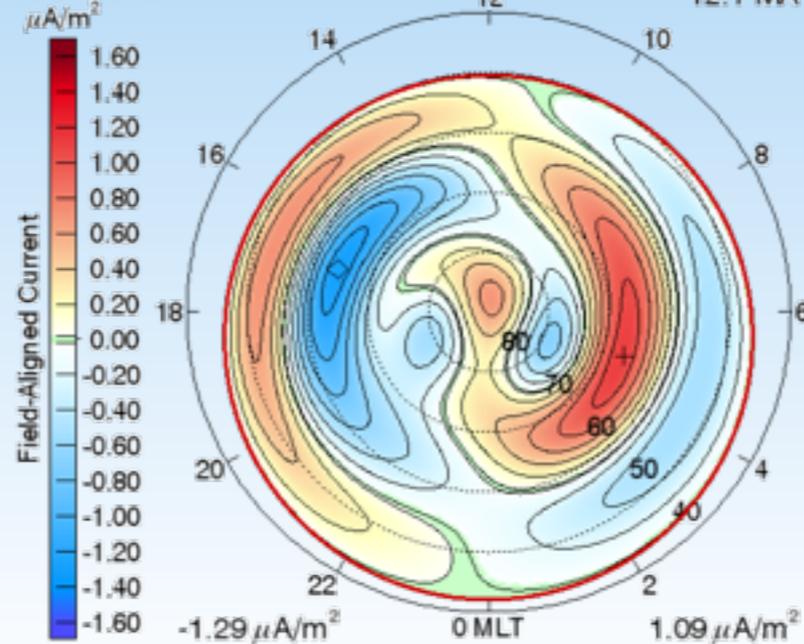


$$\mathbf{E} = -\nabla_S \Phi$$

8 Nov. 2004
06:00 UT



8 Nov. 2004
06:00 UT



$$J_{\parallel} = \nabla_S^2 \psi / \mu_o$$

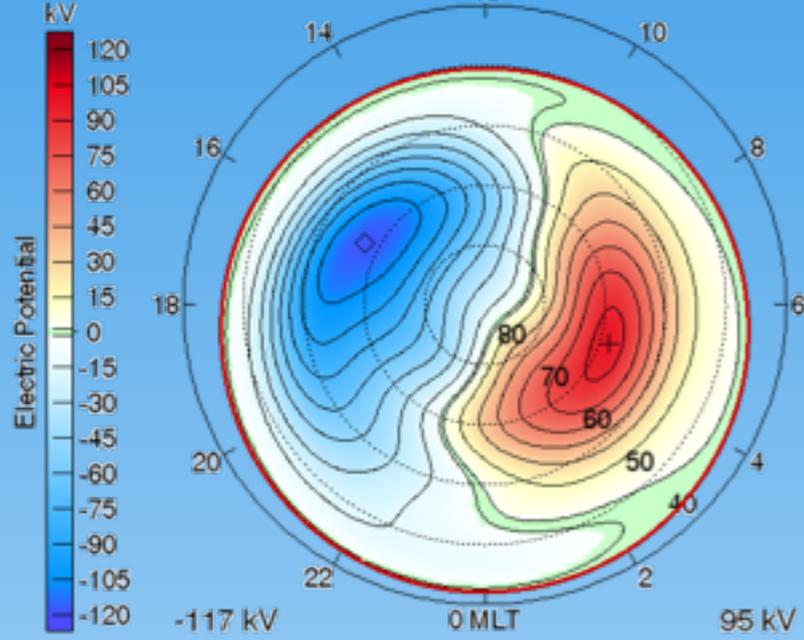
Brief Review:

The 2005 model included both electric potential and magnetic Euler potential.

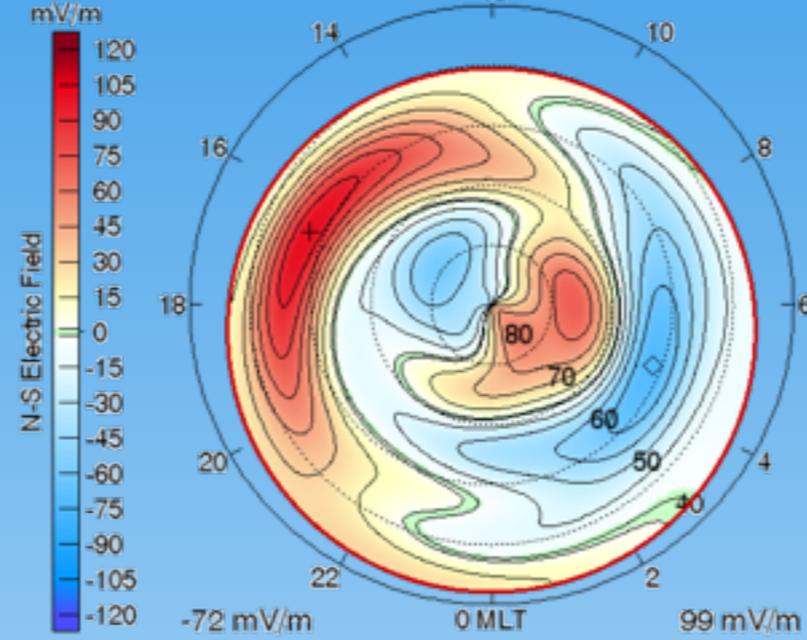
The magnetic potential model was first created in order to map field-aligned currents, without use of the “infinite current sheet” assumption. Conductivity variations in ionosphere are implicitly included in the magnetometer measurements.

These models use “spherical cap harmonic analysis” (SCHA).

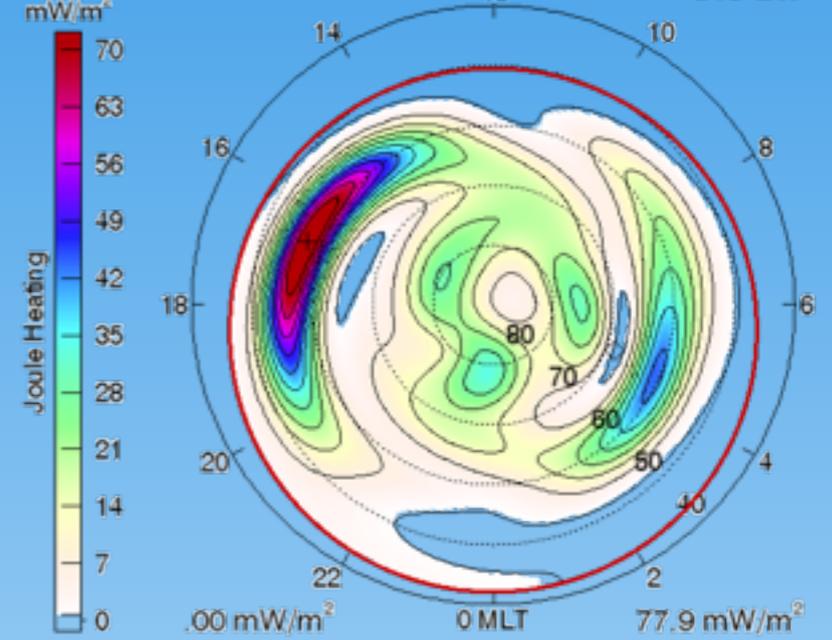
8 Nov. 2004
06:00 UT



8 Nov. 2004
06:00 UT



8 Nov. 2004
06:00 UT



$$\mathbf{E} = -\nabla_S \Phi$$

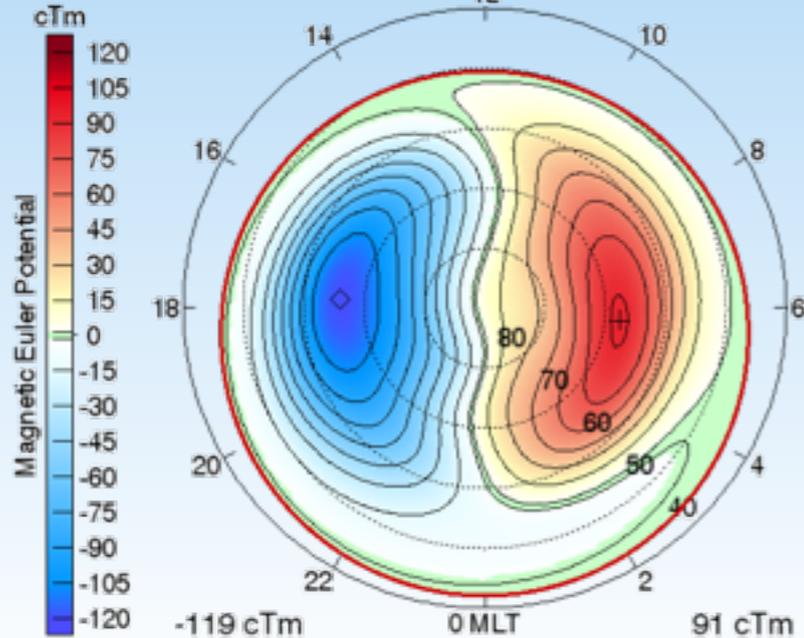
$$\mathbf{S} = \mathbf{E} \times \Delta \mathbf{B} / \mu_o$$

Although developed for a FAC model, the magnetic potentials have been even more useful, in combination with the electric potentials, to obtain the Poynting flux

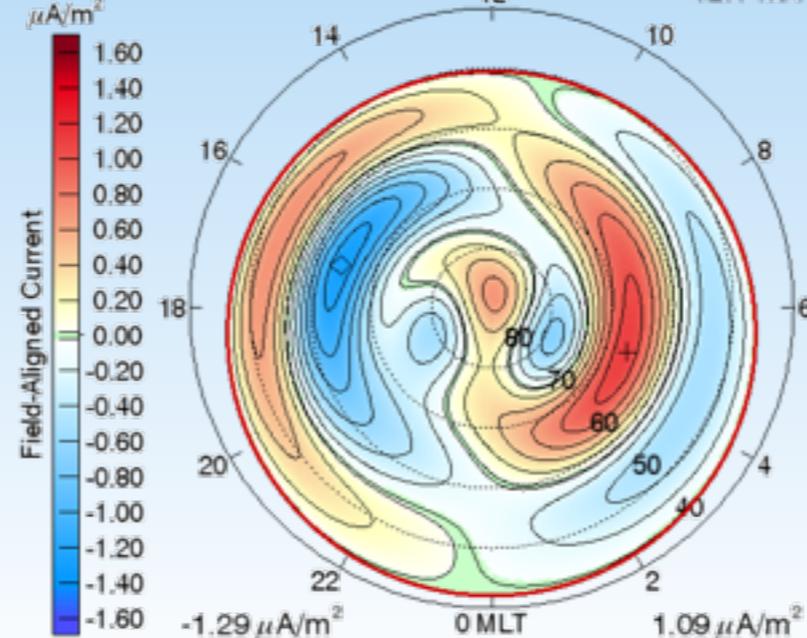
$$J_{\parallel} = \nabla_S^2 \psi / \mu_o$$

