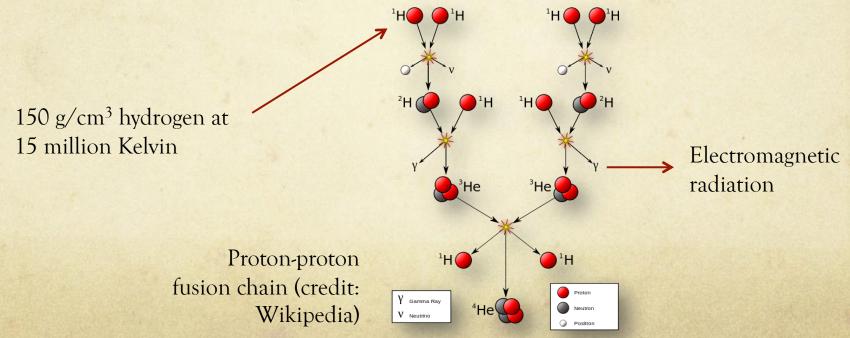




- O In this class we drill deeper into the physics of:
  - O Sun.
  - Key processes driving solar activity.
- O Some relevant iSWA product examples.

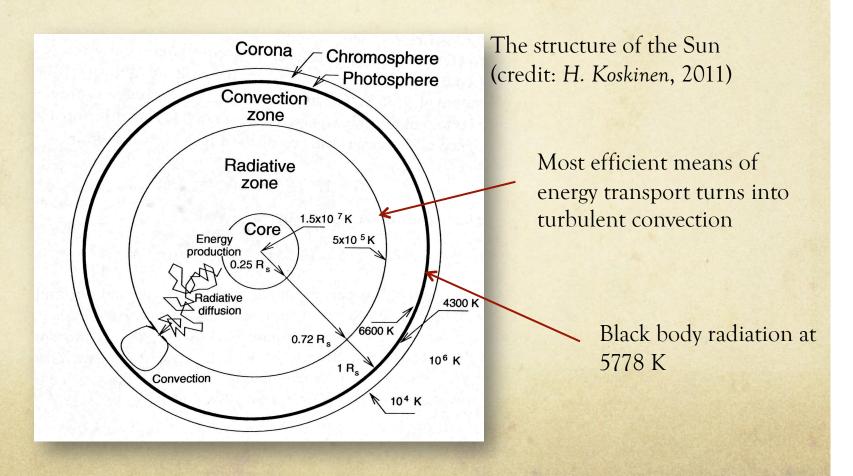


- We already learned that the Sun is the ultimate driver of space weather.
- o (primarily) Proton-proton fusion process operating in the solar core fuels everything.

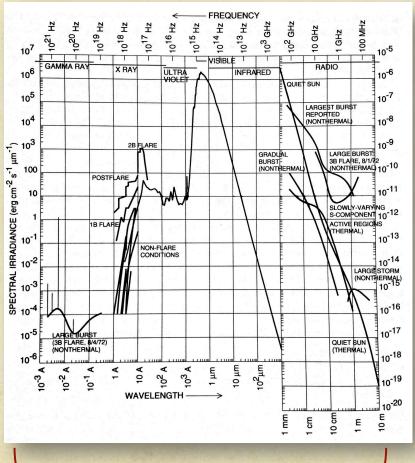




O Generated energy transported through a variety of layers.







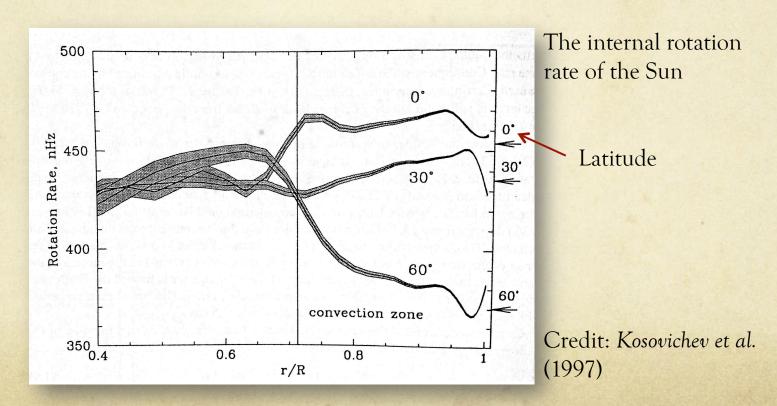
Solar spectrum from gamma rays to radio waves.