$$\Delta \mathbf{B} = \hat{\mathbf{r}} \times \nabla_S \psi$$

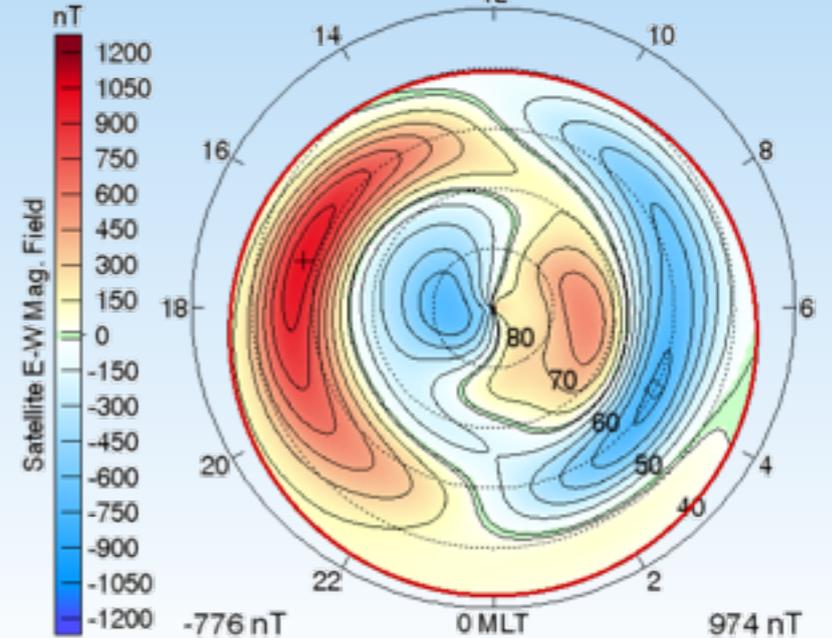
8 Nov. 2004
06:00 UT



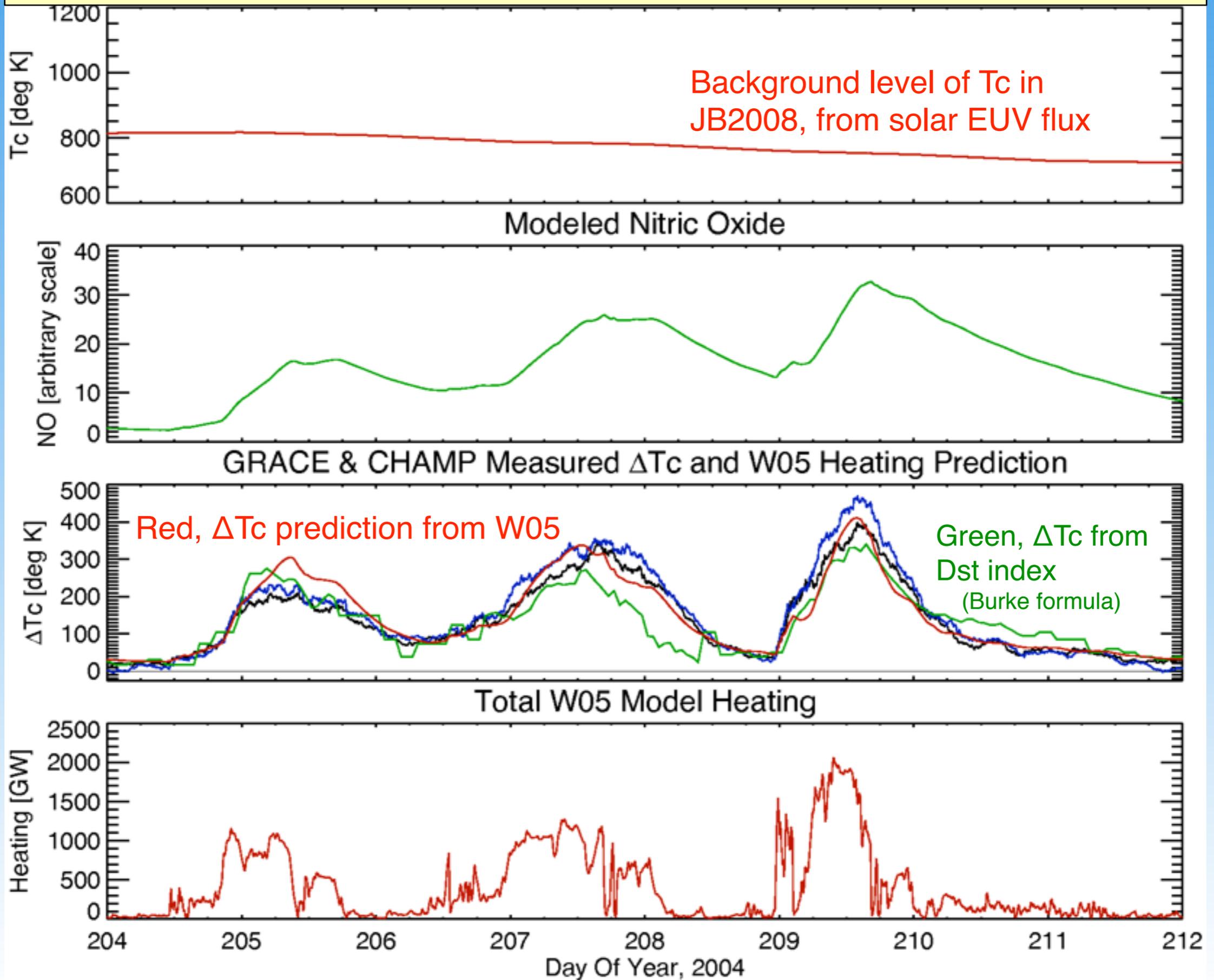
8 Nov. 2004
06:00 UT



8 Nov. 2004
06:00 UT



Poynting flux from W05 model compared with change in average exospheric temperature, derived from CHAMP/GRACE densities



The ΔT_c prediction technique:

$$\Delta T_c(t_{n+1}) = \Delta T_c(t_n) \left(1 - \frac{\Delta t}{\tau_c}\right) + \beta H_J \Delta t$$

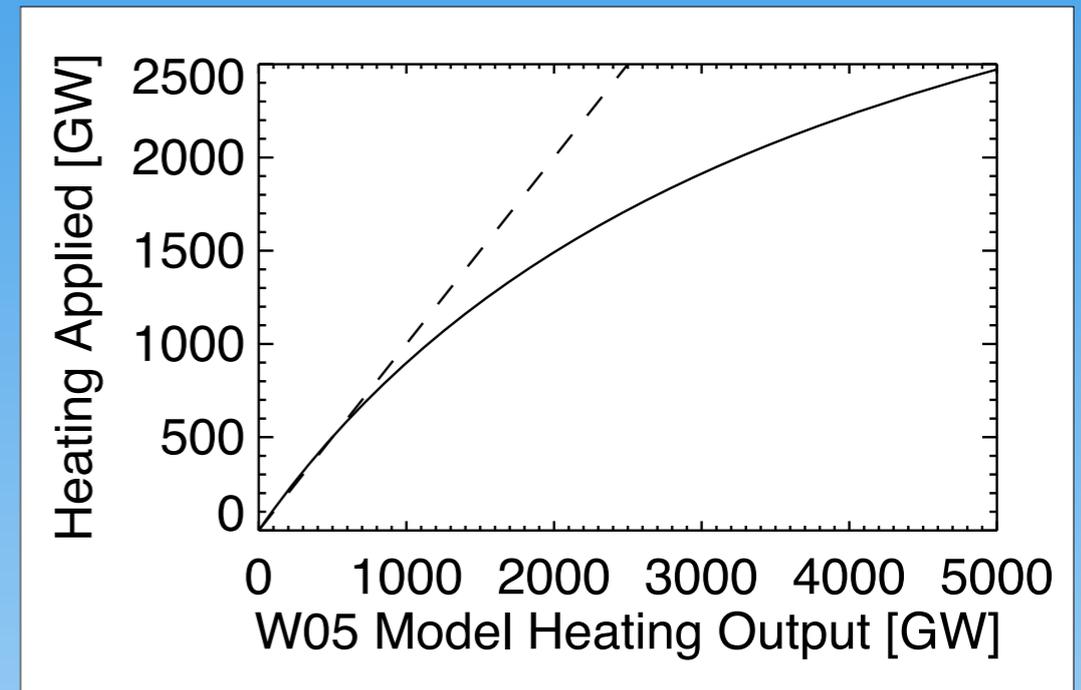
$$\beta = 6.9 \cdot 10^{-4} \text{ (}^\circ\text{K/GW-min)}$$

$$\tau_c = 14.6 \text{ (hours)} - 0.281 NO$$

$$NO(t_{n+1}) = NO(t_n) \left(1 - \frac{\Delta t}{\tau_{NO}}\right) + \gamma H_J \Delta t$$

$$\gamma = 2.5 \cdot 10^{-5} \text{ (units/GW-min)}$$

$$\tau_{NO} = 28.0 \text{ (hours)}$$

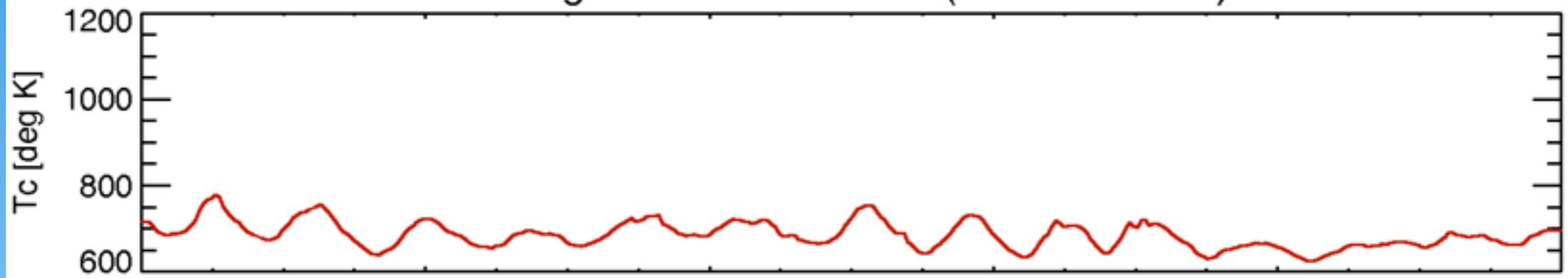


Saturation curve applied to heating

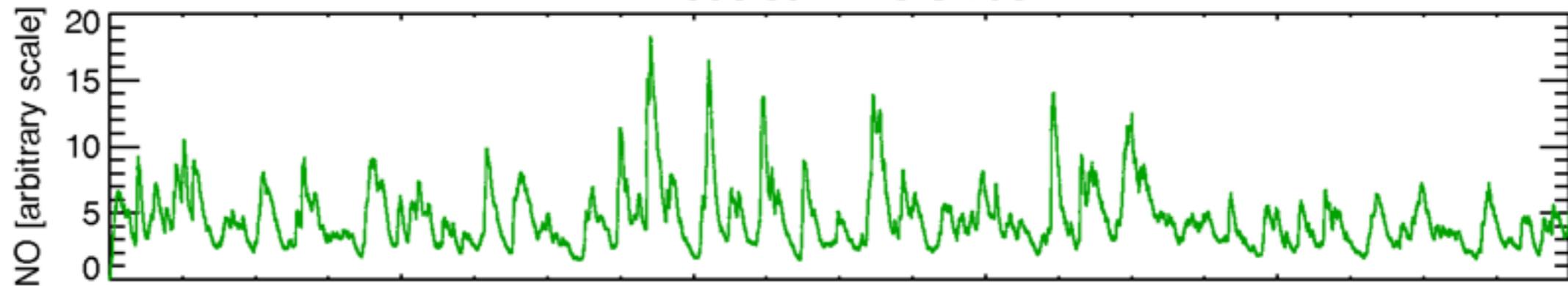
H_J is total Joule heating/Poynting flux from W05 model, with additional saturation applied. All constants (except γ) obtained by fitting five years of H_J with CHAMP and GRACE measurements of ΔT_c .

A program that does this calculation has been delivered to the CCMC.

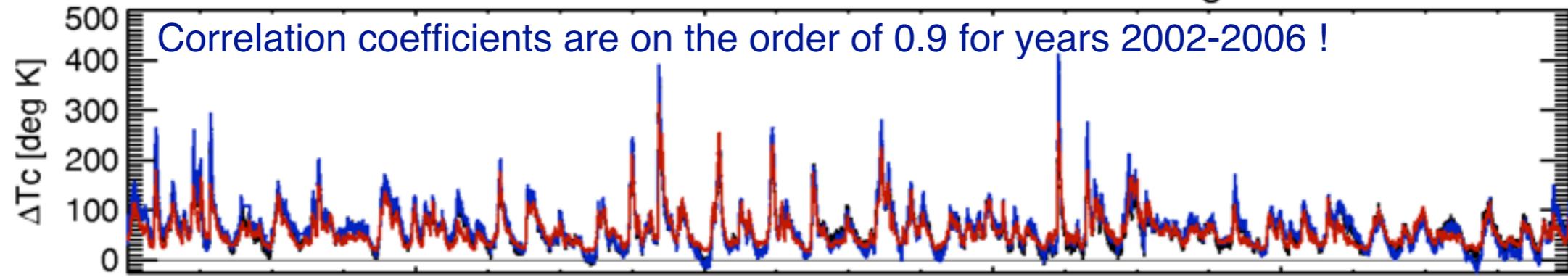
Background Tc in JB2008 (From EUV flux)



Modeled Nitric Oxide

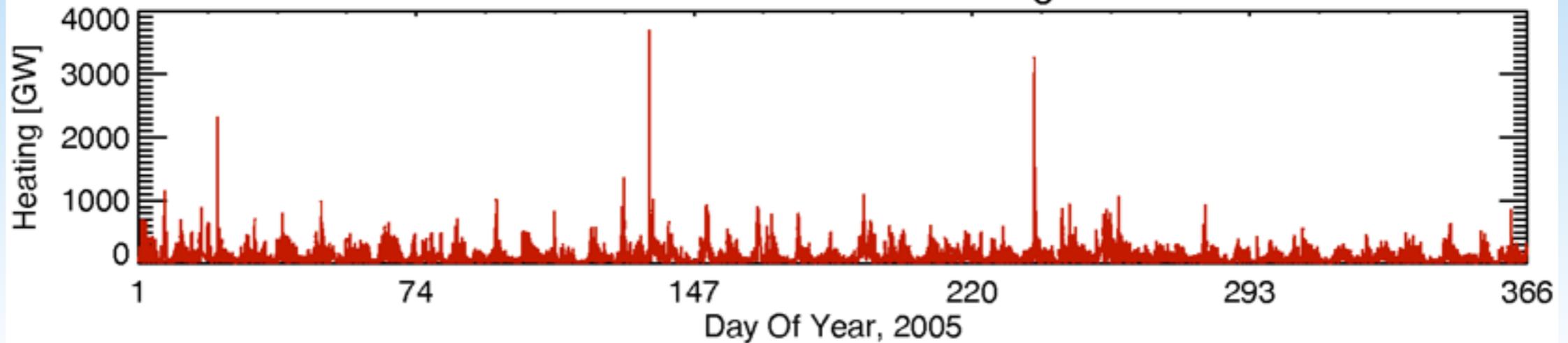


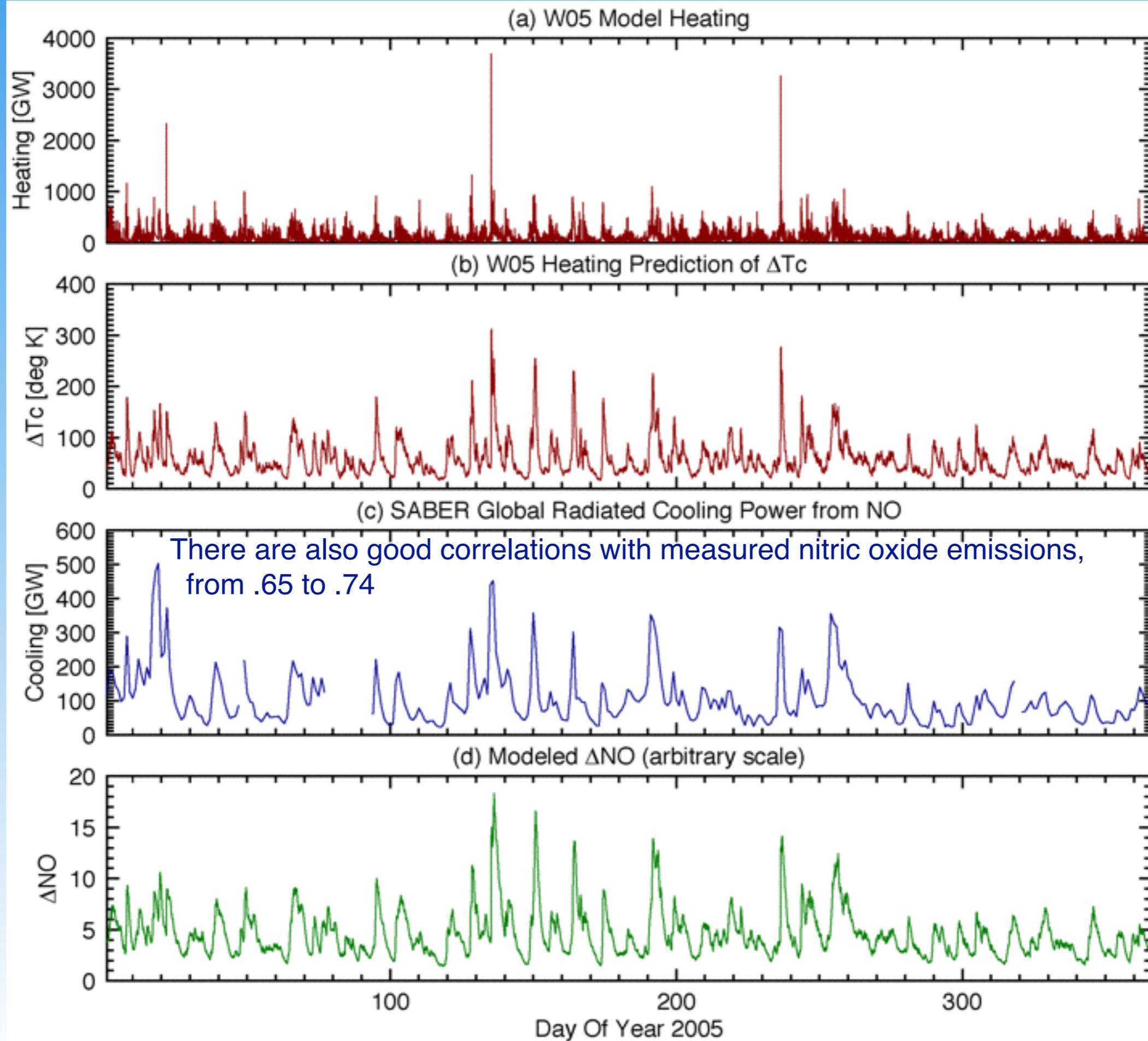
GRACE & CHAMP Measured ΔT_c and W05 Heating Prediction



Correlation coefficients are on the order of 0.9 for years 2002-2006 !