Credit: Aschwanden (2004)

Total solar irradiance or solar "constant" varying at 1 AU between about 1361 W/m<sup>2</sup> and 1362 W/m<sup>2</sup> (solar min vs max)

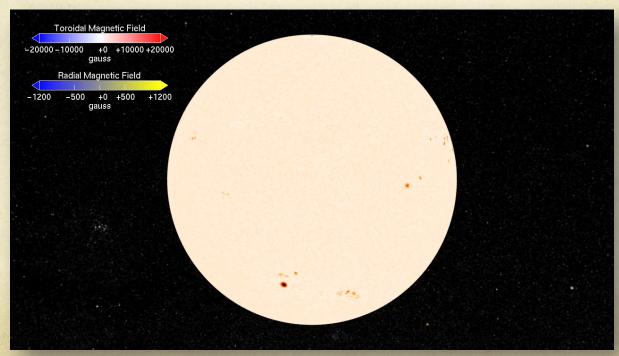


O Differential rotation one key element of the Sun. The origins of this behavior not yet fully understood.





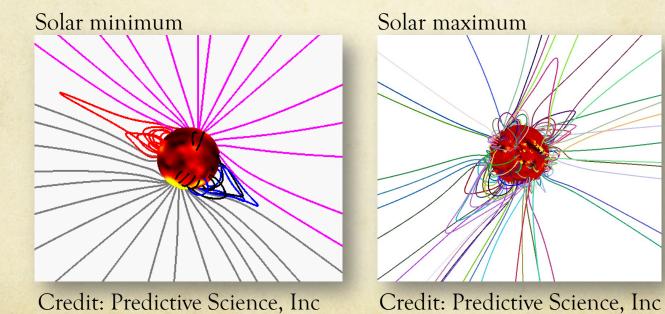
The Sun is a magnetic beast. Turbulent flows together with differential rotation in the convection zone the key ingredients for the solar dynamo.



Credit: NASA GSFC SVS



O Global structure of the solar magnetic field varies as a function of solar cycle.





## iSWA products

O Potential Field Source Surface (PFSS) models in iSWA.



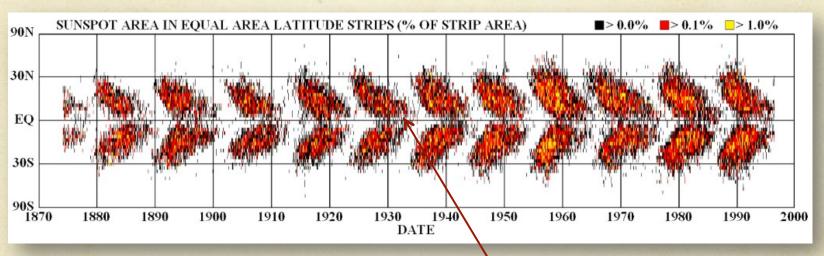
Sunspots that are at about 4100 K are reflection of "piercing" of the field generated in the convection zone through the photosphere.



Emergence of new magnetic flux through the photosphere seen in NASA SDO white light imagery.



O Cyclic solar dynamo gives rise to solar cycle and the famous butterfly diagram.



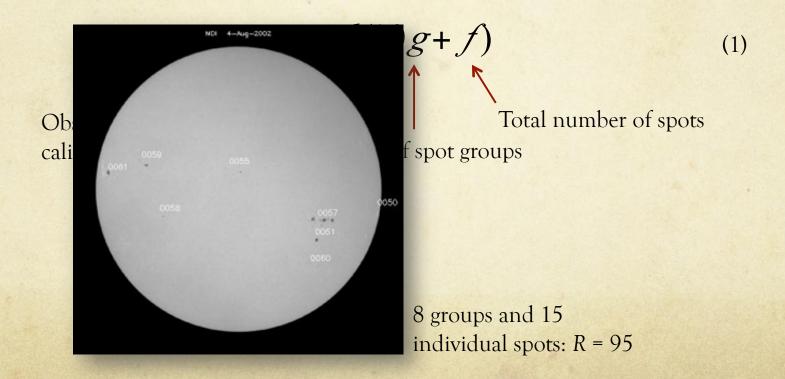
Credit: Montana State University

Solar magnetic field flips approx. over 11 year cycle so the full cycle is in fact approx. 22 years long