Total W05 Model Heating



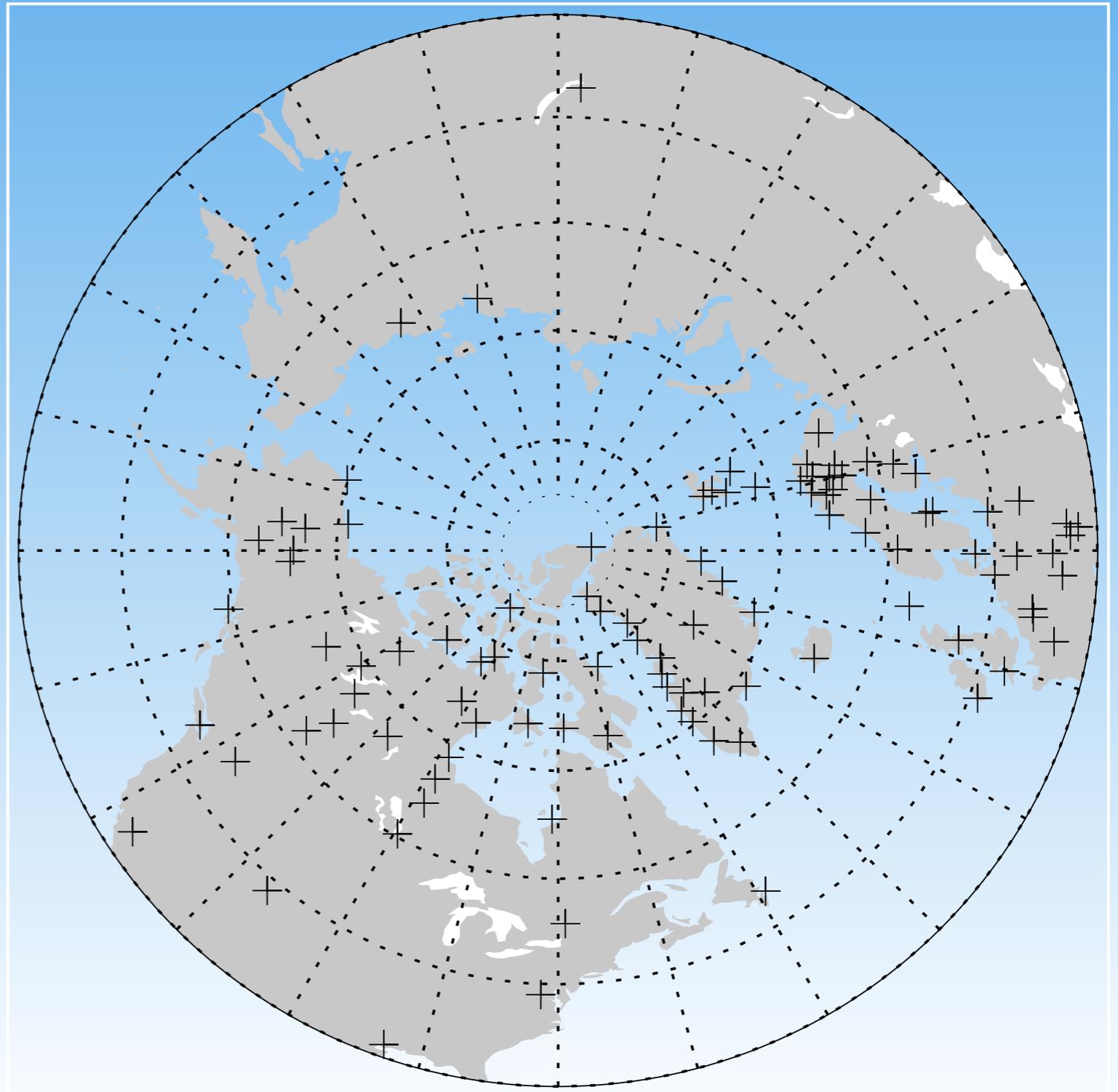


New model for prediction of geomagnetic perturbations

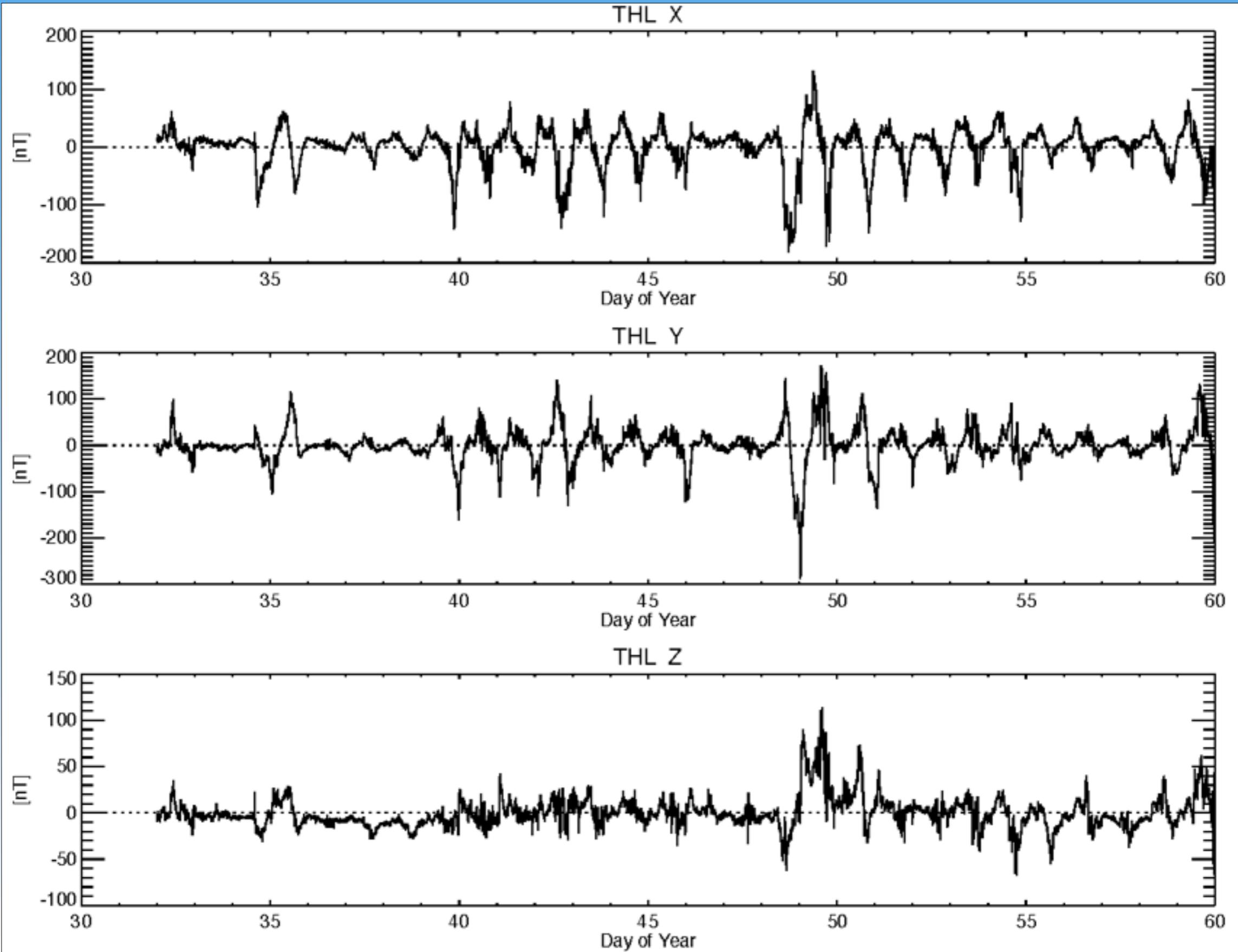
Reference: Weimer, D. R. (2013), An empirical model of ground-level geomagnetic perturbations, *Space Weather*, 11, 107–120, doi:10.1002/swe.20030

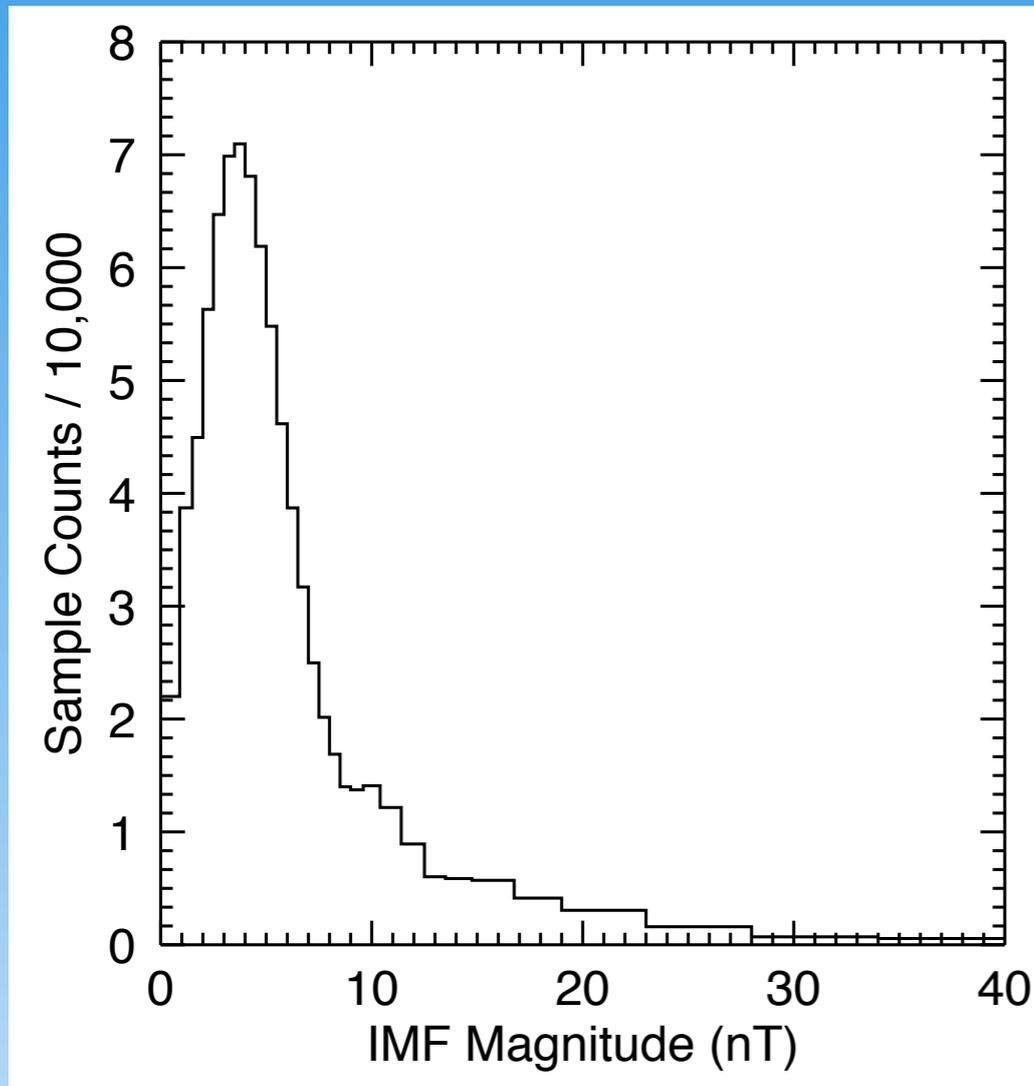
Uses data from >140 magnetometer stations in Northern hemisphere, over an 8-year period (1998-2005), solar wind & IMF, and $F_{10.7}$.

Effects of conductivity variations and induced, underground currents are implicitly included.



Example of processed data, spanning one month ($\approx 13,000$ of these were examined for quality control)





The data are divided into 29 bins, sorted according to IMF magnitude. The width of each bin increased above 9 nT, yet there are few samples in the highest bins.

Model coefficients are derived using a least-error fit for each bin, including an over-lap of data from the adjacent bins.

Each vector component is fit separately, using spherical harmonics on a 90° cap, in corrected geomagnetic apex coordinates.

Only even l - m combinations of the Legendre polynomials are used.

To obtain a smooth response curve, the C_n coefficients are interpolated, given the IMF magnitude.

$$\Delta B_X(\Lambda, \varphi) = \sum_{l=0}^{31} \sum_{m=0}^{3<l} P_l^m(\cos \Lambda) (g_k^m \cos m\varphi + h_k^m \sin m\varphi)$$

$$g_k^m = c_0 + c_1 B_T + c_2 V_{SW} + c_3 t + c_4 \sqrt{F_{10.7}} +$$

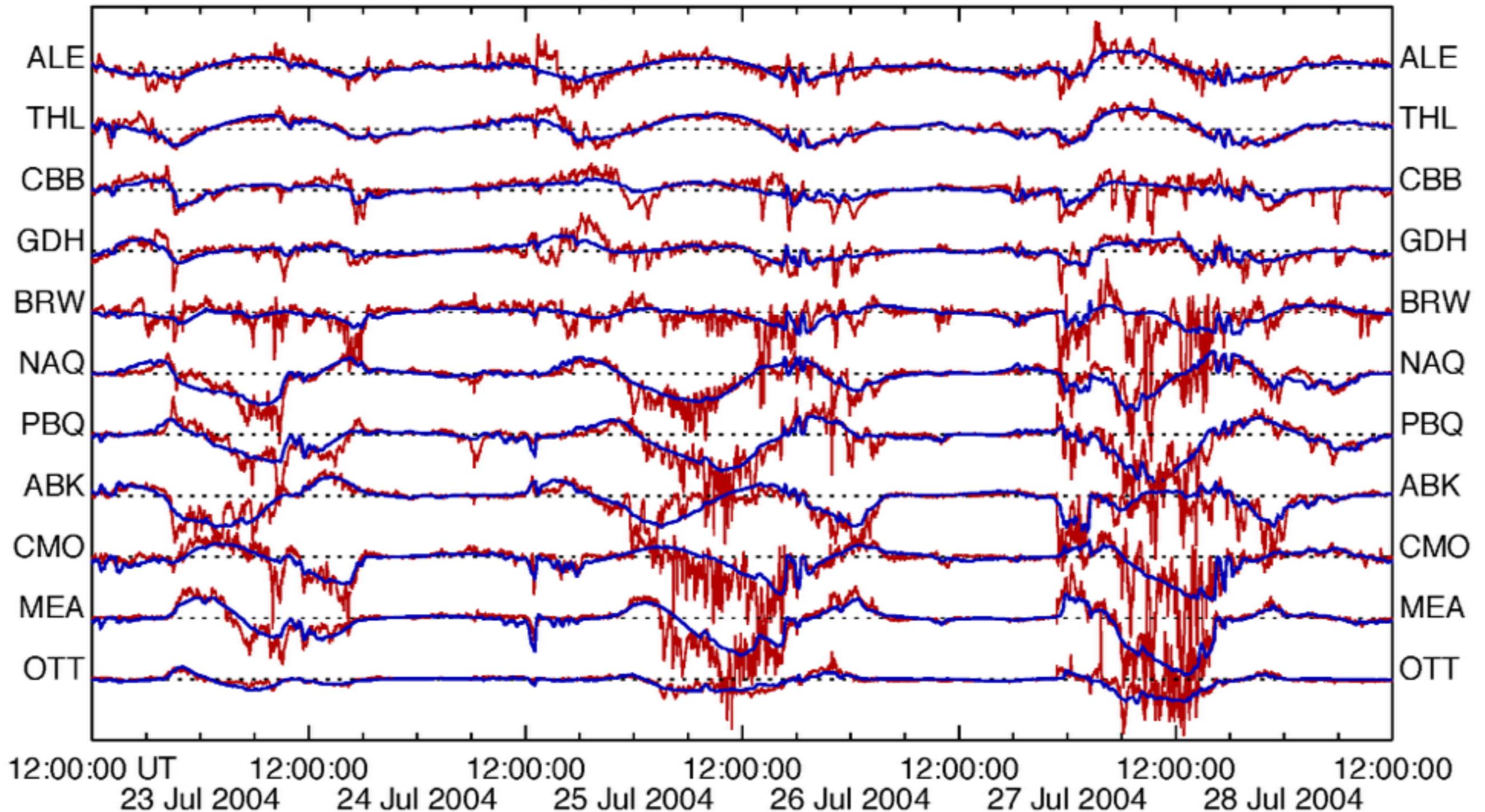
$$c_5 B_T \cos(\theta_c) + c_6 V_{SW} \cos(\theta_c) + c_7 t \cos(\theta_c) + c_8 \sqrt{F_{10.7}} \cos(\theta_c) +$$

$$c_9 B_T \sin(\theta_c) + c_{10} V_{SW} \sin(\theta_c) + c_{11} t \sin(\theta_c) + c_{12} \sqrt{F_{10.7}} \sin(\theta_c) +$$

$$c_{13} B_T \cos(2\theta_c) + c_{14} V_{SW} \cos(2\theta_c) + c_{15} B_T \sin(2\theta_c) + c_{16} V_{SW} \cos(2\theta_c)$$

The model does very well at prediction of ΔB levels, not so well on the superposed and random, higher frequency variations. Substorms are also not modeled.

ΔB North-South (X), 1200 nT between base lines



Incidental Note About Metrics

Presently used metrics for $\delta B/\delta t$ and Regional-K both ignore the sign of ΔB , as well as overall agreement with ΔB level:

$$dB / dt = \sqrt{(dB_x / dt)^2 + (dB_y / dt)^2}$$

K is calculated from the maximum “Range” of ΔB in the two horizontal directions, over three hours =

$$\max[(\Delta B_{x,\max} - \Delta B_{x,\min}), (\Delta B_{y,\max} - \Delta B_{y,\min})]$$

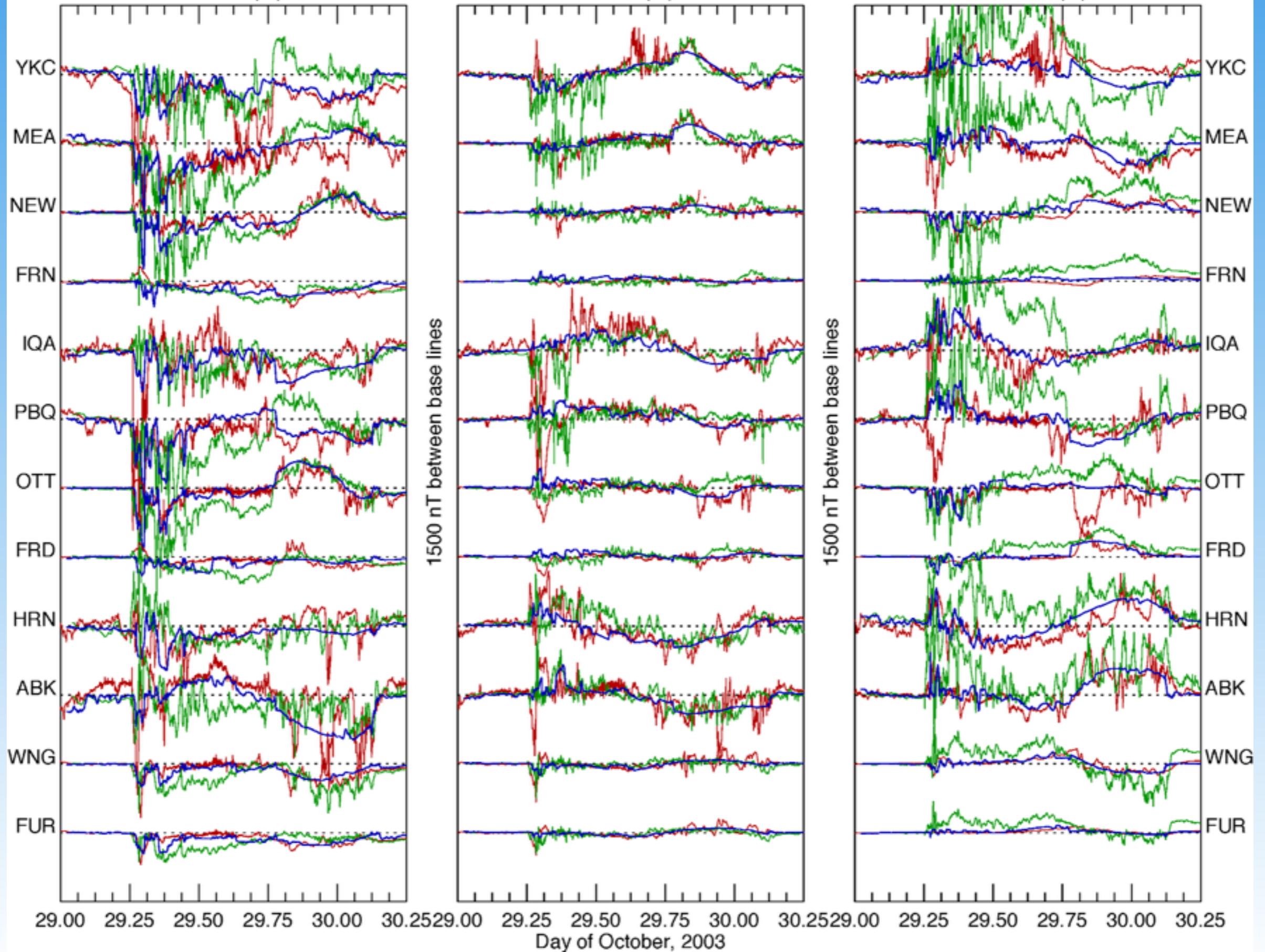
As a result, both metrics can result in “good,” high scores for predictions of ΔB having signs opposite to the actual ΔB , or very wrong magnitudes, as these metrics only test changes in ΔB during certain time intervals, and not the level.

With the present metrics for $\delta B/\delta t$ and Regional-K, it may be possible to do just as well with output from a pseudo-random noise generator, that is added to the more smoothly varying ΔB prediction, scaled to the predicted level.

BLUE: MODEL 6_WEIMER
North (X)

GREEN: MODEL 9_SWMF
East (Y)

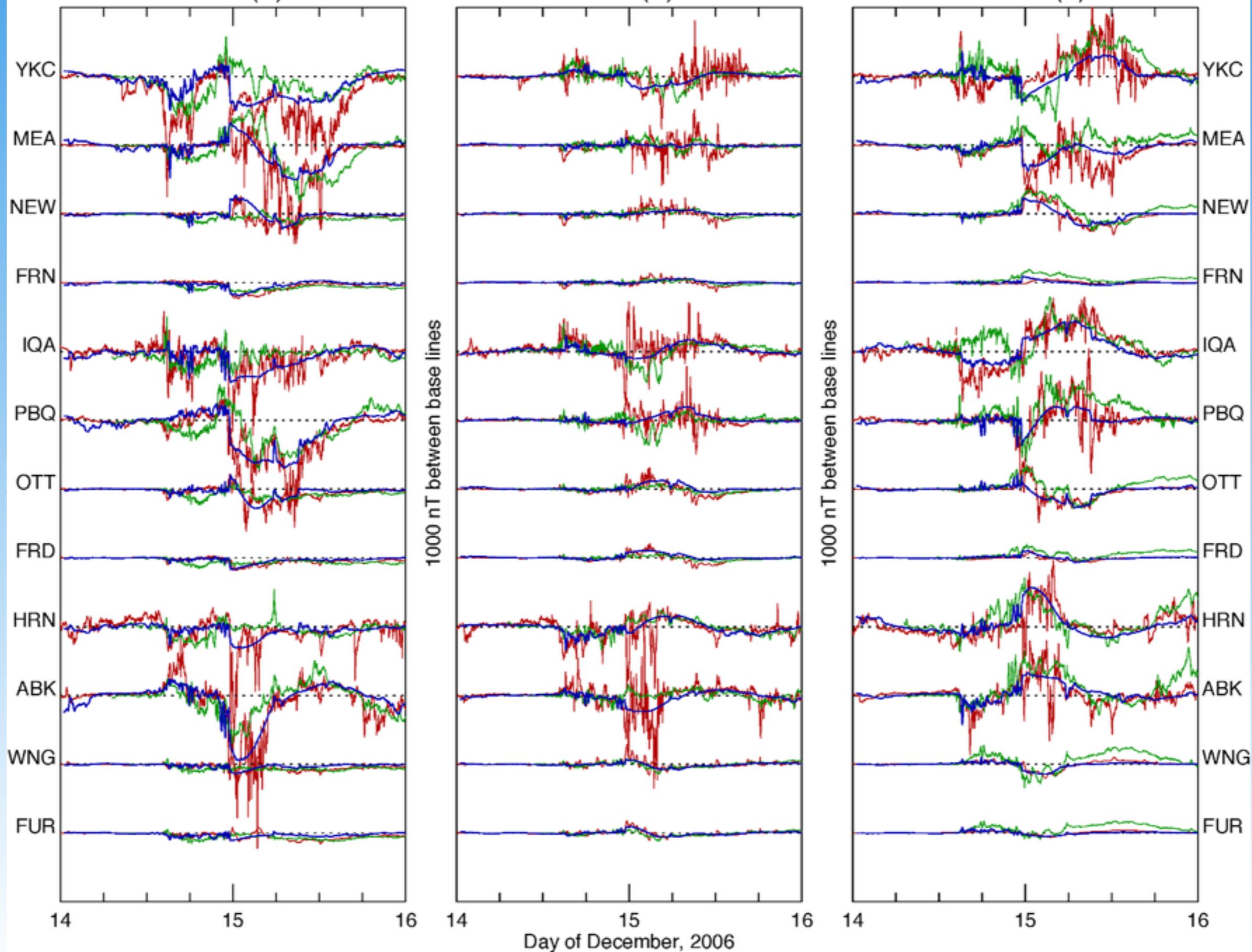
RED: DATA
Vertical (Z)



BLUE: MODEL 6_WEIMER
North (X)

GREEN: MODEL 9_SWMF
East (Y)

RED: DATA
Vertical (Z)