During the solar maximum activity close to the equator

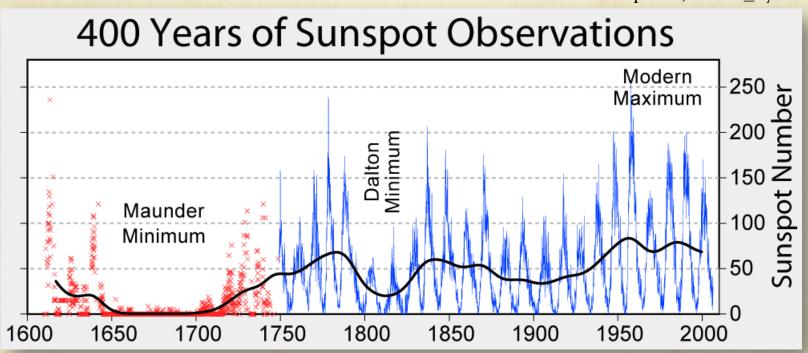


The classic (since the 18<sup>th</sup> century) means to characterize the state of the Sun is the *relative sunspot number*:





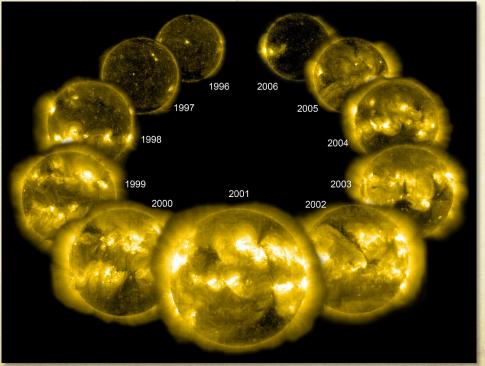
Credit: Wikipedia/Solar\_cycle



Solar cycle progression



As the global solar magnetic field structure gets more complicated also plasma configurations in the solar corona gain *complexity*.



SOHO EIT 284 Angstrom images (2 million degree plasma)

Credit: NASA/ESA



## iSWA products

O Solar EUV imagery products in iSWA.



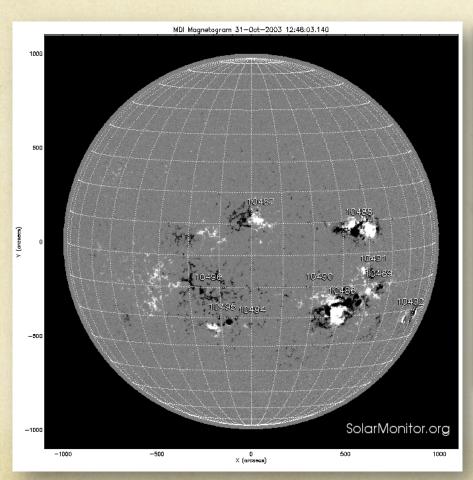
- The build up of complexity in the corona is associated with build up of *free energy* in plasma configurations.
- The build up of magnetic free energy is often characterized in terms of Mt Wilson sunspot magnetic classification of the active regions (photosphere):

Class	Description
Alpha	A unipolar sunspot group
Beta	Sunspot groups having both bipolar magnetic fields
Gamma	Sunspot groups complicated and irregular positive and negative polar structures
Beta-gamma	Sunspots with bipolar structure, but with sufficient complexity that one continuous line can be drawn between spots of opposite polarity
Delta	Umbra separated by less than 2 degrees from penumbra with opposite polarity
Beta-delta	Groups with beta magnetic classification and one or more delta spots
Beta-gamma-delta	A beta-gamma group with one or more delta spots
Gamma-delta	A sunspot group under gamma category, but with one or more delta spots

Credit: R. Benestad (2002)

These are explosive!





:Product: 20031031SRS.txt :Issued: 2003 Oct 31 0030 UTC # Prepared jointly by the U.S. Dept. of Commerce, NOAA, # Space Environment Center and the U.S. Air Force. Joint USAF/NOAA Solar Region Summary SRS Number 304 Issued at 0030Z on 31 Oct 2003 Report compiled from data received at SWO on 30 Oct I. Regions with Sunspots. Locations Valid at 30/2400Z Nmbr Location Lo Area Z LL NN Mag Type 0484 N01W95 356 0210 Dao 10 06 Beta-Gamma 0486 S18W23 284 2600 Fkc 18 80 Beta-Gamma-Delta 0487 N12E06 255 0280 Dko 07 23 Beta 0488 N08W28 289 1750 Fkc 17 34 Beta-Gamma-Delta 0489 S12W36 297 0130 Dao 06 09 Beta 0490 S12W14 275 0010 Hrx 01 03 Alpha 0491 S06W32 293 0120 Dso 07 10 Beta 0492 S23W62 323 0340 Eko 11 17 Beta 0494 S23E08 253 0010 Axx 00 01 Alpha 0495 S22E20 241 0240 Dso 08 10 Beta

Credit: NOAA SWPC

Credit: SolarMonitor.org



# iSWA products

O SolarScape.



### The Sun and drivers

O And from here we get to flares and CMEs...