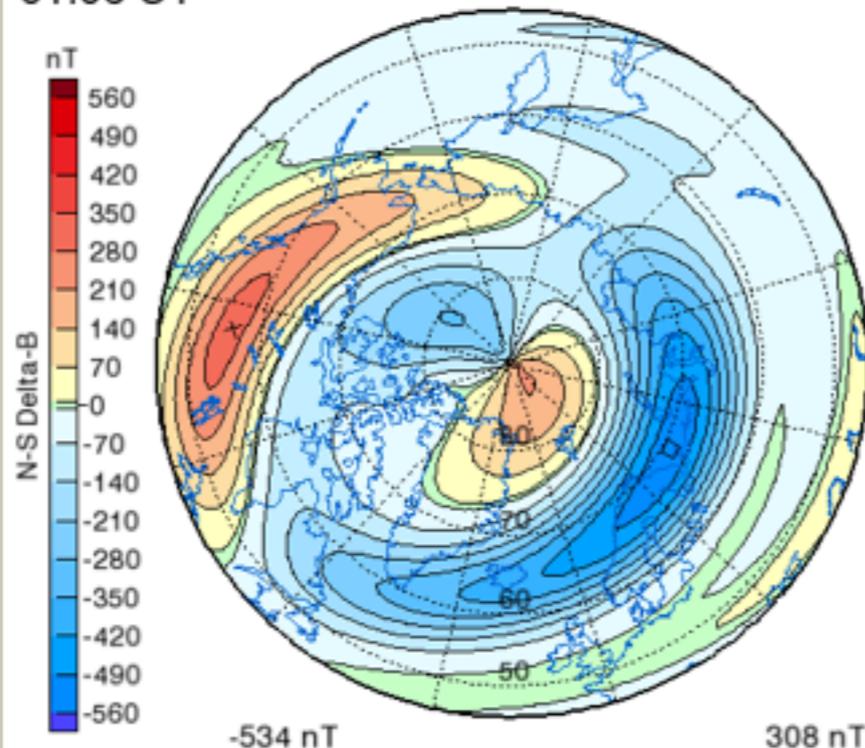


Real-time maps are shown at <http://mist.nianet.org/weimerGeomag.html>
Operating continuously since 2011; plots archived since Sept. 2012



Real-Time Space Weather Predictions Geomagnetic Perturbations at Ground Level

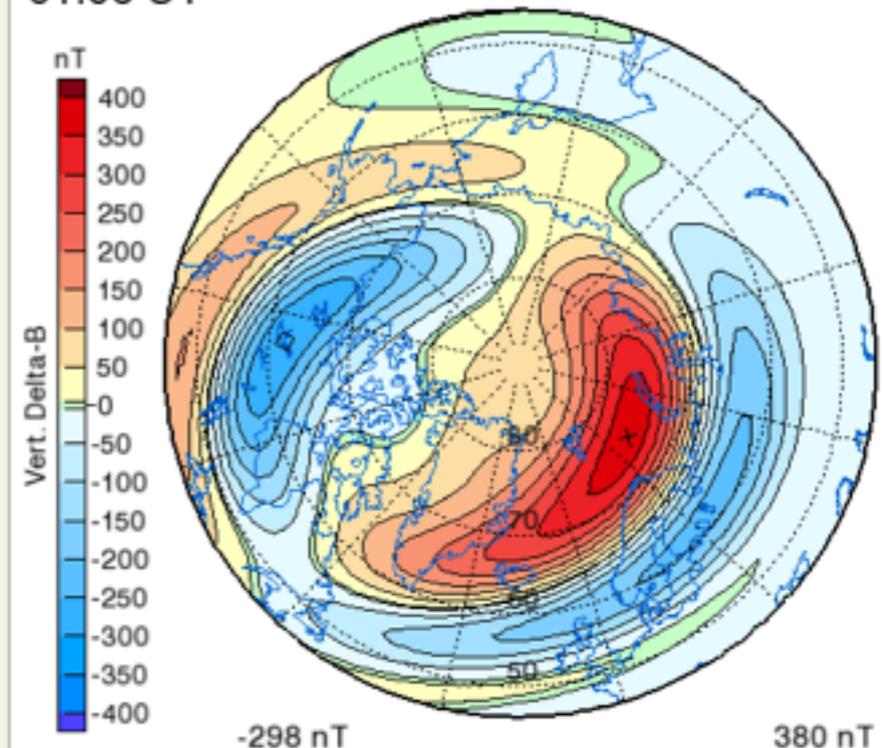
30 April 2011
01:06 UT



North-South Magnetic Perturbations at Surface

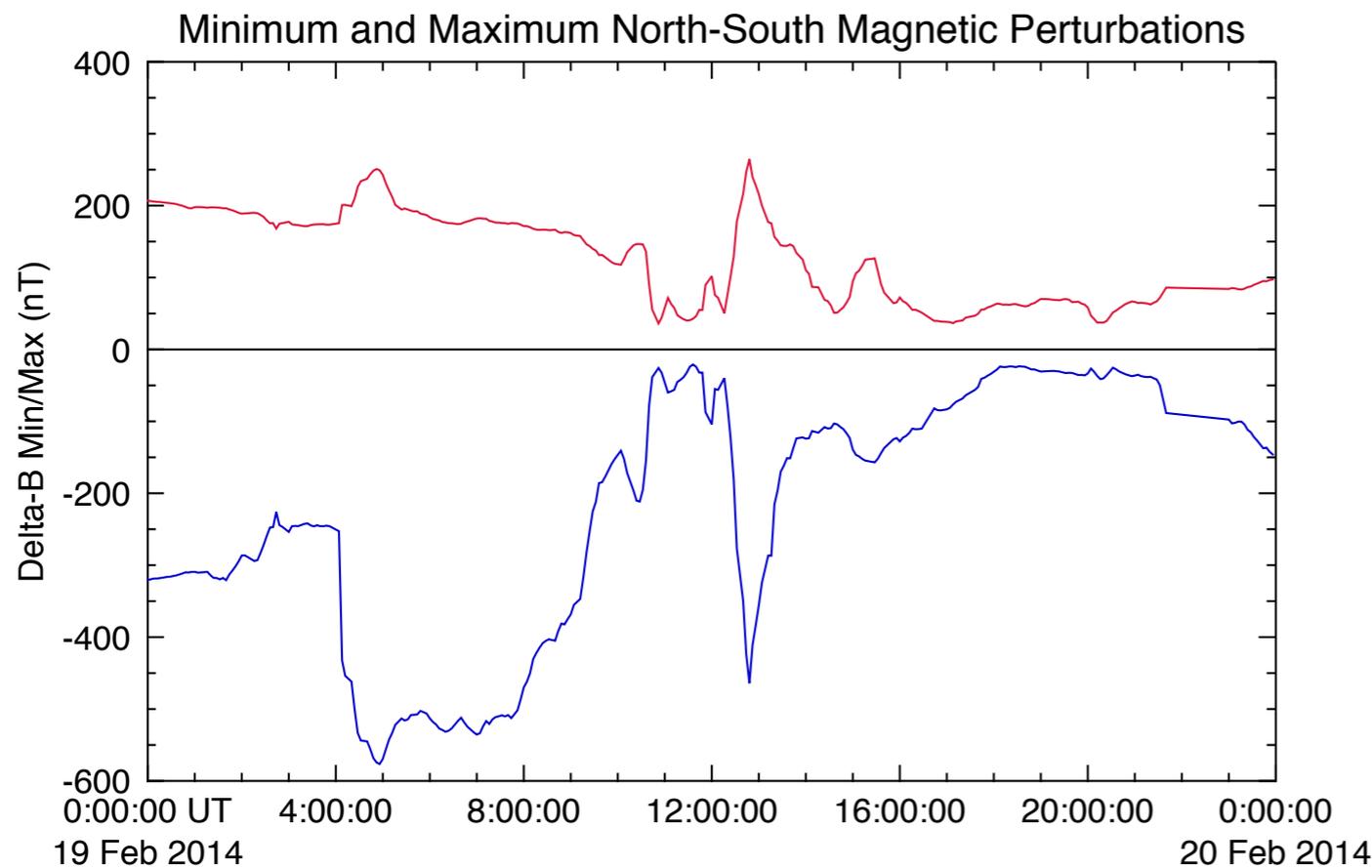
[Download PDF file](#)

30 April 2011
01:06 UT



Vertical Magnetic Perturbations at Surface

[Download PDF file](#)



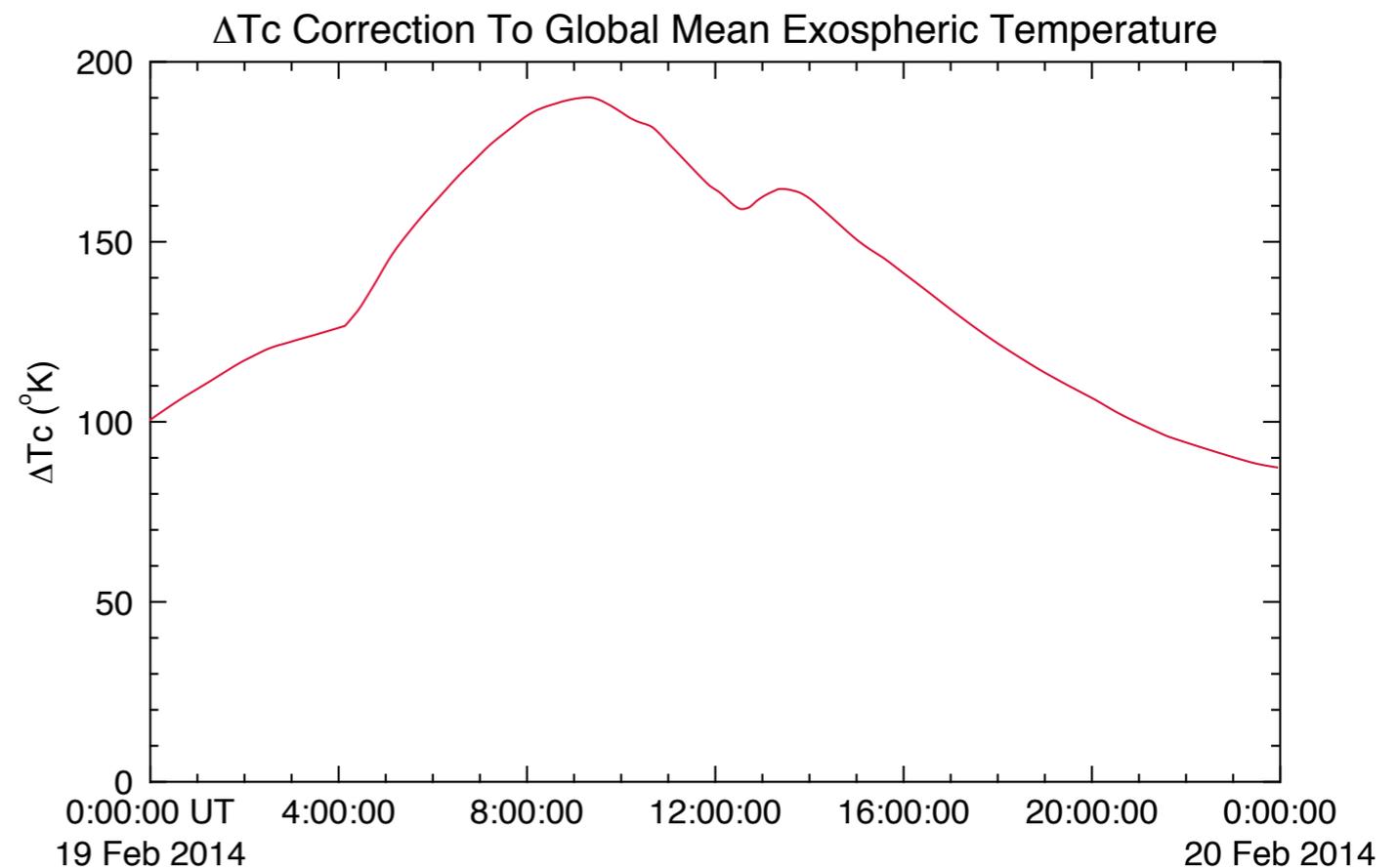
Recent addition: Daily summary plots, at <http://mist.nianet.org/weimerDaily.html>

Includes min/max ΔB -North, polar cap potential drop, total heating, and real-time exosphere temperature correction, ΔT_C

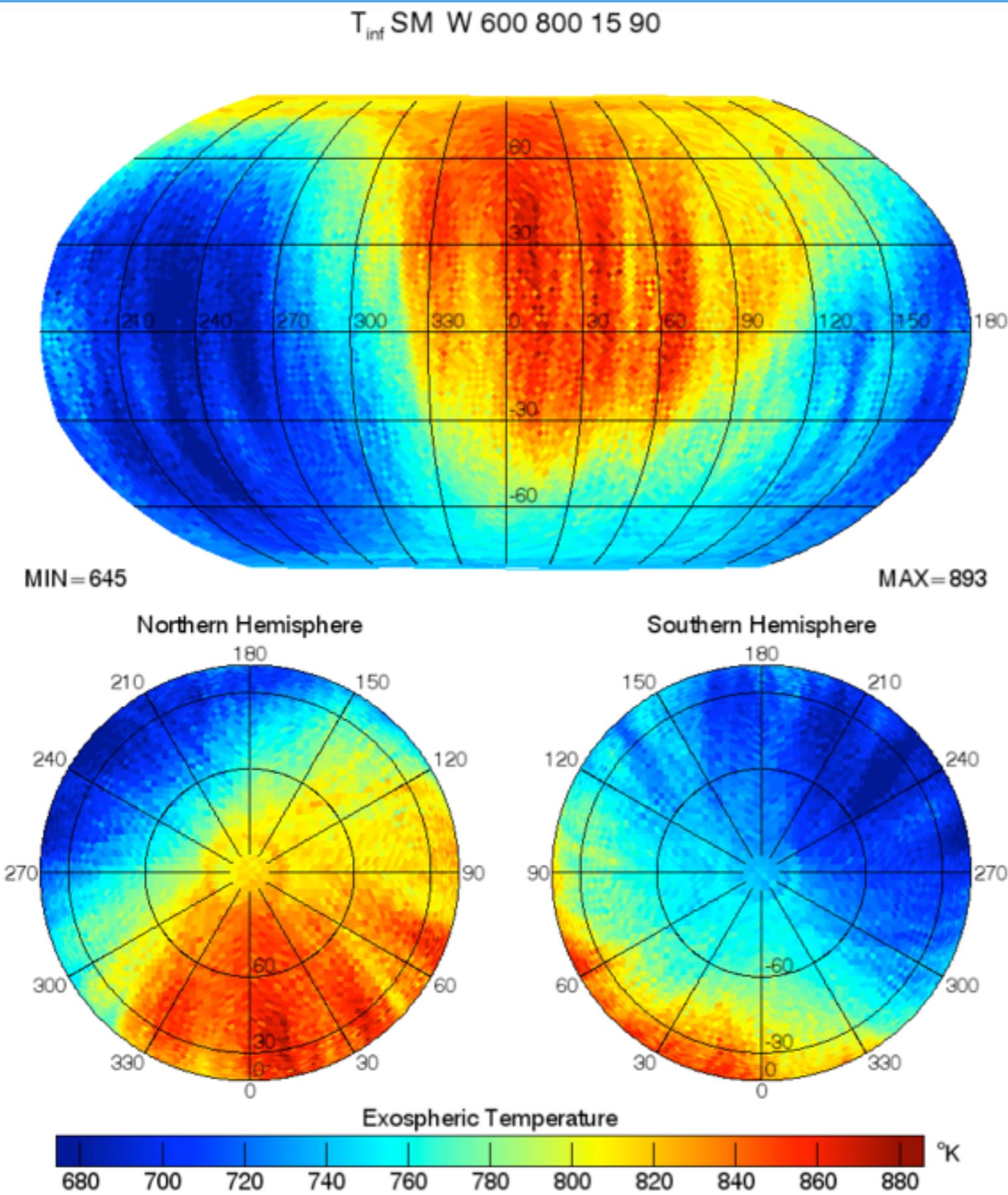
Future models in development:

New FAC model, using data from magnetometers on Oersted, CHAMP, and Swarm.

New thermosphere model



Example of Global Map from CHAMP and GRACE data



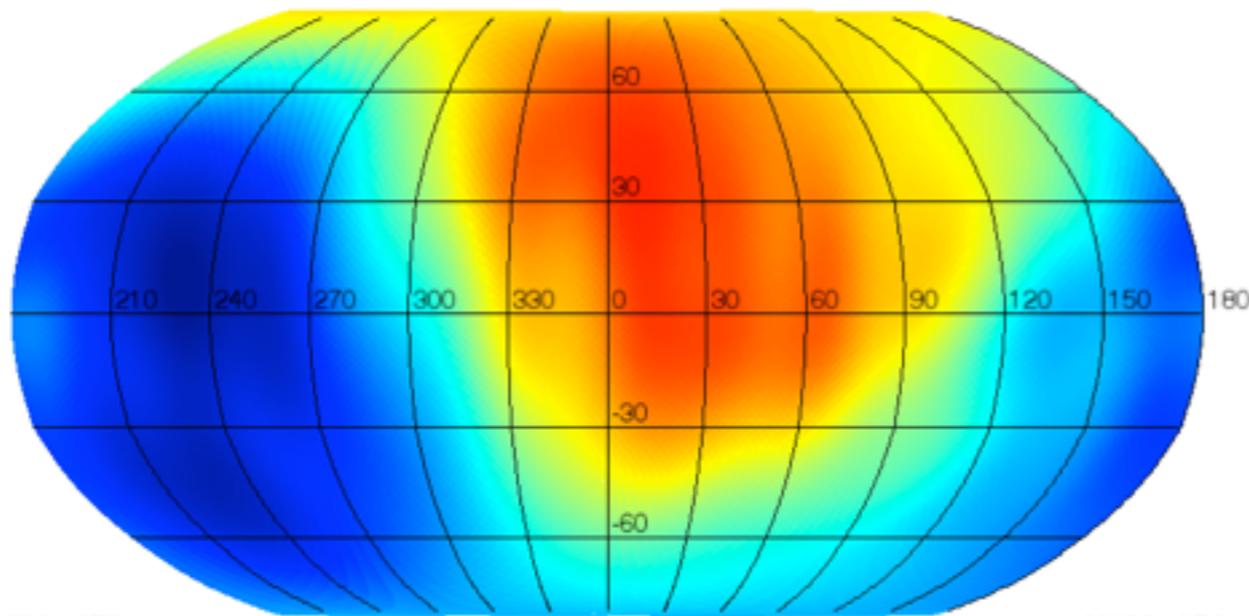
- Years 2002—2006
- “HEALPix” Equal-Area Grid has 12,288 cells
- Means in each cell shown, from times with subsolar lat. $> 15^\circ$
- JB2008+W05 models predicts that $600 < T_{\infty\text{min}} < 800$ °K
- Locations in Solar Magnetic coordinates (SM), rather than geographic
- Resolves density/temperature perturbations in auroral ovals
- Meridional bands are side-effect of 27-day period in solar activity, while satellites precess in local time

Wavelet Filtering Applied to Same Data, Compared with JB2008

Smoothing with “Undecimated Wavelet Transform” at 5 scale levels, first two scales set to zero

JB2008 Model at Summer Solstice, and Minimum Temperature T_c Set to 700 °K

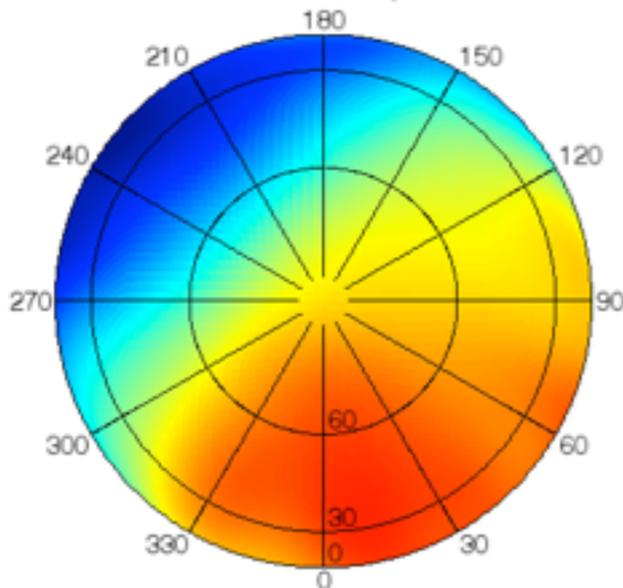
T_{inf} SM W 600 800 15 90
WT: NBRSCALE=5 ZMAX=2



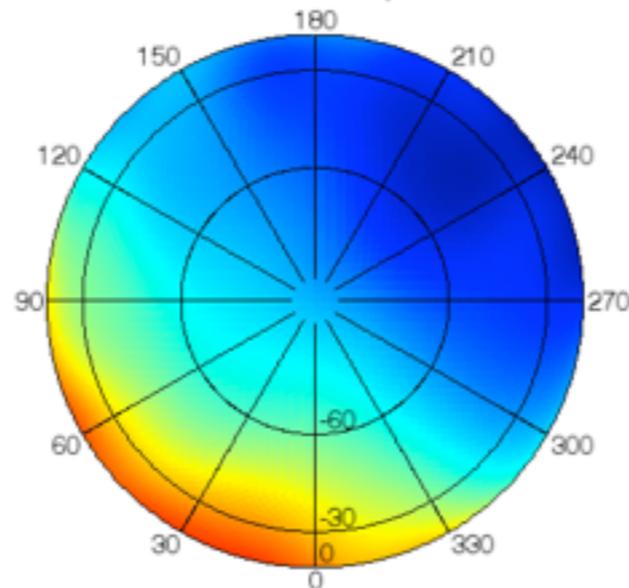
MIN=678

MAX=856

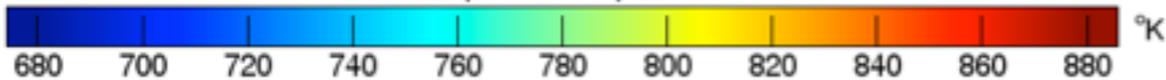
Northern Hemisphere



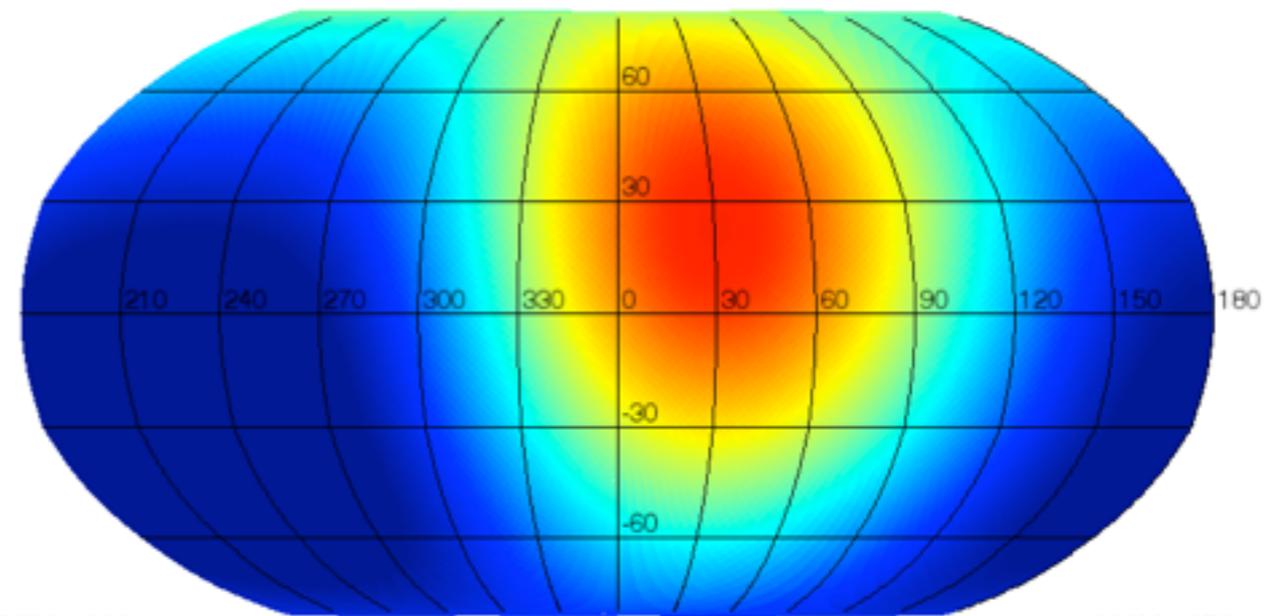
Southern Hemisphere



Exospheric Temperature



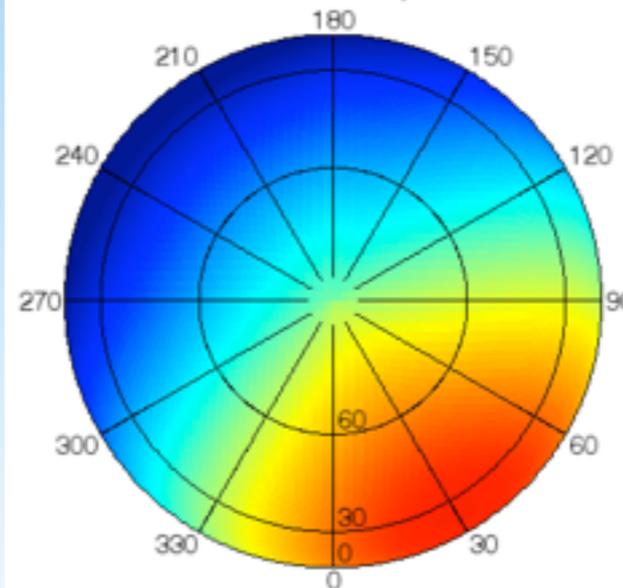
Exospheric Temperature, Derived from JB2008 Model Densities
21 Jun 2012 12:00 UT Fixed $T_c=700$ $DT_c=0$



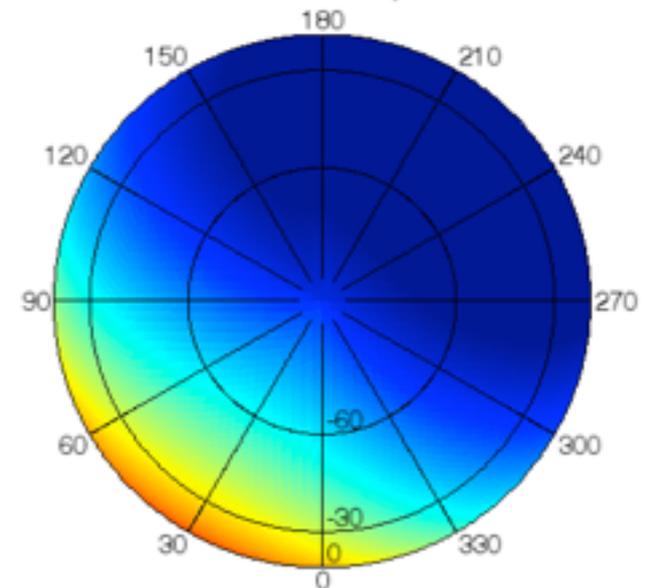
MIN=660

MAX=858

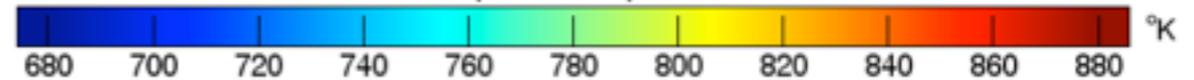
Northern Hemisphere



Southern Hemisphere



Exospheric Temperature



Acknowledgments

Prior model work supported by NSF, NASA, and AFRL. The thermosphere density prediction supported by NASA LWS Program grant #NNX09AJ58G, and now Geospace #NNX13AD73G to Virginia Tech.

The author thanks Eric Sutton for providing calibrated CHAMP and GRACE measurements of neutral density. Bruce Bowman and Kent Tobiska provided the code for the JB2008 model and solar indices. TIMED SABER measurements of nitric oxide emissions were provided by Martin Mlynczak, NASA LaRC.

HEALPix software, from <http://healpix.jpl.nasa.gov/>, developed by Krzysztof M. Gorski

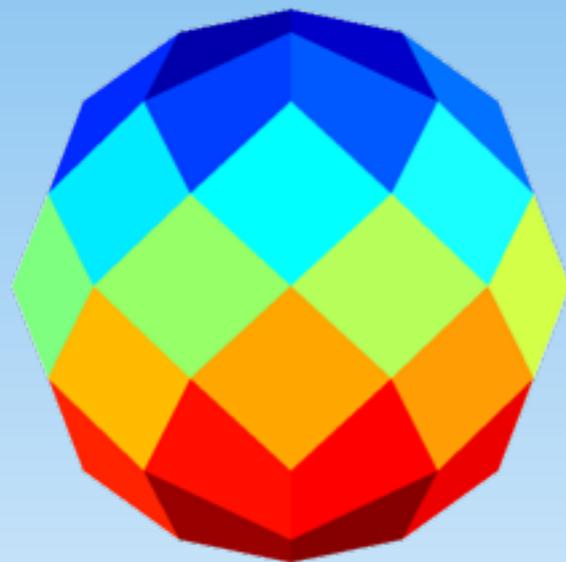
Spherical wavelet software in “Interactive Sparse Astronomical Data Analysis Packages (iSAP)” , from <http://jstarck.free.fr/isap.html>, is by J.-L. Starck

The magnetic perturbation prediction was funded by the National Space Weather Program through NSF grant ATM-0817751 to Virginia Tech. The author thanks the many personnel and institutions that maintain the numerous magnetometer arrays:

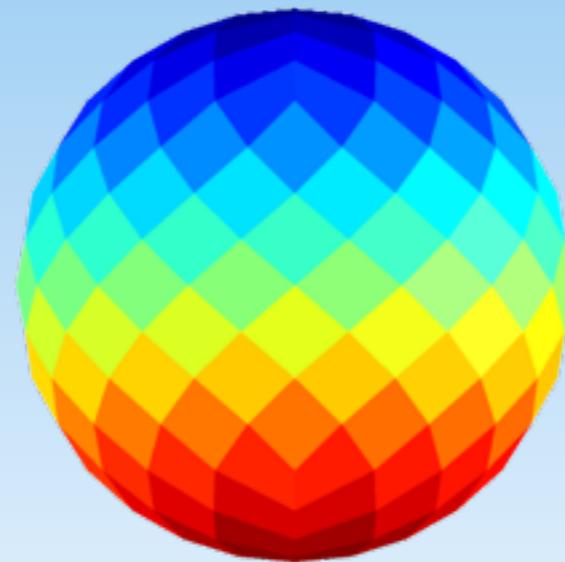
CARISMA, operated by the University of Alberta, funded by the Canadian Space Agency, the member institutes of INTERMAGNET, the Geophysical Institute of the University of Alaska, Augsburg College MACCS Project, the Danish Meteorological Institute, and the IMAGE Magnetometer Array

Using the HEALPix Grid for Mapping of Exospheric Temperatures

- HEALPix: **H**ierarchical **E**qual **A**rea iso**L**atitude **P**ixelization of a sphere.
- Each pixel covers the **same surface area** as every other pixel.
- Pixel centers located on a discrete number of rings of constant latitude.
- Developed within the astrophysics community for all-sky mapping.
- Extensive toolset for analysis, including (spherical) wavelet transforms.

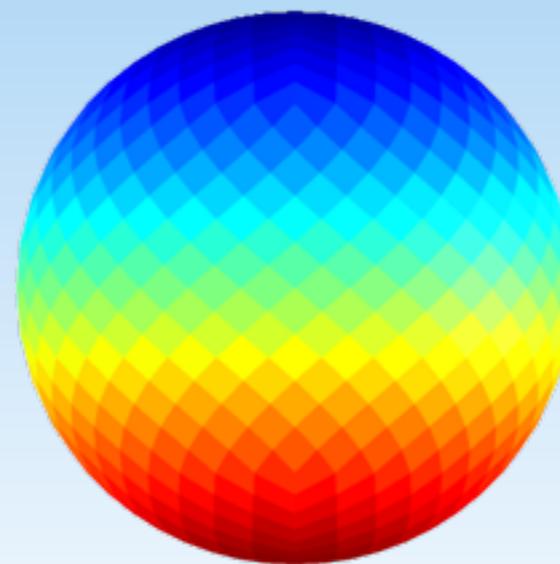


n=2, 48 pixels

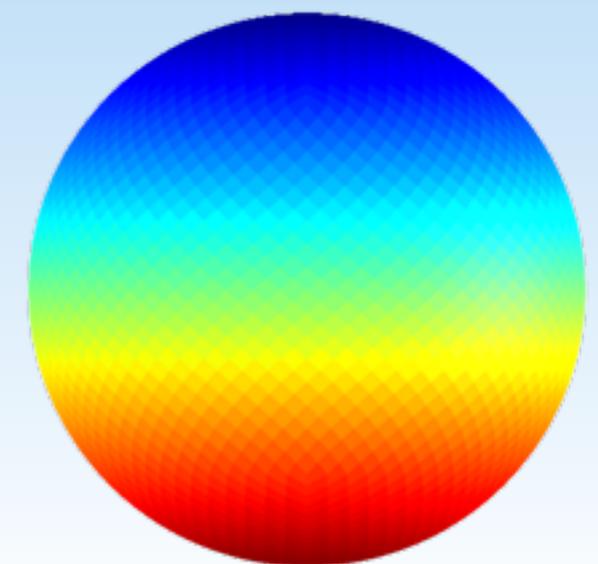


n=4, 192 pixels

n=8, 768 pixels



n=16, 3072 pixels



Progressively higher resolutions

Appendix A: Energy budget of the thermosphere

The total energy content of the thermosphere is proportional to ΔT_c . *Burke* [2008] found that a 1°K increase in ΔT_c raises the total energy in the thermosphere above 100 km altitude by $1.01 \cdot 10^{14}$ Joules.

$$\Delta T_c(t_{n+1}) = \Delta T_c(t_n) \left(1 - \frac{\Delta t}{\tau_c}\right) + \beta H_J \Delta t$$

$$\beta = 6.9 \cdot 10^{-4} \text{ (°K/GW-min)}$$

$$\tau_c = 14.6 \text{ (hours)} - 0.281 \text{ NO}$$

$$\text{NO}(t_{n+1}) = \text{NO}(t_n) \left(1 - \frac{\Delta t}{\tau_{NO}}\right) + \gamma H_J \Delta t$$

$$\gamma = 2.5 \cdot 10^{-5} \text{ (units/GW-min)}$$

$$\tau_{NO} = 28.0 \text{ (hours)}$$

With β equal to $6.9 \cdot 10^{-4}$ °K/GW/min, an output of 362 GW from the W05 model over a period of 4 min is needed to raise the T_c temperature by 1°K. A heat input of 362 GW during a 4 min interval amounts to $0.869 \cdot 10^{14}$ J, which is just slightly under the $1.01 \cdot 10^{14}$ Joule figure obtained by *Burke* [2008] for the change in energy. The particle precipitation can account for the other 14%.